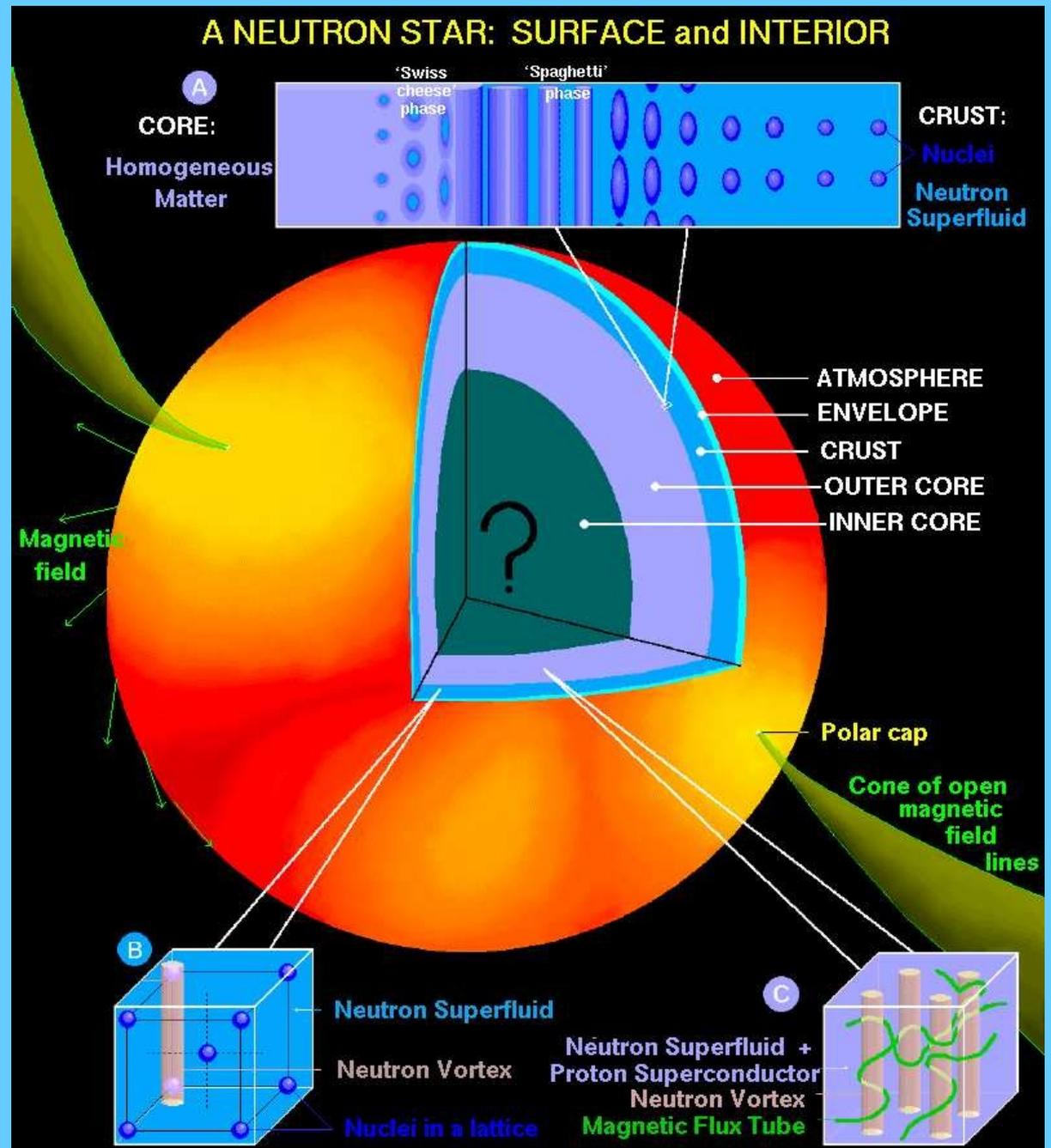


Neutron matter Equation of State at very low density

StPetersburg june 2008

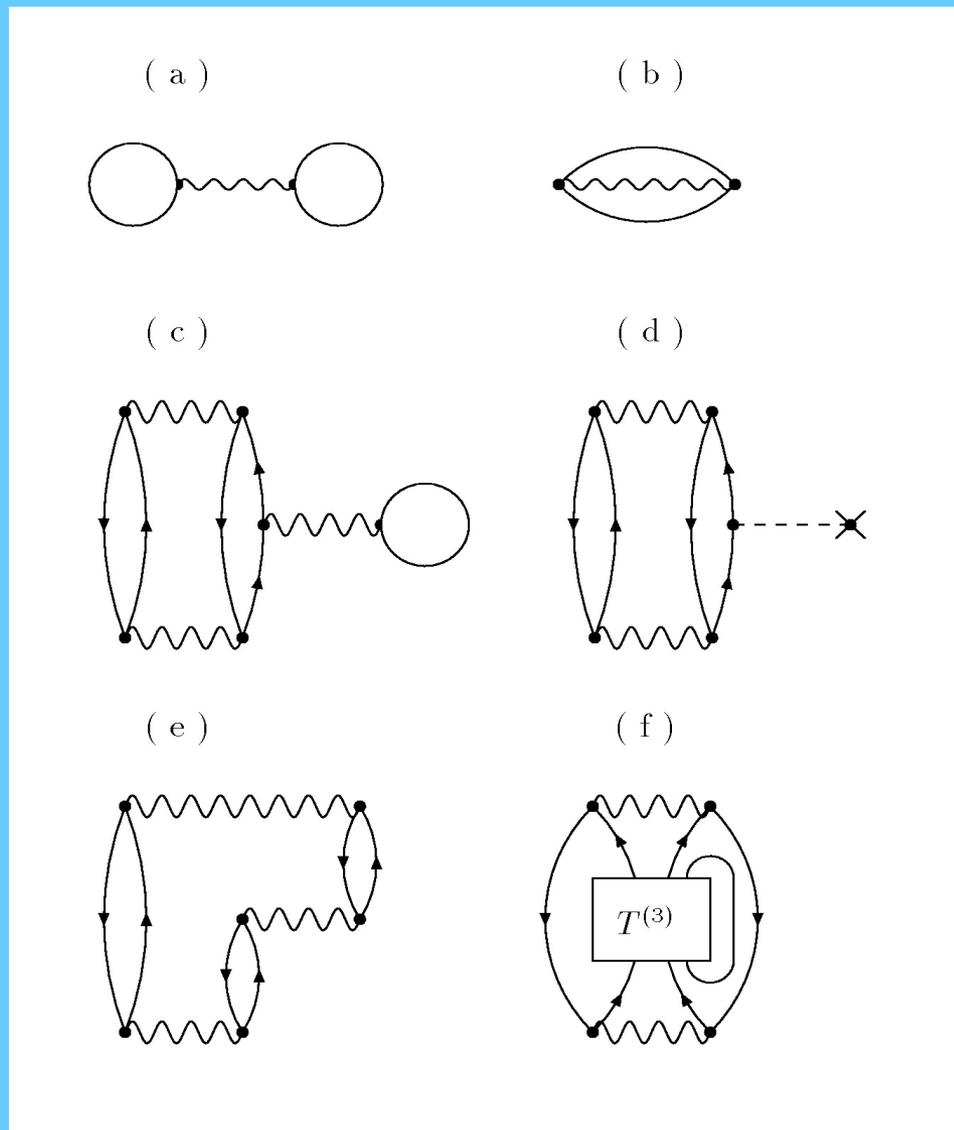
A section (schematic)
of a neutron star



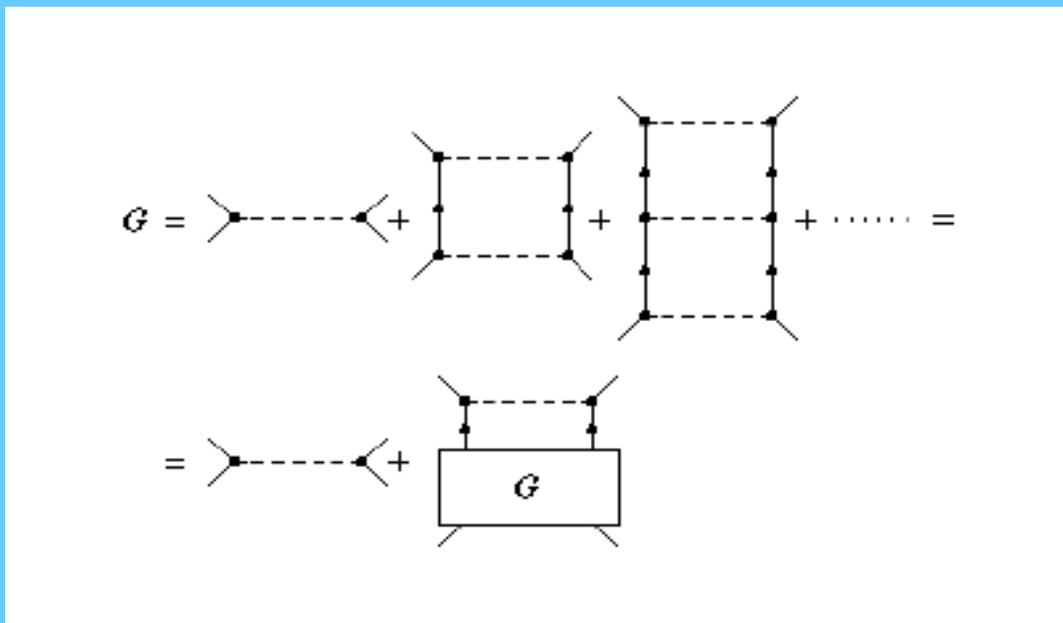
Neutron matter studies

Motivations

1. Low density neutron matter in the drip region
3. “Close” to the unitary limit ($-a \longrightarrow \infty$)
3. Set the uncertainty in the many-body treatment

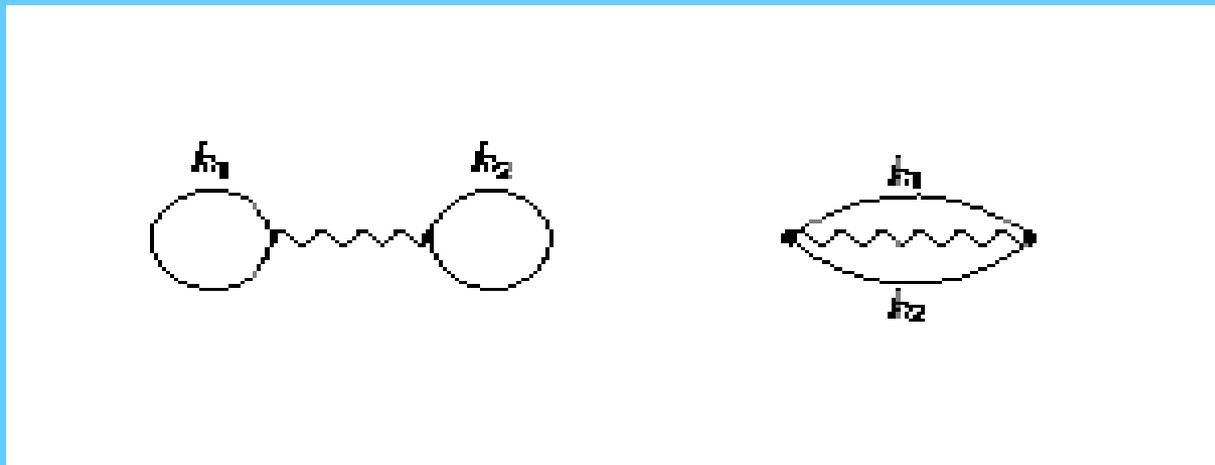


Two and three hole-line diagrams in terms of the Brueckner G-matrix



Ladder diagrams for the scattering G-matrix

$$G = V + V \frac{Q}{e} G$$



$$T^{(3)} = \begin{array}{c}
 \begin{array}{c}
 | \quad | \quad | \\
 \bullet \text{---} \text{wavy} \text{---} \bullet \\
 k_1 \quad k_2 \quad k_3
 \end{array}
 + \begin{array}{c}
 | \quad | \quad | \\
 \bullet \text{---} \text{wavy} \text{---} \bullet \text{---} \text{wavy} \text{---} \bullet \\
 \uparrow \quad \uparrow
 \end{array}
 + \begin{array}{c}
 | \quad | \quad | \\
