

Two years of GW170817

Eleonora Troja

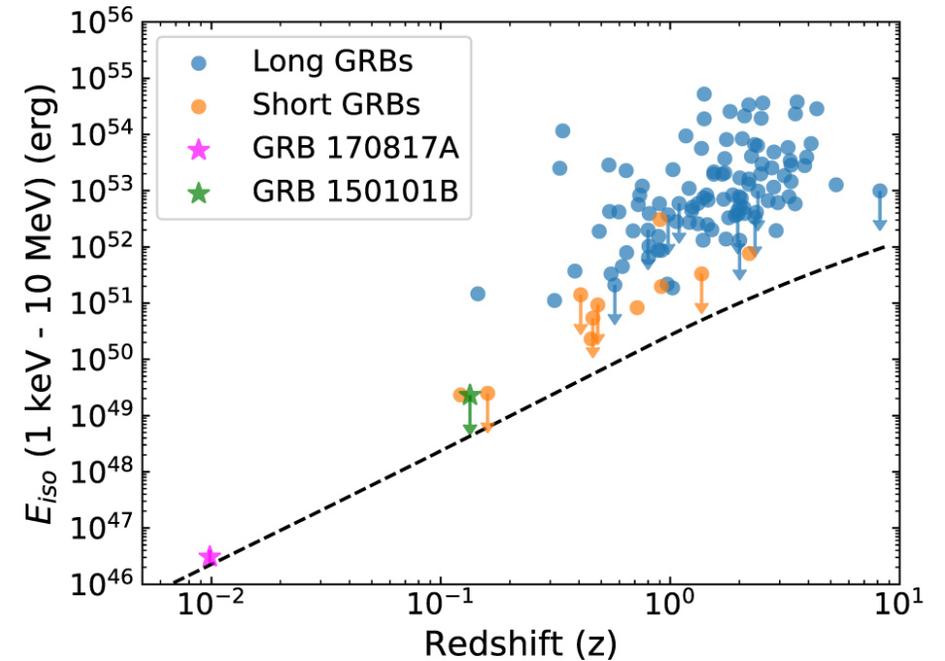
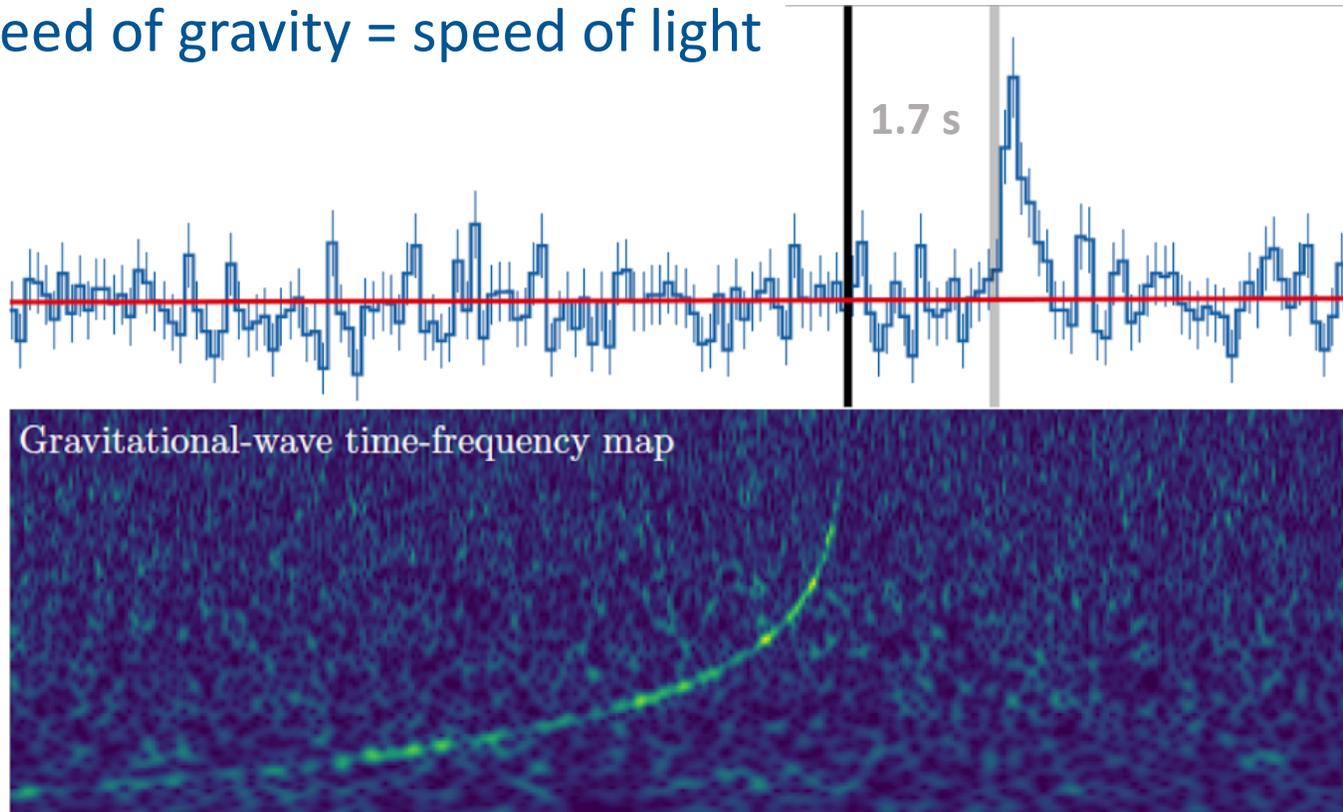
University of Maryland, College Park

