

3D explosion dynamics of a critical-mass neutron star in a binary system

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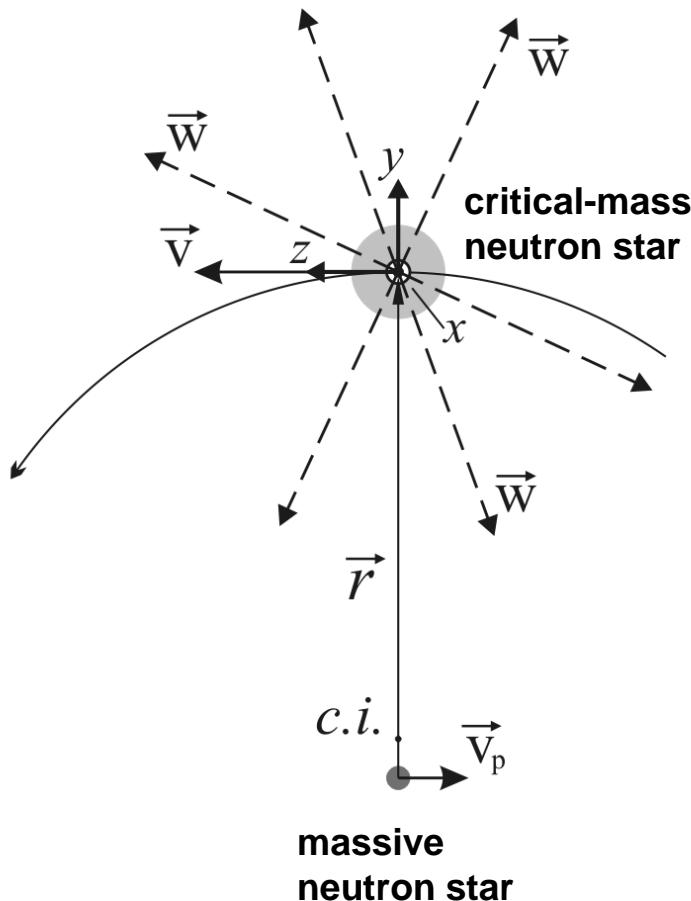
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Problem statement

Imshennik V.S., Manukovskiy K.V., AstrL, 33, 2007



$$a = \frac{m^2}{M+m} \frac{G}{V_p^2} = \frac{M^2}{M+m} \frac{G}{V_{ns}^2}$$
$$[V] = \left(\frac{GM}{a} \right)^{1/2}$$

$$\nu = V_{ns}/[V] = \left(\frac{M}{M+m} \right)^{1/2}$$
$$w = (2\varepsilon_0/m_0)^{1/2}/[V]$$

$$\varepsilon_0 = 4.7 \text{ MэB/nucleon}$$

$$V_p = 1000 \text{ km/s}$$

$$m = 0.1M_\odot$$

$$M/m = 18 : \quad \nu = 0.973$$
$$w = 1.622$$



Analytical solution ($M/m \rightarrow \infty$)

Runge-Lenz vector

$$\mathbf{A} = -GM \frac{\mathbf{r}}{r} + (\mathbf{v} + \mathbf{w}) \times \mathbf{J}$$

$$A = \sqrt{(GM)^2 + 2EJ^2} = GM e \quad \mathbf{J} = \mathbf{r} \times (\mathbf{v} + \mathbf{w})$$

Landau L.D., Lifshitz E.M., Mechanics

Asymptotic velocity

$$\mathbf{v}_\infty = \frac{1}{e} \left[-\sqrt{2E} \hat{\mathbf{A}} + \left(\frac{2EJ}{GM} \right) \hat{\mathbf{J}} \times \hat{\mathbf{A}} \right] \quad (E > 0, t \rightarrow \infty)$$

$$\hat{\mathbf{A}} = \frac{\mathbf{A}}{A} \quad \hat{\mathbf{J}} \times \hat{\mathbf{A}} = \frac{\mathbf{J} \times \mathbf{A}}{JA}$$

$$e^2 = 1 + \frac{2EJ^2}{(GM)^2} \quad (E > 0) \quad 2E = |\mathbf{v} + \mathbf{w}|^2 - \frac{2GM}{r}$$