 \bullet \text{---} \text{wavy} \text{---} \bullet \text{---} \text{wavy} \text{---} \bullet \\
 \uparrow \quad \uparrow
 \end{array}
 + \dots \\
 \\
 \begin{array}{c}
 | \quad | \quad | \\
 \bullet \text{---} \text{wavy} \text{---} \bullet \text{---} \text{wavy} \text{---} \bullet \\
 \uparrow \quad \uparrow \quad \uparrow
 \end{array}
 + \begin{array}{c}
 | \quad | \quad | \\
 \bullet \text{---} \text{wavy} \text{---} \bullet \text{---} \text{wavy} \text{---} \bullet \\
 \uparrow \quad \uparrow \quad \uparrow
 \end{array}
 + \begin{array}{c}
 | \quad | \quad | \\
 \bullet \text{---} \text{wavy} \text{---} \bullet \text{---} \text{wavy} \text{---} \bullet \\
 \uparrow \quad \uparrow \quad \uparrow
 \end{array}
 + \dots
 \end{array}$$

The ladder series for the three-particle scattering matrix

$$T_3 = G + GX \frac{Q_3}{e} T_3$$

$$E_{3h} =$$

$$\frac{1}{2} \sum_{k_1 k_2 k_3} \sum_{[k' k'']} \langle k_1 k_2 | G | k_1' k_2' \rangle_A$$

$$\frac{1}{e} \langle k_1' k_2' k_3 | XT_3 X | k_1'' k_2'' k_3 \rangle \frac{1}{e}$$