NASA Goddard Space Flight Center



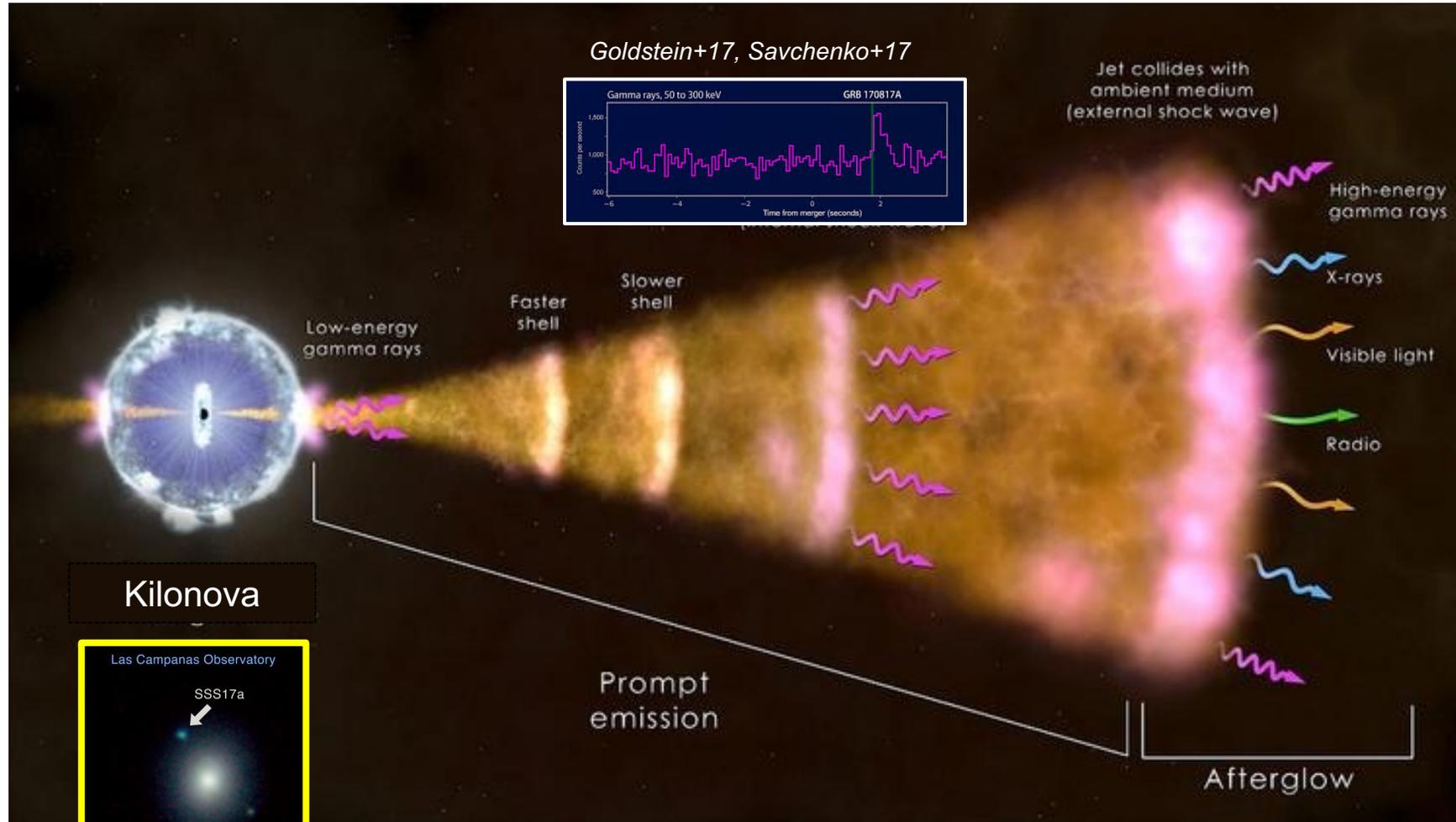
A short GRB following GW170817

speed of gravity = speed of light

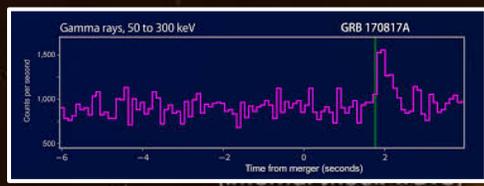


Abbott *et al.* (LIGO Scientific Collaboration and Virgo Collaboration)
2017, Phys. Rev. Lett. **119**, 161101