Colpi M., Wasserman I., *Astrophys. J.*, 581, 1271(2002)



Velocity components

Isotropic explosion

$$w_x = w \sin \theta \cos \varphi$$

$$w_y = w \sin \theta \sin \varphi$$

$$w_z = w \cos \theta$$

$$v_{x\infty} = \frac{(w \sin \theta \cos \varphi) \sqrt{\Phi}}{1 + \Phi \Psi} \left[w \sin \theta \sin \varphi + \sqrt{\Phi} (\Psi - 1) \right]$$

$$v_{y\infty} = \frac{\sqrt{\Phi}}{1 + \Phi \Psi} \left[\sqrt{\Phi} \Psi w \sin \theta \sin \varphi - (\Psi - 1) \right]$$

$$v_{z\infty} = \frac{(v + w \cos \theta) \sqrt{\Phi}}{1 + \Phi \Psi} \left[w \sin \theta \sin \varphi + \sqrt{\Phi} (\Psi - 1) \right]$$

$$\Psi = (v + w \cos \theta)^2 + w^2 \sin^2 \theta \cos^2 \varphi$$

$$\Phi = (w^2 + v^2 + 2wv \cos \theta) - 2$$



Energy spectrum

Velocity magnitude $v_\infty^2 = \Phi = (v^2 + w^2 + 2vw\cos\theta) - 2$

Critical angle $\cos\theta_{cr} = \frac{2 - w^2 - v^2}{2vw}$ $\sqrt{2} - v \leq w \leq \sqrt{2} + v$

Energy spectrum $f(E) = \frac{1}{(1 - \cos\theta_{cr})vw} = \frac{2}{(v + w)^2 - 2} = const$

Maximum energy

hyperbolic tracks

$$e_{\max}^{hyp} = \frac{1}{2}(v + w)^2 - 1$$

elliptic tracks

$$e_{\max}^{ell} = \frac{1}{2} \left(\frac{GM}{J} \right)^2 (e + 1)^2 \xrightarrow[J \rightarrow 0]{} \infty$$