$$\langle k_1'' k_2'' | G | k_1 k_2 \rangle_A$$

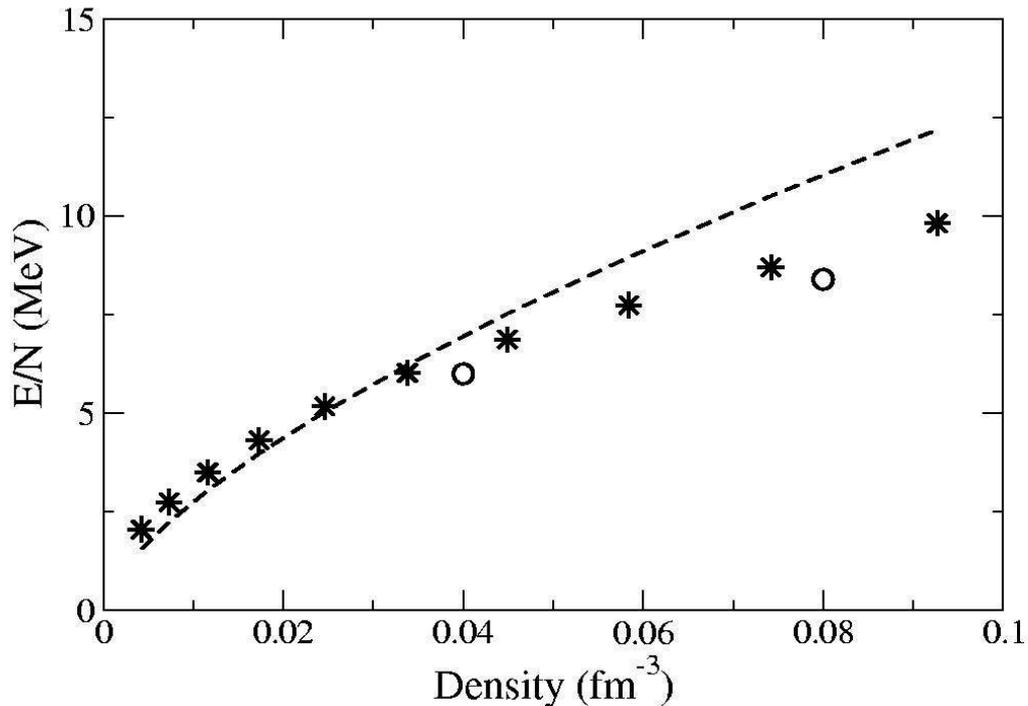
$$k_1, k_2, k_3 \leq k_F$$

$$k_1', k_2', k_1'', k_2'' \geq k_F$$

Three hole-line contribution

“Low” density

PHYSICAL REVIEW C 68, 025802 (2003)

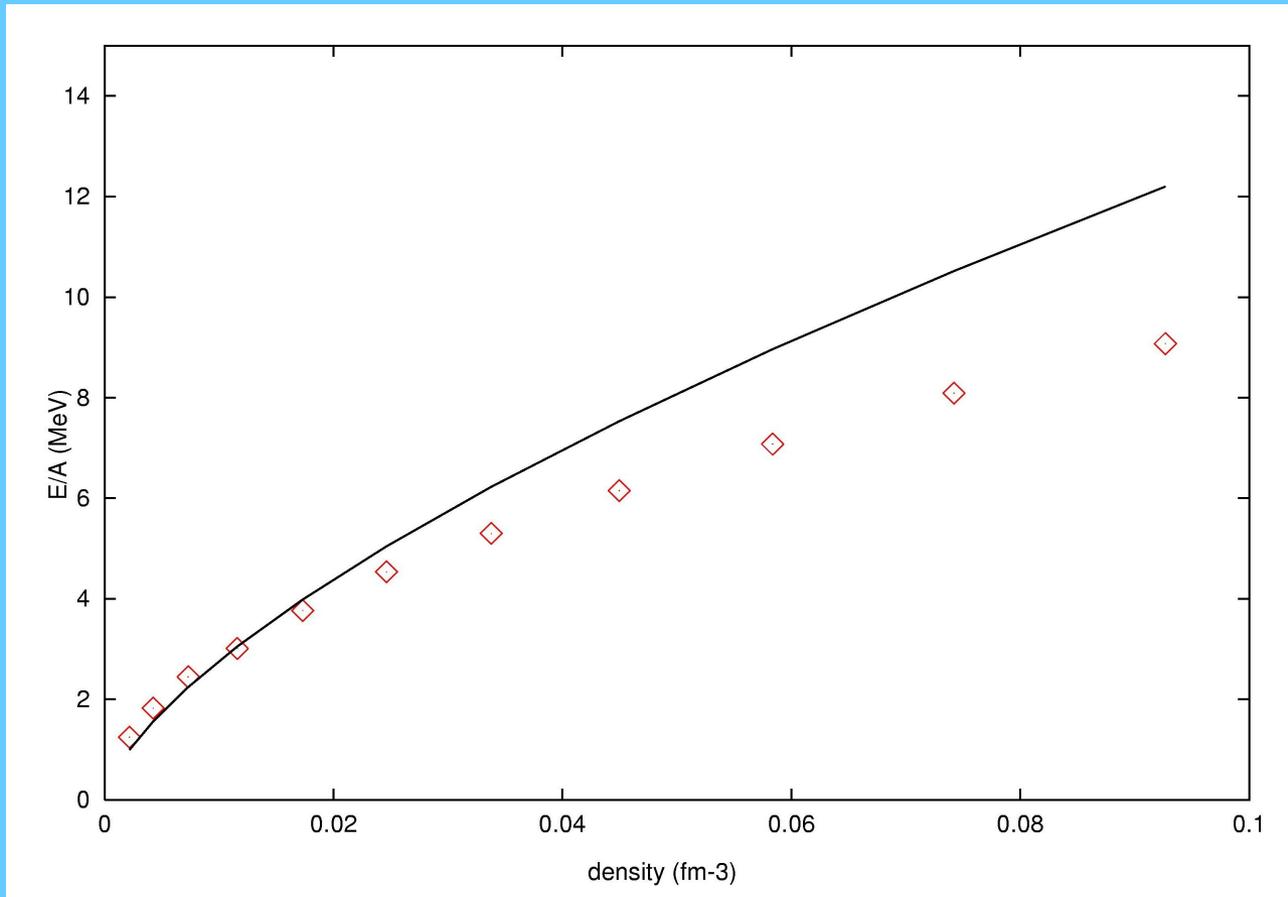


Puzzling “quadratic” behaviour of interaction energy

Dotted line : $\frac{1}{2} E(\text{free gas})$

Stars : Friedman & Pandharipande, Nucl. Phys. A361,501(1981)

Urbana potential



Similar behaviour in BBG calculations (v18 potential)

These results are suggestive of the “unitary limit” behaviour of neutron matter :