The aftermath of a NS merger



Goldstein+17, Savchenko+17



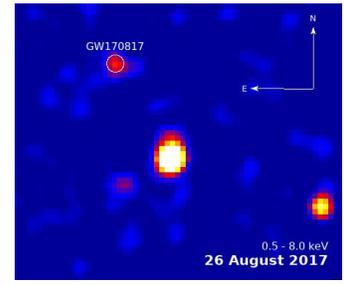
Kilonova



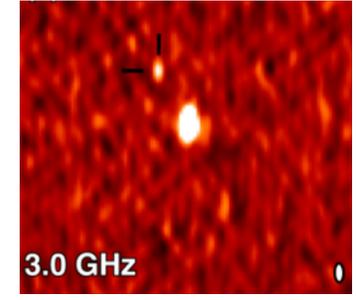
Coulter+17

Arcavi+17, Lipunov+17, Soares-Santos+17, Tanvir+17, Valenti+17

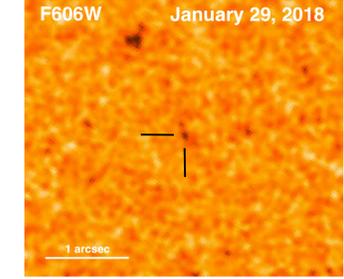
Troja+17



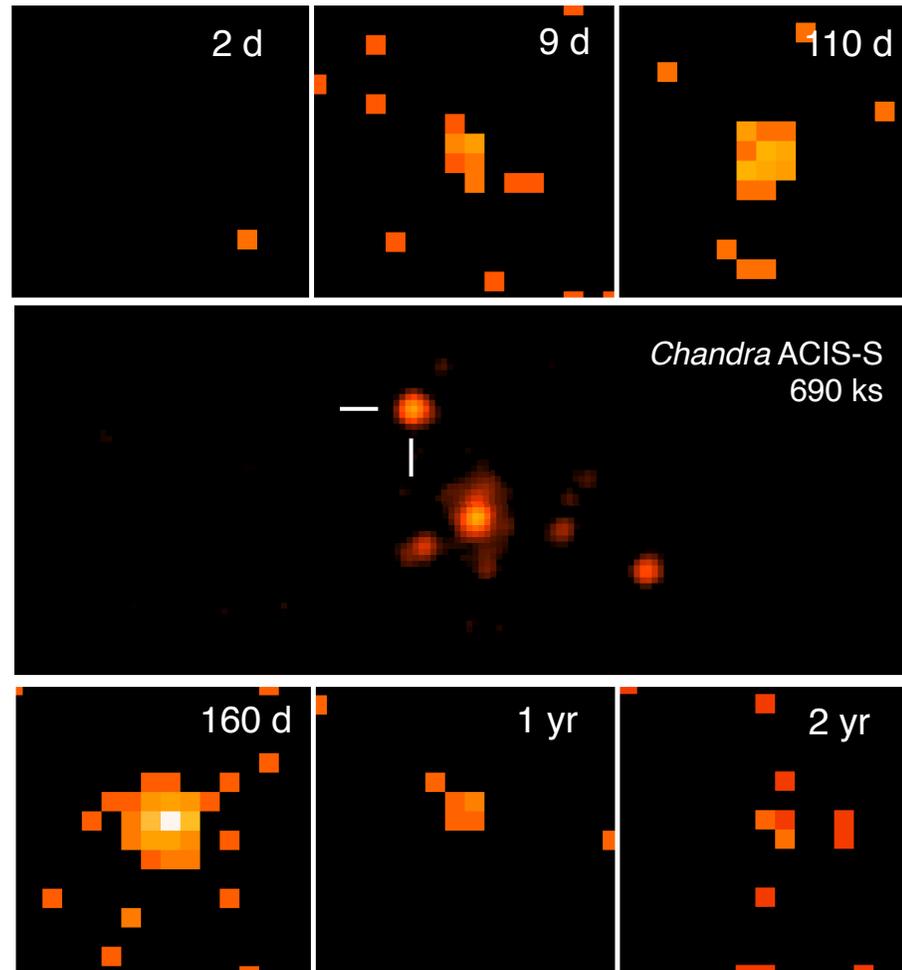
Hallinan, Corsi+17



Lyman+18



The X-ray afterglow

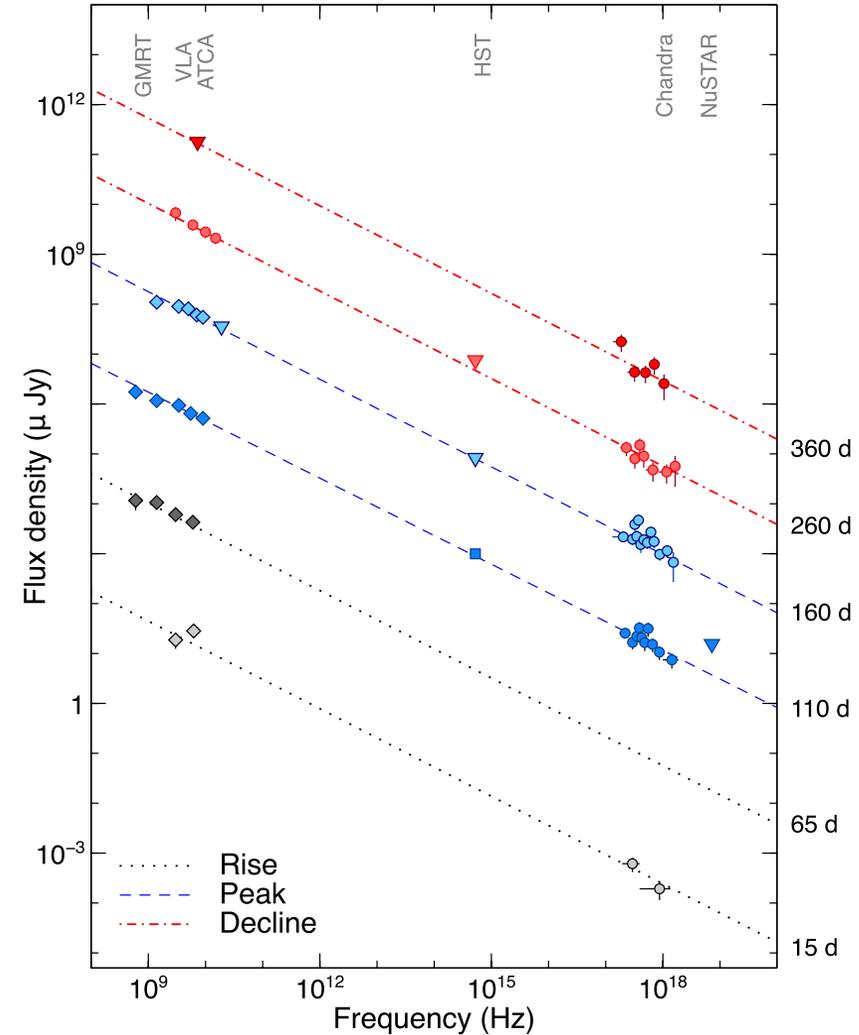
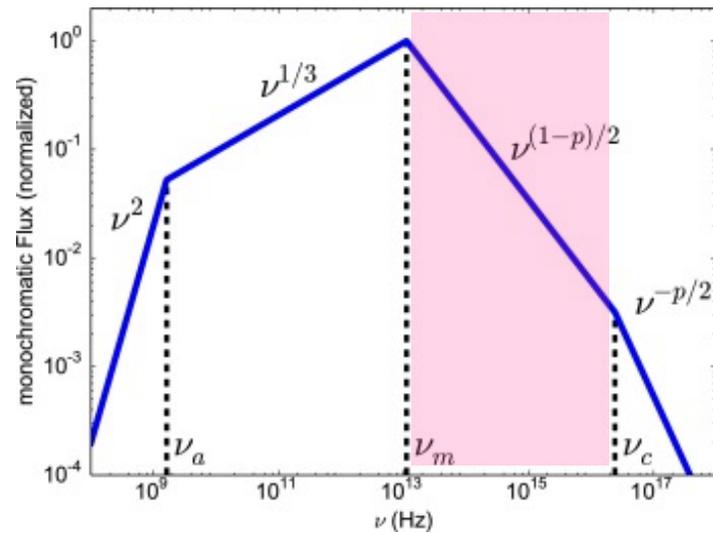