kinetic energy

Initial energy $e_0 = \frac{m}{2} \left(\frac{GM}{a} \right) (v^2 + w^2)$

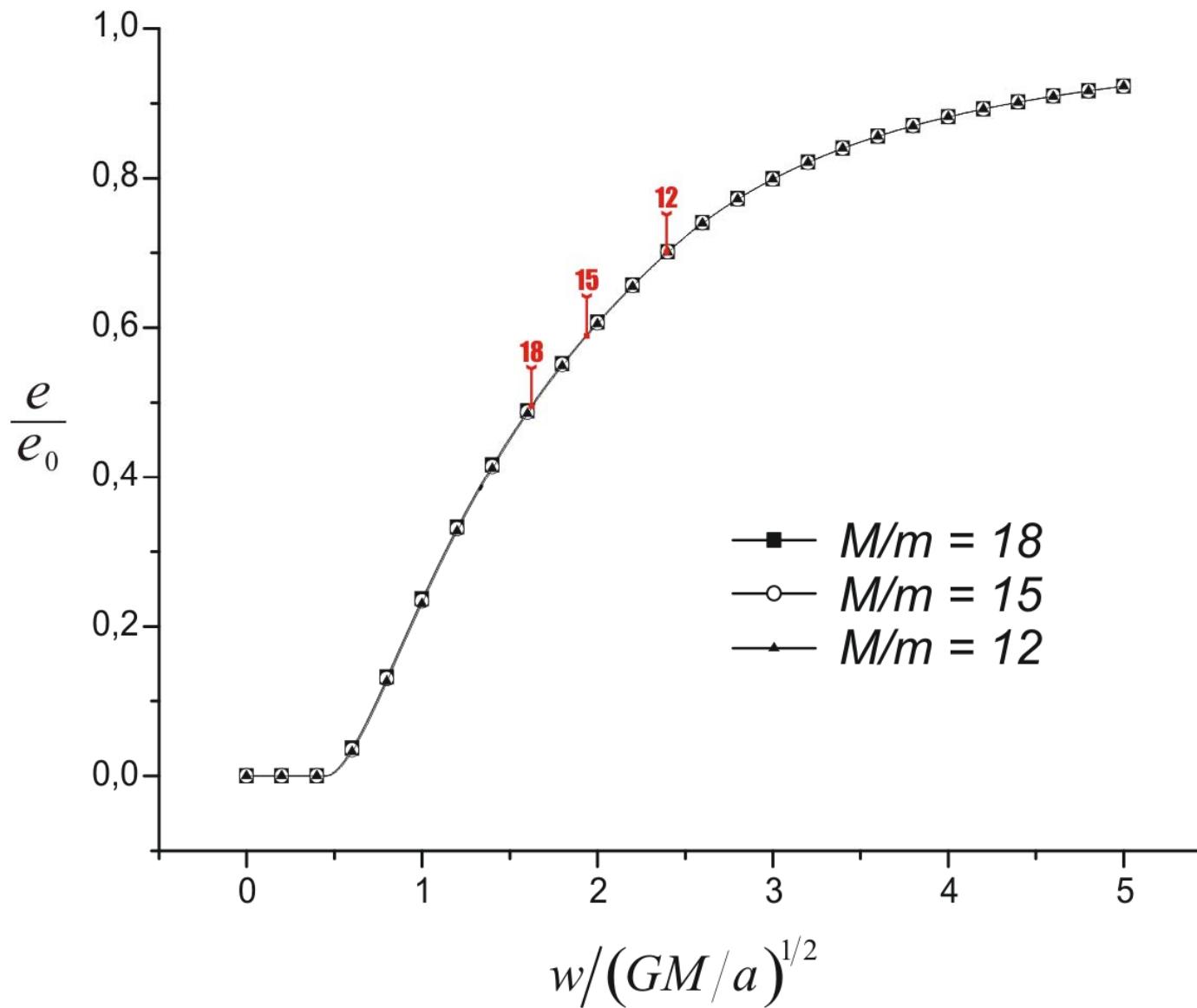
Final energy

$$e = \frac{m}{4} (1 - \cos \theta_{cr}) \left[(v^2 + w^2 - 2) + wv(1 + \cos \theta_{cr}) \right] \left(\frac{GM}{a} \right)$$

$$\begin{cases} e = \frac{m}{2} \left(\frac{GM}{a} \right) (v^2 + w^2 - 2), & w \geq \sqrt{2} + v \\ e = \frac{m}{16wv} \left(\frac{GM}{a} \right) \left[(v + w)^2 - 2 \right]^2, & \sqrt{2} - v \leq w < \sqrt{2} + v \\ e = 0, & w < \sqrt{2} - v \end{cases}$$



Kinetic energy



Recoil momentum

$$(M + \Delta m)v'_p + (m - \Delta m)v_e = 0 \quad \Delta m/m = \chi = \frac{1 + \cos \theta_{cr}}{2}$$