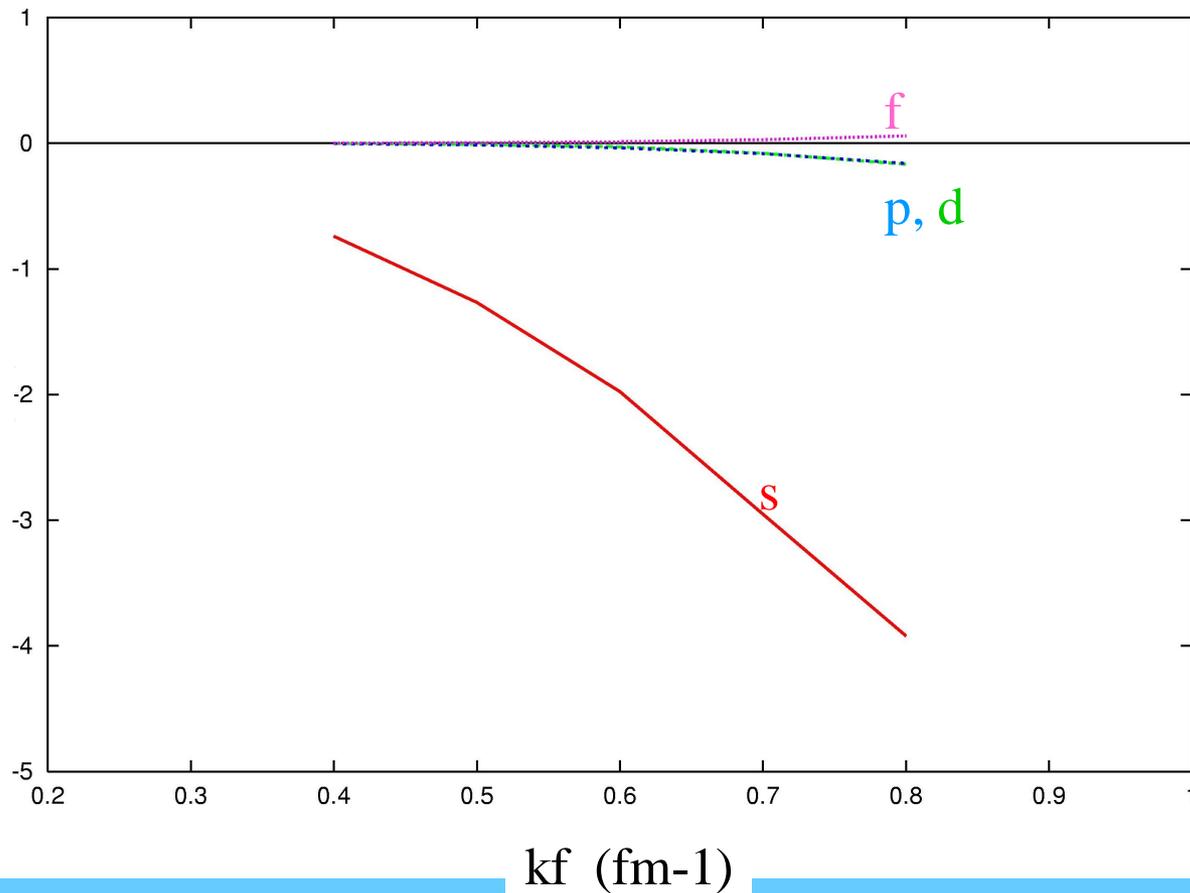
$$-a \gg d \gg r$$

where a is the scattering length, r is the effective range and d is the average distance between particles. For $a \gg d$ — the only scale is given by $d \approx 1/k_F$ and the corresponding energy scale can be only the kinetic energy

$$E = \xi E_F$$

with ξ a density independent factor

B/A
(MeV)



“Low”
density
region

1. The s-wave dominates
2. The three hole-lines are small (< 0.2 MeV)
3. Three-body forces are negligible (< 0.01 MeV)
4. Effect of self-consistent U is small (see later)

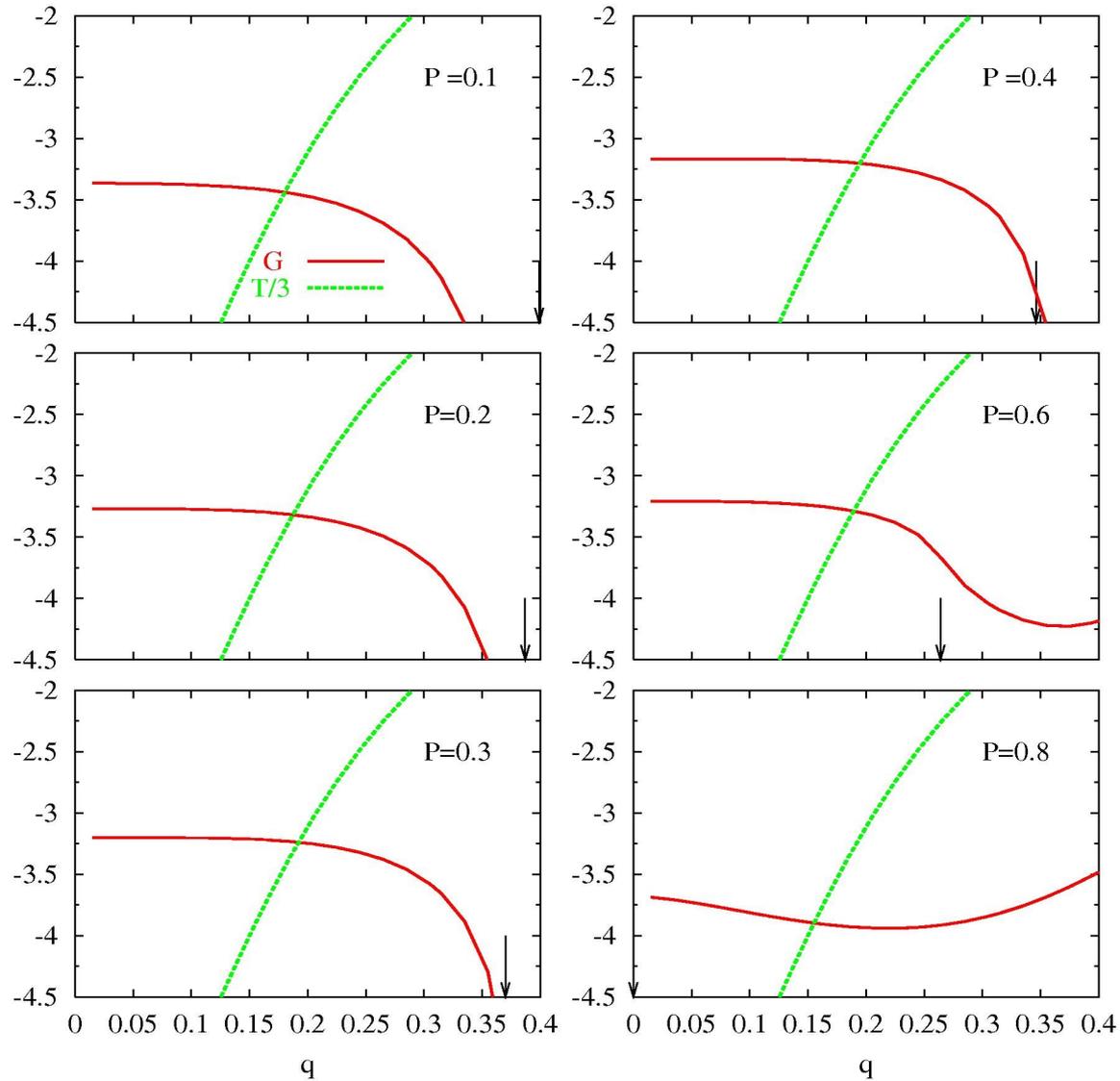
M.B. & C. Maieron

Three hole-line contribution

k_F	0.4	0.5	0.6	0.7	0.8	(fm ⁻¹)
D_3	0.023	0.091	0.107	0.153	0.148	(MeV)
B	-0.630	-0.416	-0.526	-0.611	-0.592	
BU	0.485	0.389	0.515	0.648	0.651	
R	0.156	0.122	0.123	0.121	0.095	
H	0.012	-0.004	-0.005	-0.005	-0.006	

M.B. & C. Maieron, PRC 77, 015801 (2008)

$k_F=0.4$, q dependence at fixed P



Using effective theory

Pethick and Schwenk PRL **95** 160401 (2005)

(Pauli operator only)

Introduce a local and energy dependent interaction

$$\mathbf{V}(\mathbf{E}) = \delta(\mathbf{r} - \mathbf{r}') \cdot \mathbf{g}^2 / (\Delta - \mathbf{E})$$

together with a cutoff Λ and adjust \mathbf{g} and Δ to reproduce the scattering length \mathbf{a} and effective range \mathbf{r}_e to lowest order in Λ . Taking only cutoff independent terms in the ladder sum one gets

$$\mathbf{T}_{\text{med}} = 4\pi / \mathbf{D}$$

$$\mathbf{D} = \frac{1}{\mathbf{a}} - \frac{1}{2} \mathbf{r}_e \mathbf{k}^2 - \frac{\mathbf{k}_f + 0.5\mathbf{P}}{\pi} + \frac{\mathbf{k}}{\pi} \log\left(\frac{\mathbf{k}_F + 0.5\mathbf{P} + \mathbf{k}}{\mathbf{k}_F + 0.5\mathbf{P} - \mathbf{k}}\right) + \frac{\mathbf{k}^2 + 0.25\mathbf{P}^2 - \mathbf{k}_F^2}{\pi\mathbf{P}} \log\left(\frac{(\mathbf{k}_F + 0.5\mathbf{P})^2 - \mathbf{k}^2}{(\mathbf{k}_F - 0.5\mathbf{P})^2 - \mathbf{k}^2}\right)$$