Troja+17
Haggard+17
Margutti+17
Troja+18
Piro+19
Nynka+18
Pooley+18
Alexander+18
Troja+19

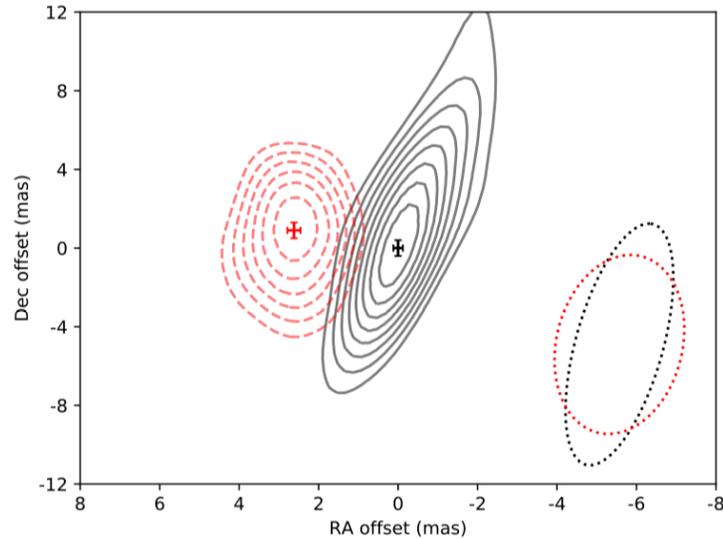
No Spectral Evolution

Simple power-law spectrum
over 10 decades in energy

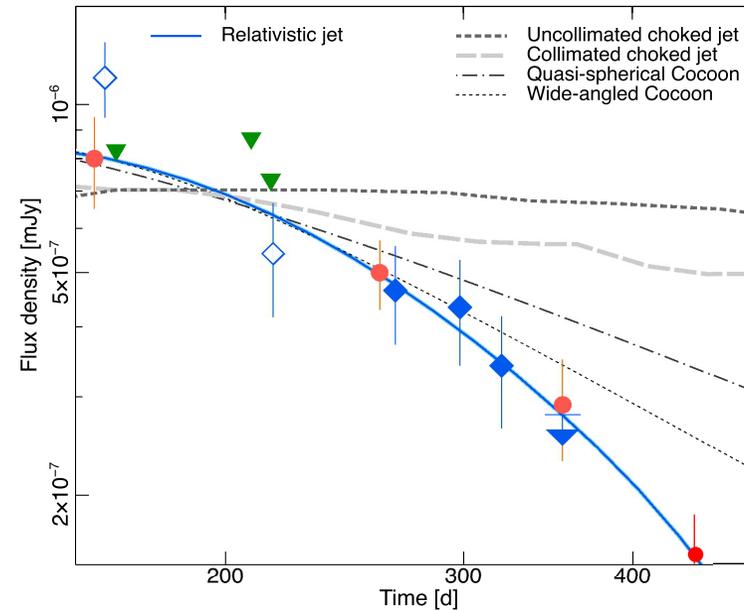
Consistent with synchrotron emission
regime $\nu_m < \nu_r < \nu_x < \nu_c$



Evidence for a successful relativistic jet



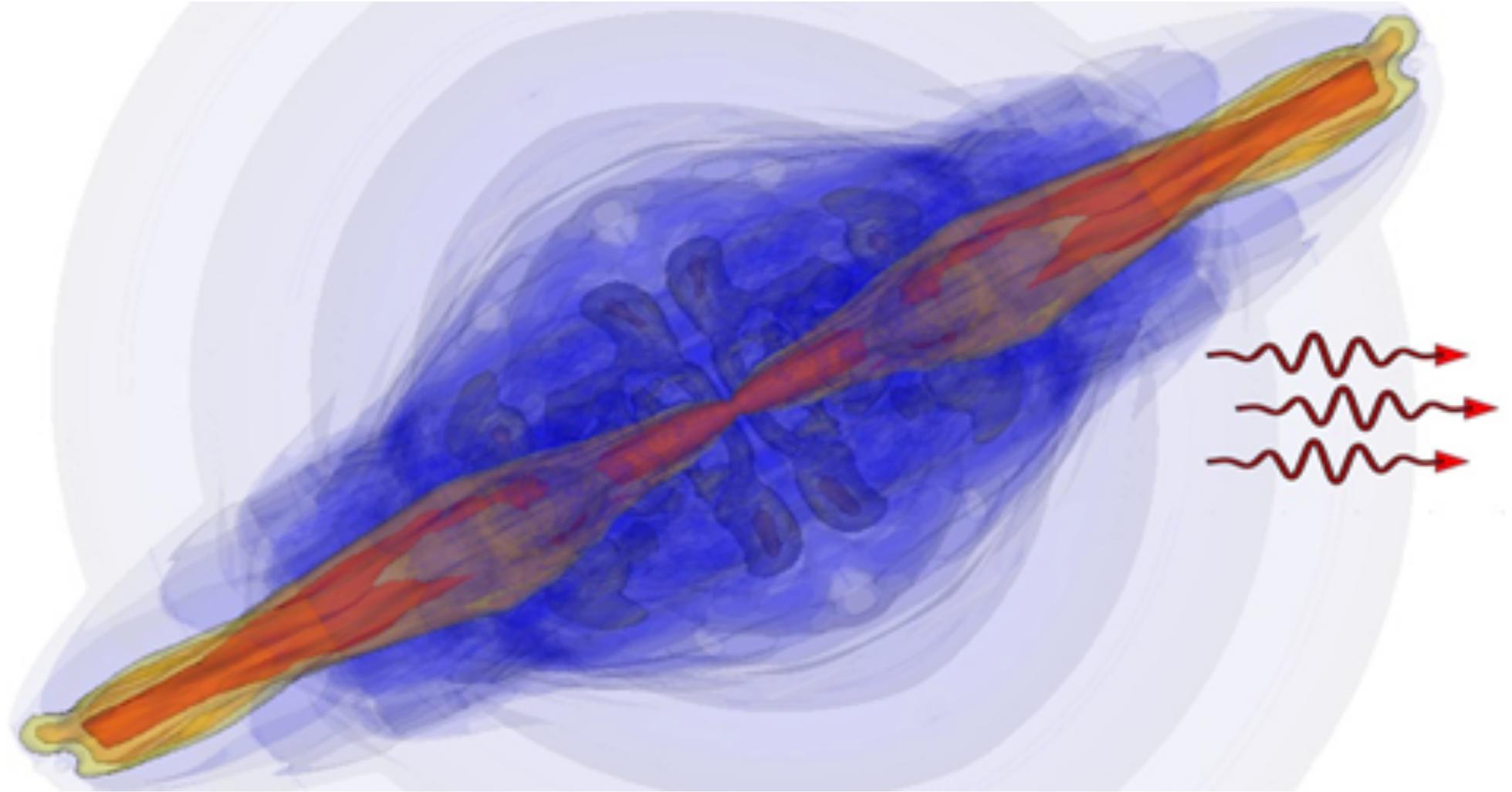
Mooley et al., 2018
Ghirlanda et al., 2019



Troja, van Eerten et al., 2019

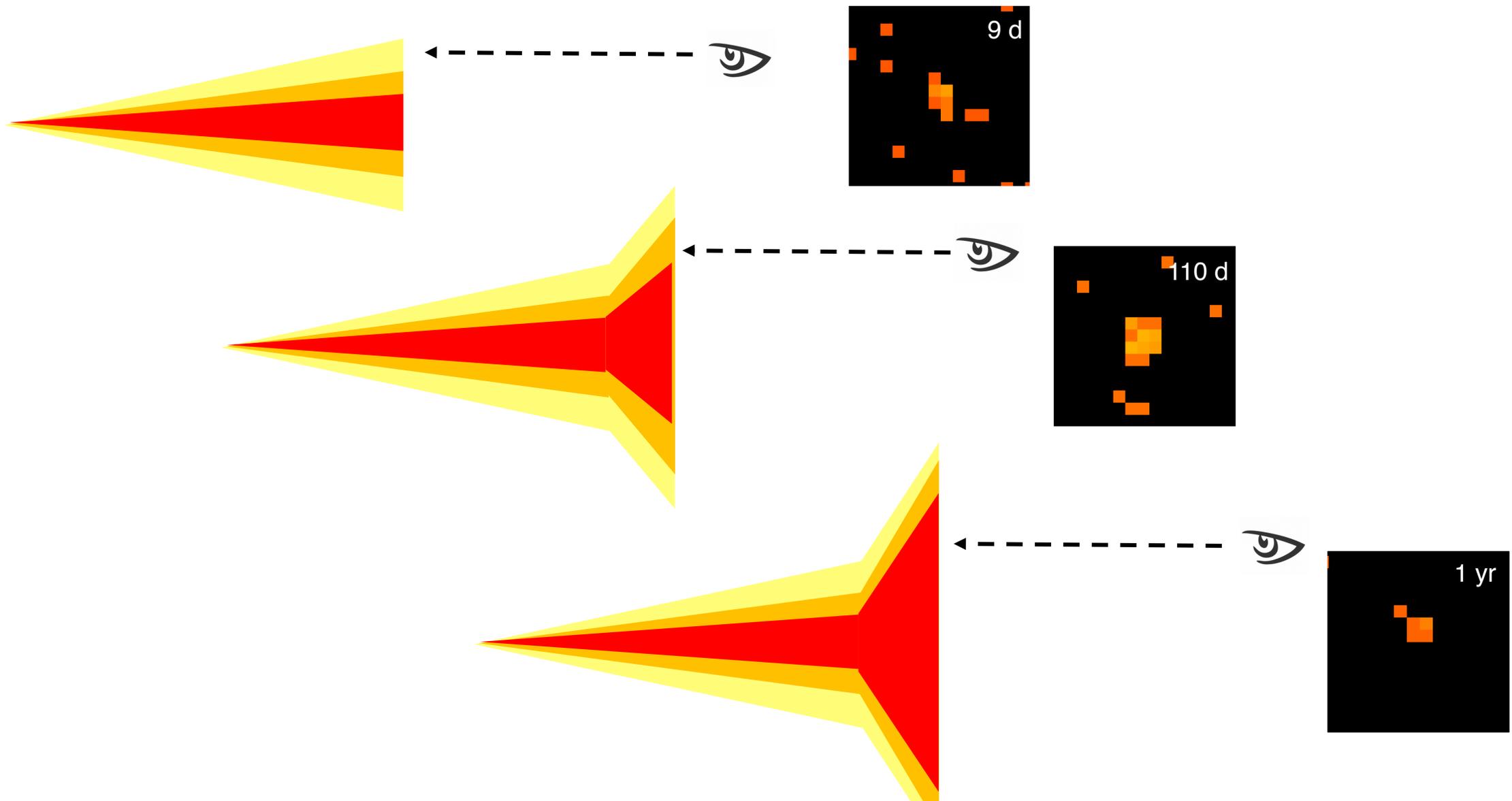
- **High-resolution radio imaging:** compact unresolved radio source
superluminal motion
- **Temporal monitoring:** rapid afterglow decline