The “in medium” scattering length and effective range turn out to be

$$\mathbf{a}' = \frac{\mathbf{a}}{1 - 2\mathbf{a}\mathbf{k}_F/\pi} \quad \mathbf{r}'_e = \mathbf{r}_e - \frac{4}{\pi\mathbf{k}_F}$$

For $\mathbf{a}\mathbf{k}_F \gg 1$, $\mathbf{a}' \sim -\frac{1}{\mathbf{k}_F}$. However this is valid only if the term in the effective range \mathbf{r}_e is small, and this implies

$$-\frac{1}{\mathbf{a}} \ll \mathbf{k}_F \ll \frac{2}{\mathbf{r}_e}$$

For the **n-n** system $1/\mathbf{a} \sim -0.2$ and $\frac{2}{\mathbf{r}_e} \sim 0.7$, and the conditions cannot be satisfied.

A simple exercise in nuclear matter

Calculate the neutron matter EOS at low density

Take a separable representation for the 1S0 channel

$$V(k, k') = \lambda \varphi(k) \varphi(k') \quad \text{with e.g.} \quad \varphi(k) = [k^2 + \beta^2]^{-1}$$

for which the free scattering matrix reads

$$T(k, k') = \lambda \varphi(k) \varphi(k') / [1 - \lambda \langle \varphi | G_0(E) | \varphi \rangle]$$

where $G_0(E)$ is the free two-body Green's function. Then fix the parameters λ , β in order to reproduce the scattering length and effective range for this channel (low energy data)

The in-medium G-matrix reads

$$G(k, k') = \lambda \varphi(k) \varphi(k') / [1 - \lambda \langle \varphi | Q G_0(E) | \varphi \rangle]$$

where Q is the Pauli operator. Compare G-matrix and T-matrix.

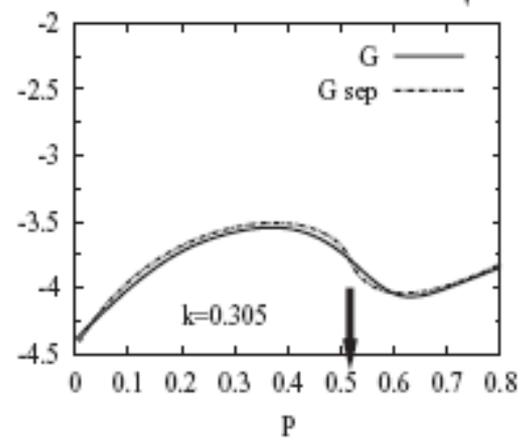
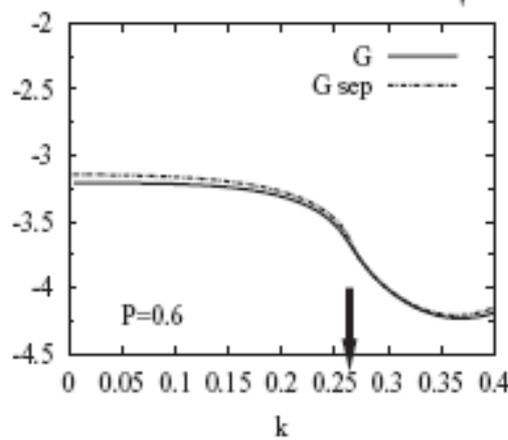
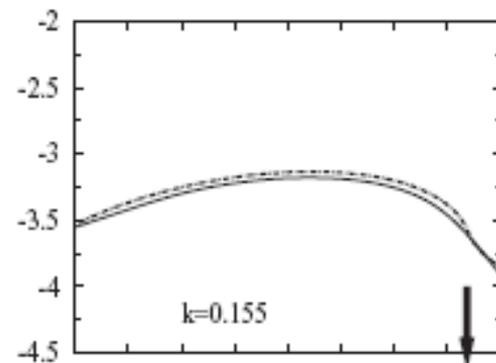
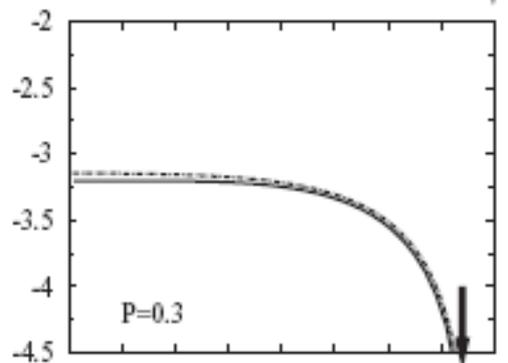
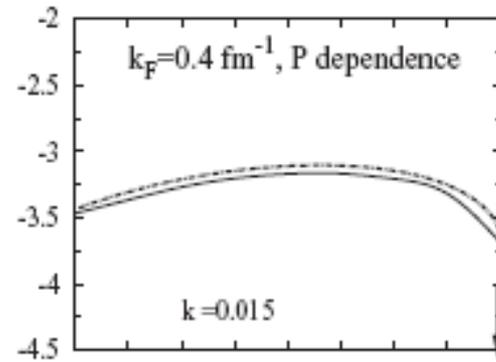
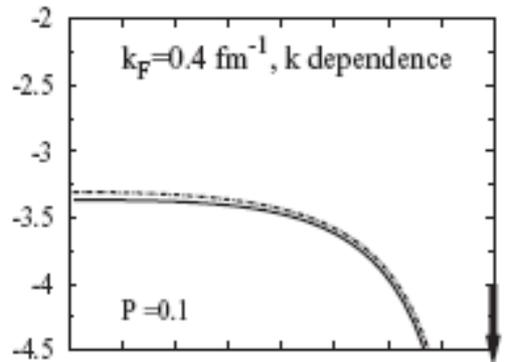
Everything is analytical. The neutron matter energy can be calculated by simple integration.