A structured jet seen off-axis

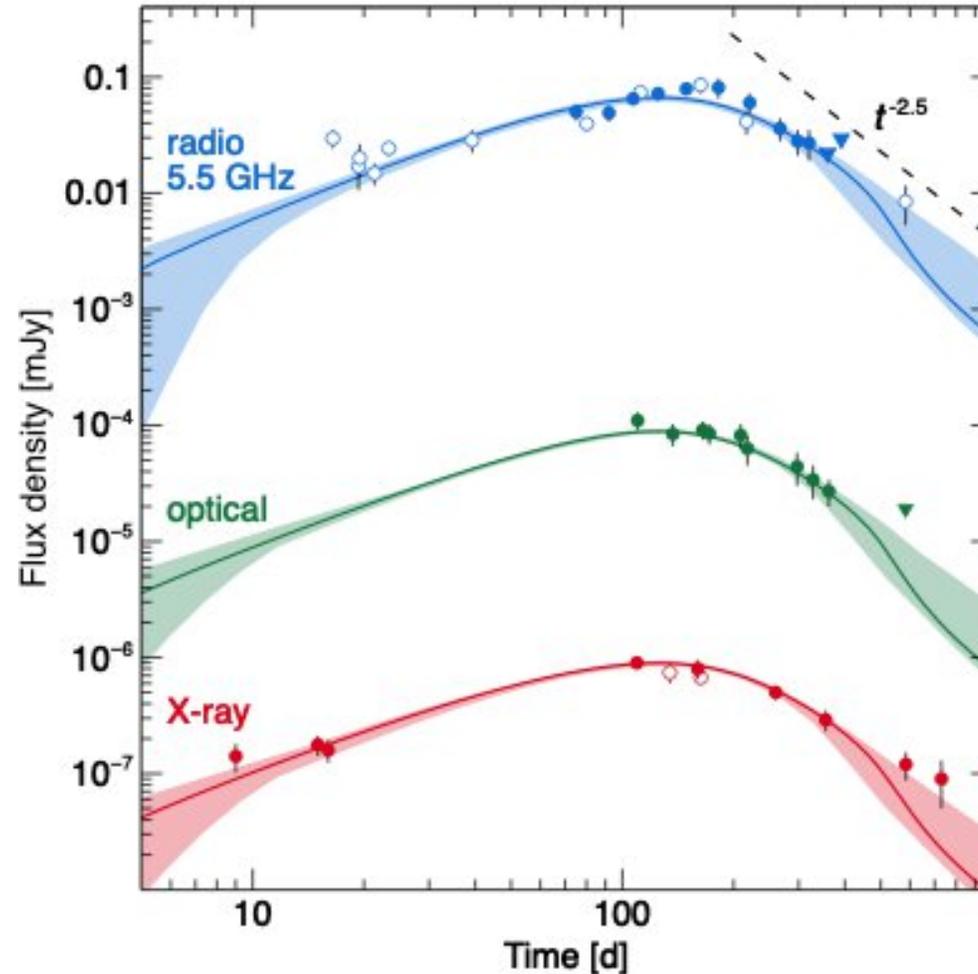


Zhang & Meszaros 02
Rossi+02
Aloy+05
Janka+05
Lazzati+18
Xu+18
Kathiramaraju+18
Ryan+19

The afterglow in the off-axis structured jet model



Broadband afterglow modeling

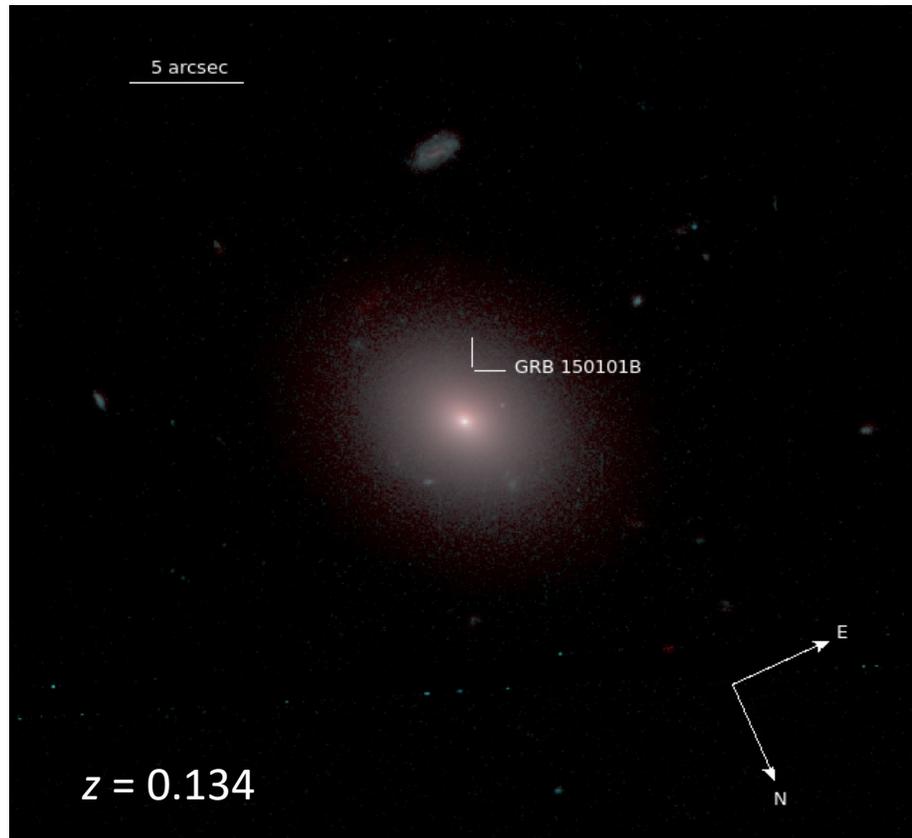


$\theta_{\text{jet}} \sim 5 \text{ deg}$
 $\theta_{\text{view}} \sim 25 \text{ deg}$
 $n \sim 10^{-2} - 10^{-4} \text{ cm}^{-3}$
 $E \sim 10^{50} \text{ erg}$

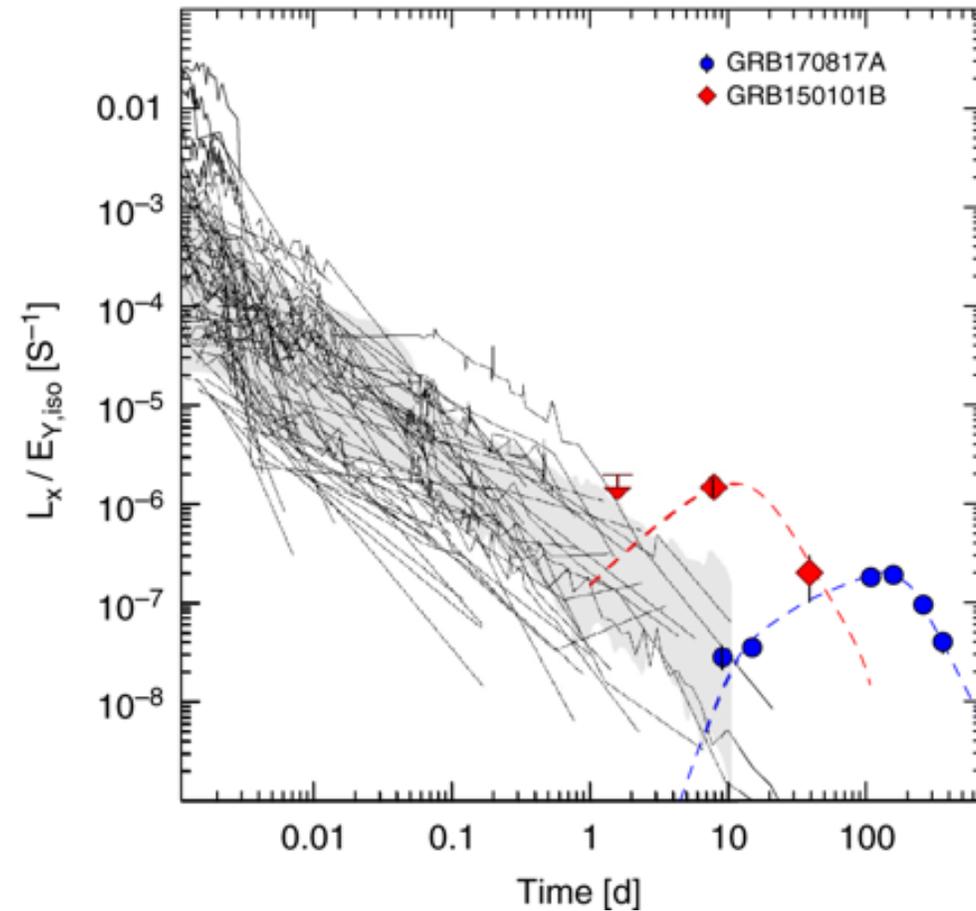
Typical of short GRB afterglows

EM viewing angle consistent
with the binary inclination from
GW data

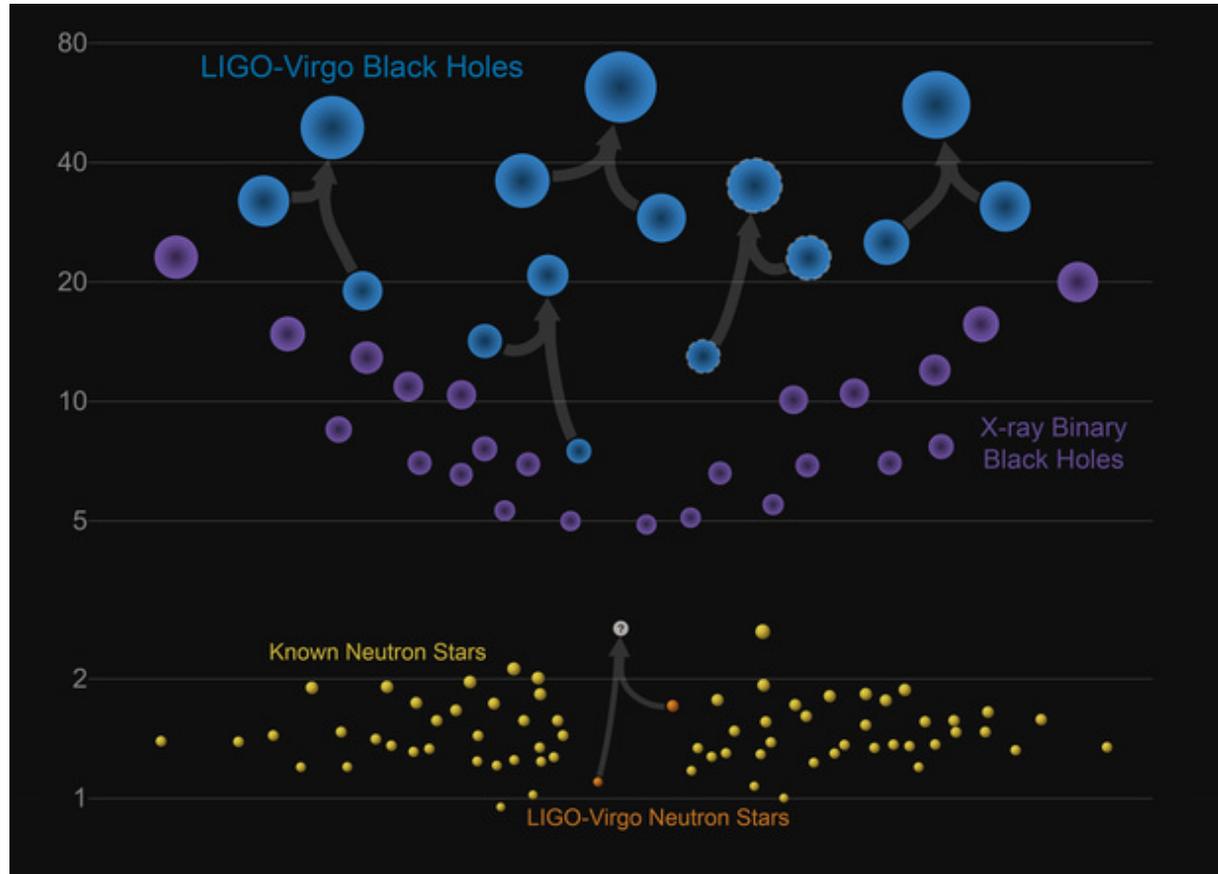
Analogues in the GRB sample: GRB 150101B