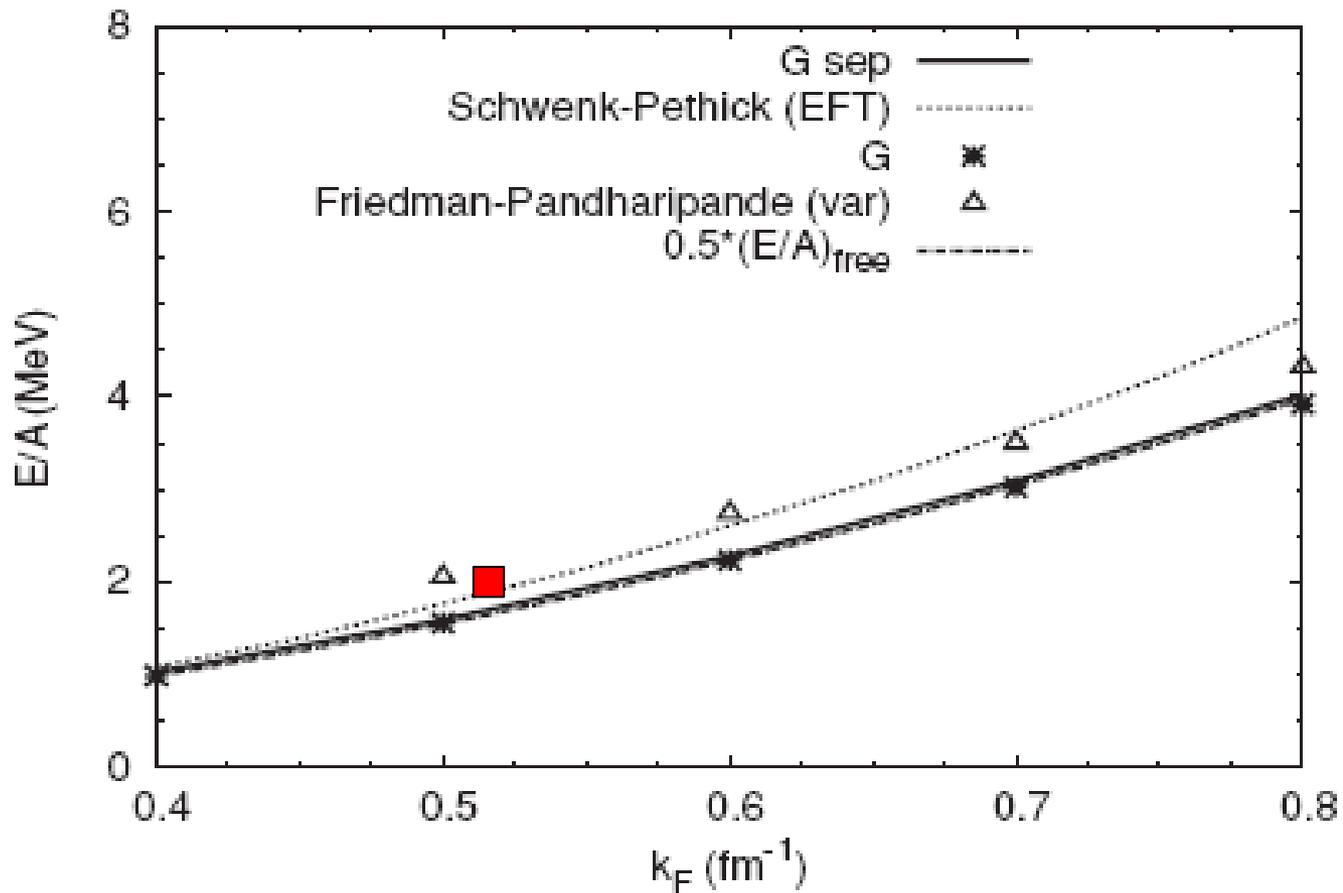
Explicit expression of the separable G-matrix

$$(k|G(P, k_F)|k) = 1/\left[\left(1/a - \frac{1}{2}r_0k^2 + \frac{1}{2}k^4/(b^3\beta)\right) + A(k, P, k_F)\right].$$

$$A = -\frac{1}{\pi b}(b^2 - k^2) \arctan\left(\frac{k_F + P/2}{b}\right) + \frac{1}{\pi}k \log\left(\frac{k + k_F + P/2}{-k + k_F + P/2}\right) + \frac{1}{\pi P}(k_F^2 - P^2/4 - k^2) \times \log\left|\frac{(k_F + P/2)^2 + b^2}{k_F^2 - P^2/4 + b^2} \cdot \frac{k_F^2 - P^2/4 - k^2}{(k_F + P/2)^2 - k^2}\right|,$$

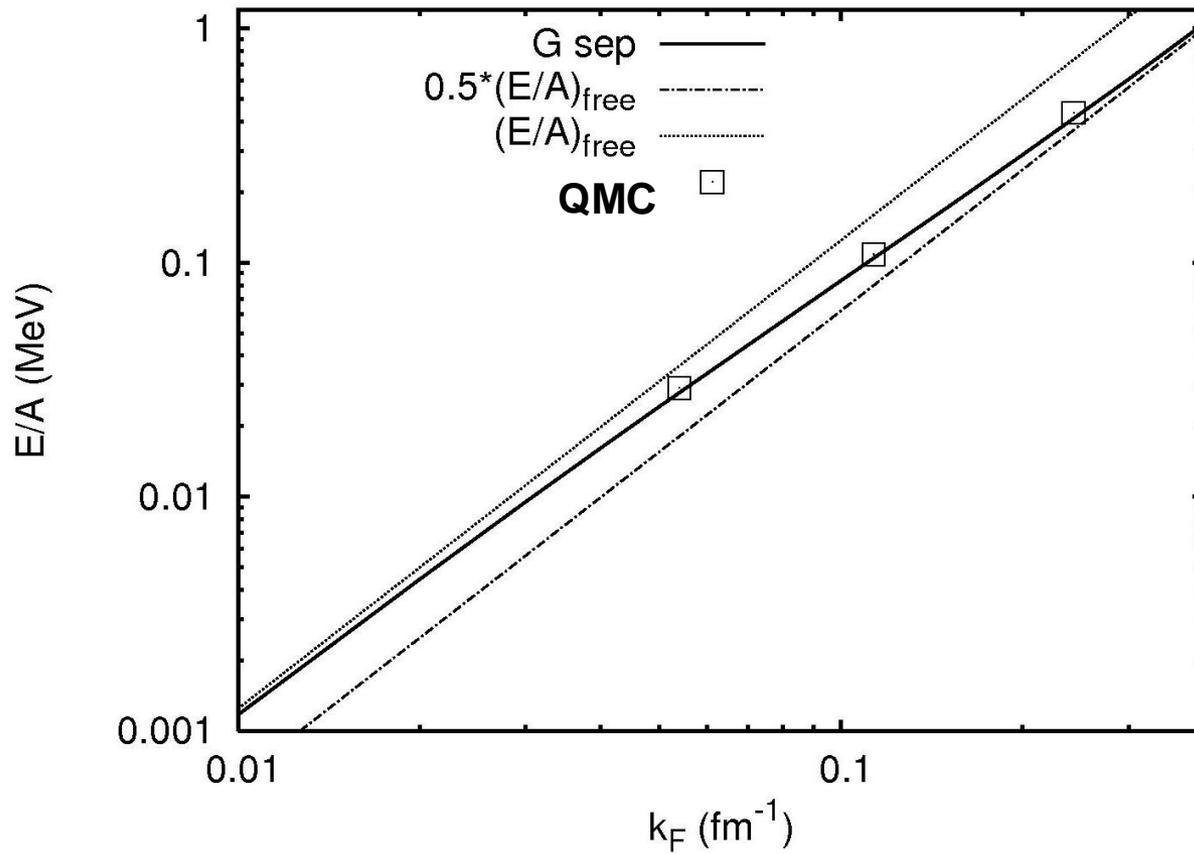
$$a = \frac{1}{b} \left(\frac{2\beta}{1 + \beta}\right); \quad r_0 = \frac{1}{b} \left(\frac{\beta - 2}{\beta}\right).$$





M.B. & C. Maieron, PRC 77, 015801 (2008)

- **A. Gezerlis and J. Carlson, Phys. Rev. C 77,032801 (2008)**
Quantum Monte Carlo calculation



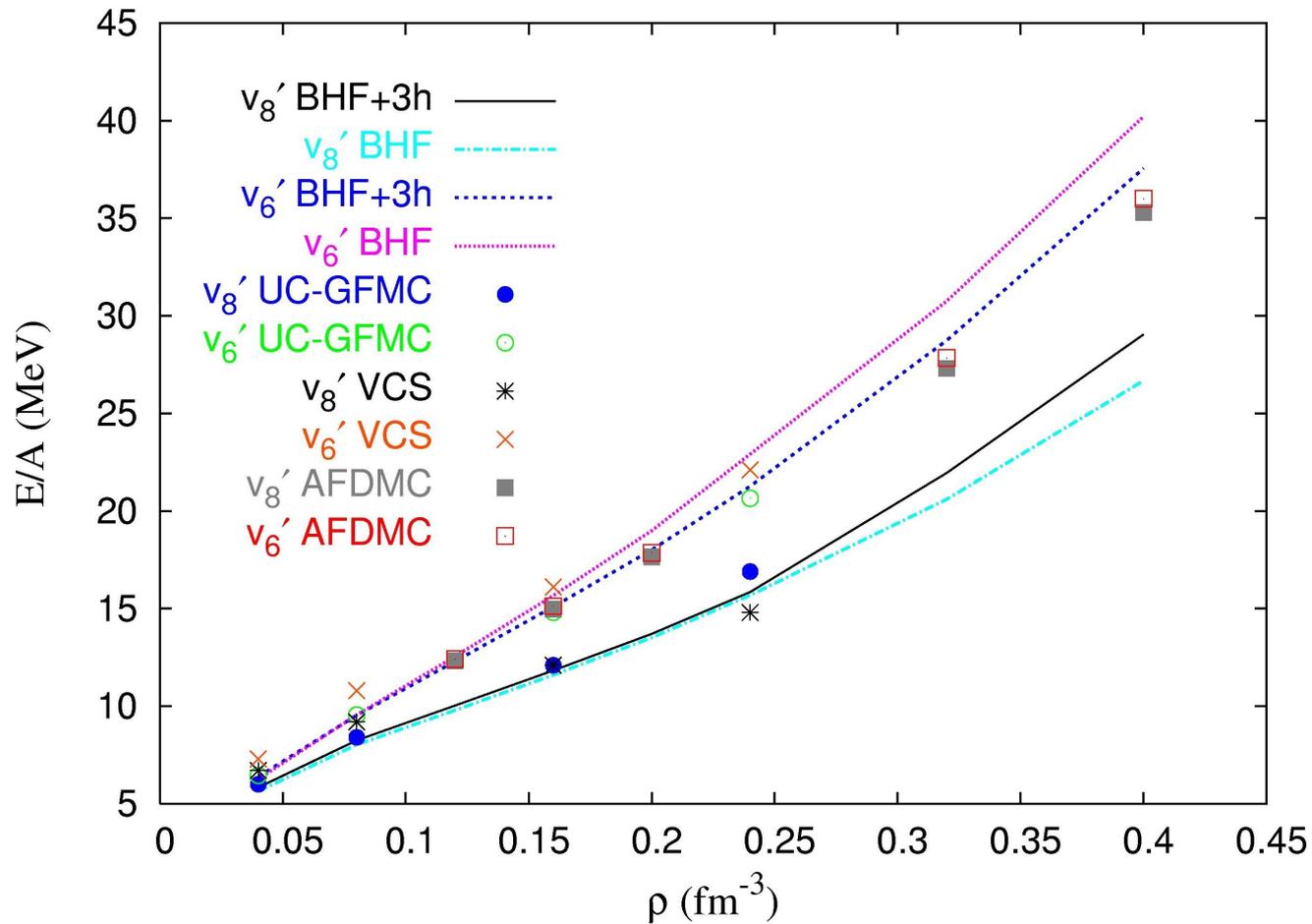
M.B. & C. Maieron, PRC 77, 015801 (2008)

Conclusions for the “very low” density region

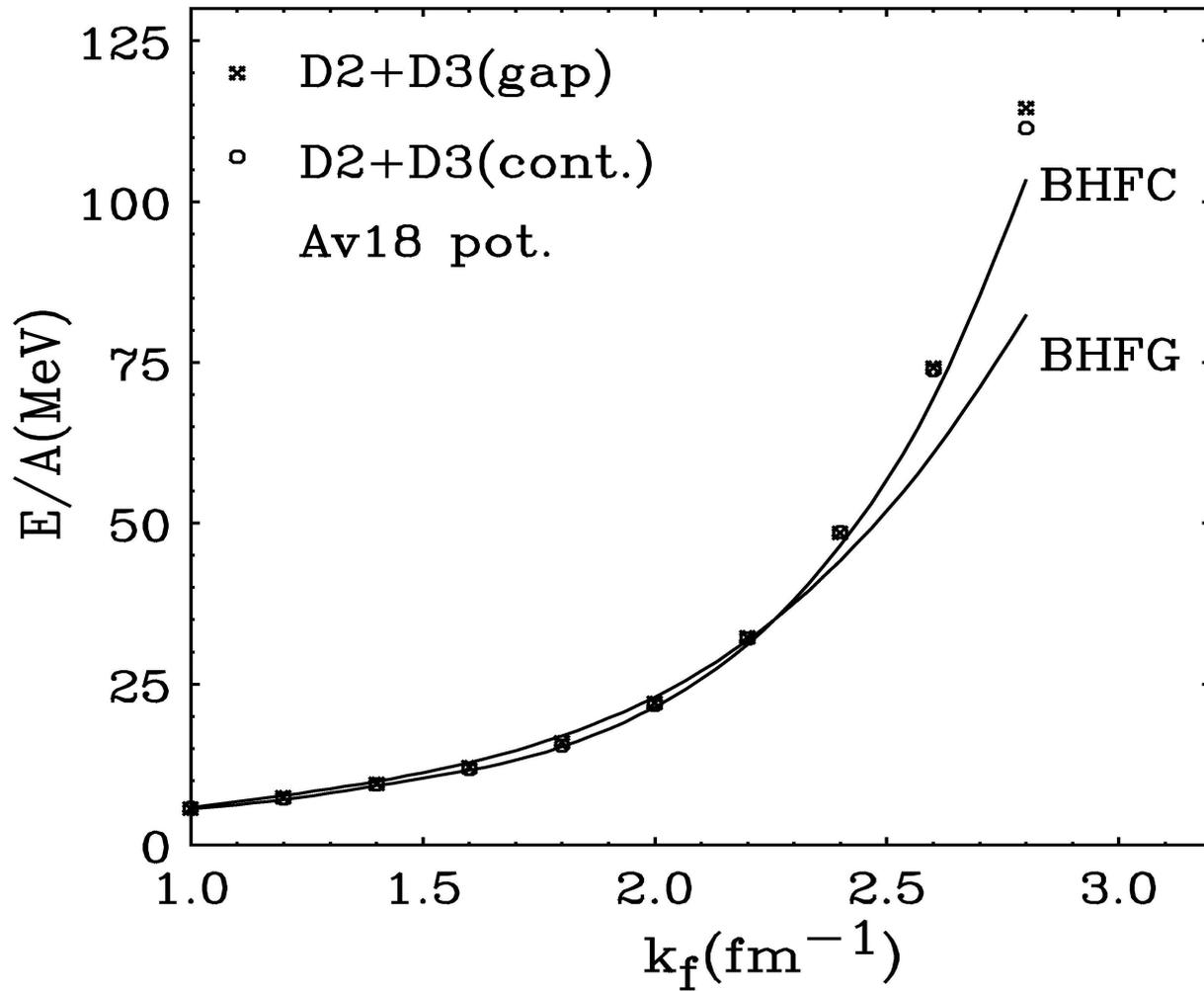
1. Only s-wave matters, but the “unitary limit” is actually never reached. Despite that the energy is $\frac{1}{2}$ the kinetic energy in a wide range of density (for unitary 0.4-0.42 from QMC).
5. The dominant correlation comes from the Pauli operator
7. Both three hole-line and single particle potential effects are small and essentially negligible
10. Three-body forces negligible
12. The rank-1 potential is extremely accurate : scattering length and effective range determine completely the G-matrix.
15. Variational calculations are slightly above BBG.
Good agreement with QMC.