Troja, Ryan et al. 18



Constraints on the remnant: NS or BH?



X-ray emission is very sensitive to the GRB central engine: sporadic emission of energy (flares) or continuous spin-down energy injection (plateaus).

$$B < 10^{12} \text{ G}$$

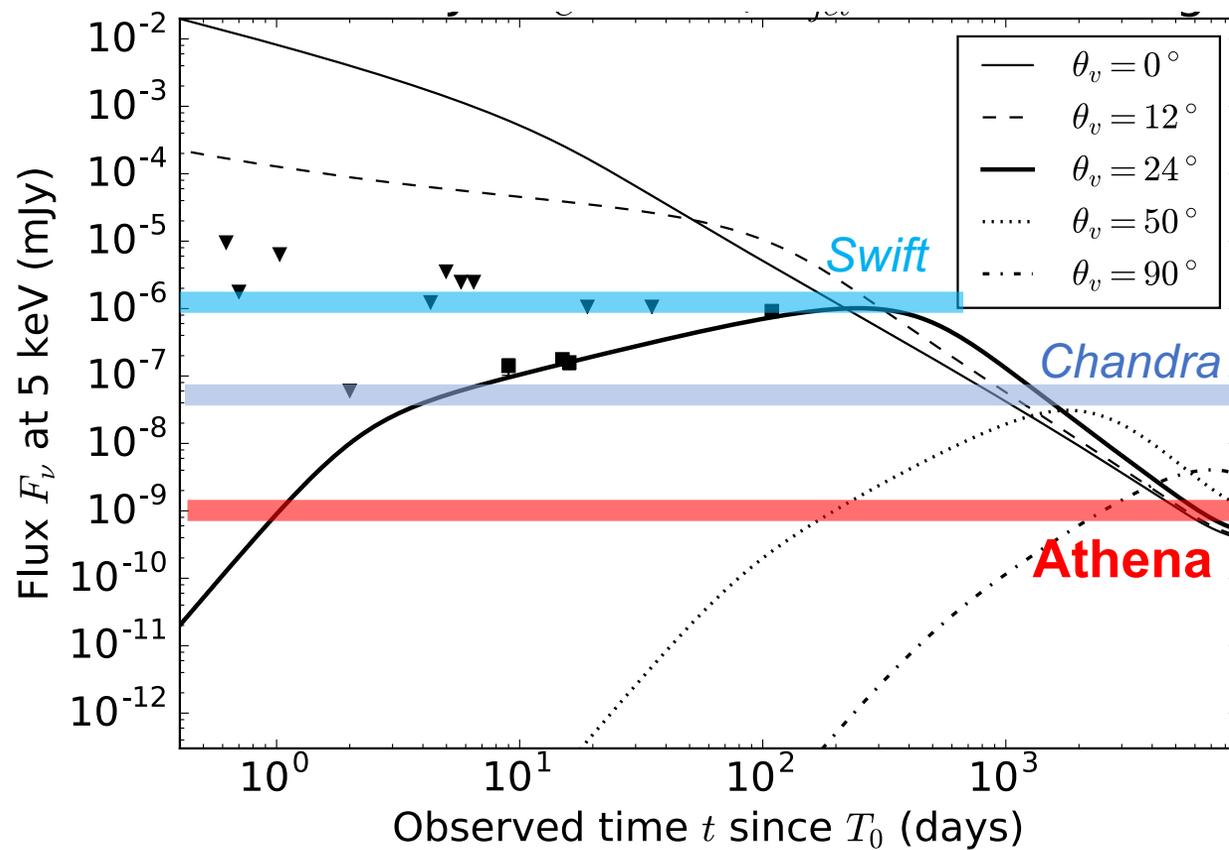
BH

Pooley+18

or stable low-B NS

Piro+19

Future perspectives



Summary

- NS mergers can launch collimated relativistic outflows (jets) powering GRBs
- Viewing angles play an important role:
Similar explosions might look very different
- GW170817: an extraordinary ordinary short GRB
A few similar events might have been seen before
- ATHENA will probe a wider range of explosions