**In this density range one can get the “exact”
neutron matter EOS**

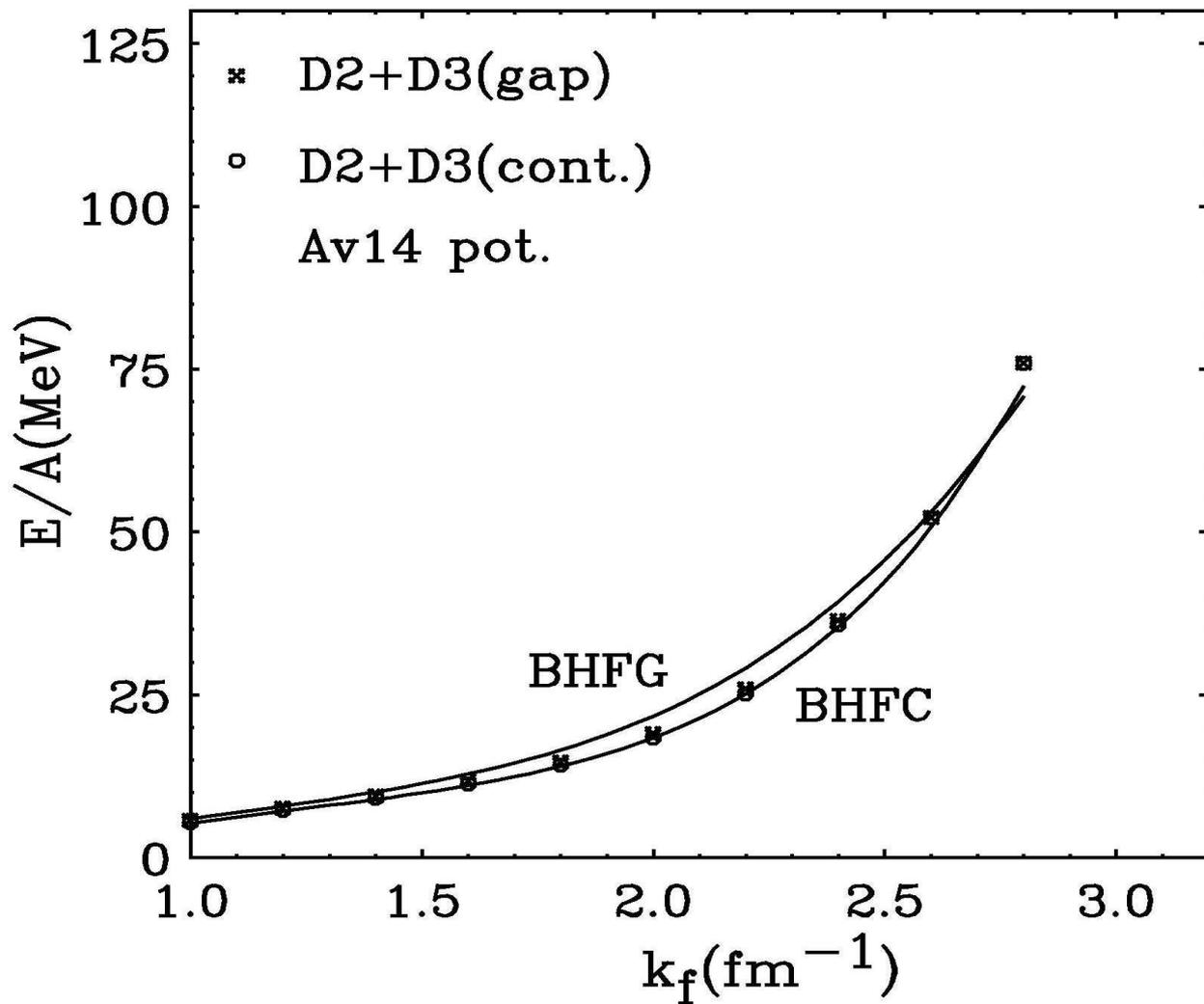
Confronting with “exact” GFMC for v6 and v8



Variational and GMFC : Carlson et al. Phys. Rev. C68, 025802(2003)
BBG : M.B. and C. Maieron, Phys. Rev. C69,014301(2004)



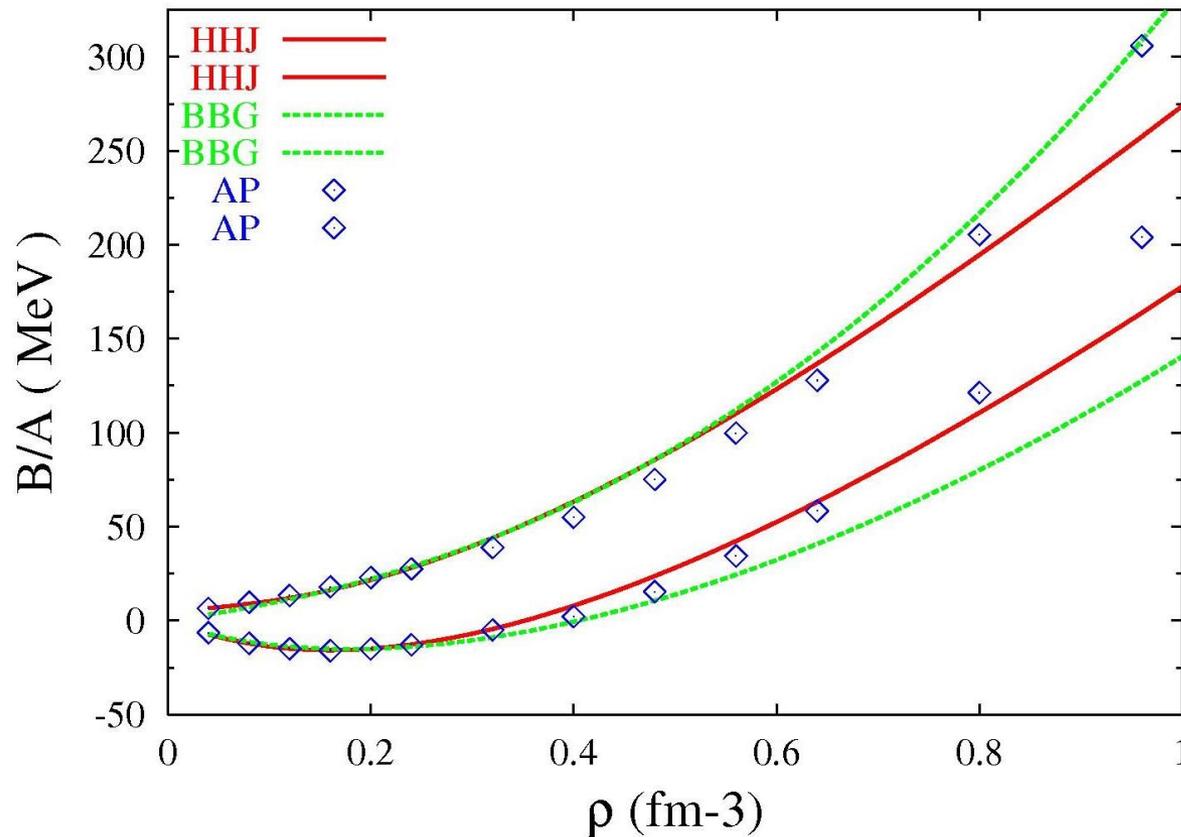
Neutron matter



Neutron matter

The baryonic Equations of State

**V18 NN potential
+ TBF**



HHJ : Astrophys. J. 525, L45 (1999)

BBG : PRC 69 , 018801 (2004)

AP : PRC 58, 1804 (1998)