$$v'_p = \frac{1}{M/m + \chi} (g^2 + f^2)^{1/2}$$

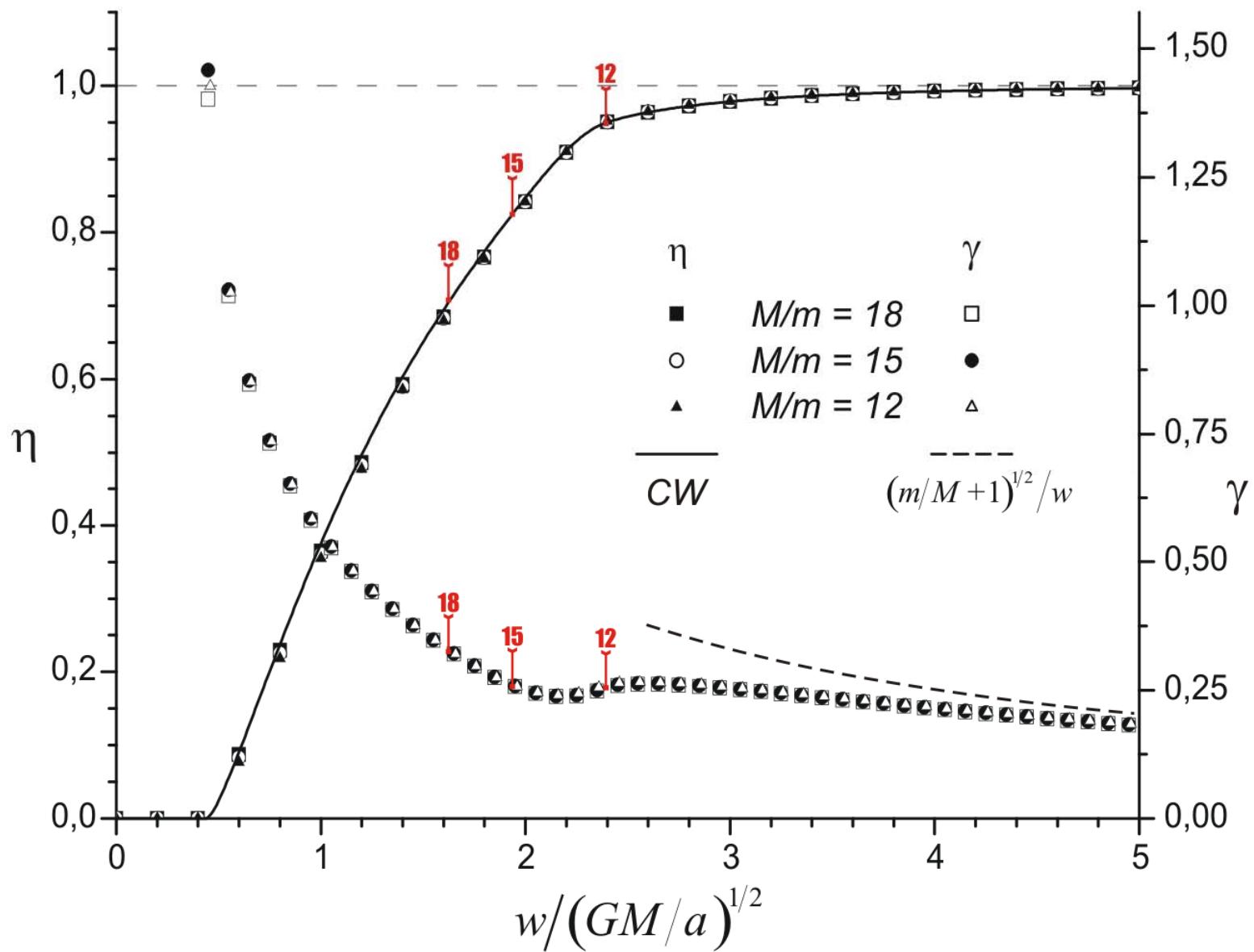
$$f = \frac{1}{4\pi} \int_0^{\theta_{cr}} (\nu + w \cos \theta) \Phi \sin \theta d\theta \int_0^{2\pi} \left(\frac{\Psi - 1}{1 + \Phi \Psi} \right) d\phi$$

$$g = -\frac{1}{4\pi} \int_0^{\theta_{cr}} \sqrt{\Phi} \sin \theta d\theta \int_0^{2\pi} \left(\frac{\Psi - 1}{1 + \Phi \Psi} \right) d\phi$$

Normalized velocity $\eta = \frac{v'_p}{v_p} = \frac{v'_p}{\frac{m}{M}\nu} = \frac{1}{1 + \chi} \frac{1}{\frac{m}{M}\nu} (g^2 + f^2)^{1/2}$



Pulsar velocity and angle of rotation



Simulation using particles

equations

$$\frac{d\mathbf{x}_i}{dt} = \mathbf{u}_i$$

$$m_i \frac{d\mathbf{u}_i}{dt} = \mathbf{F}_i$$

$$\mathbf{F}_i = \frac{G m_i M}{|\mathbf{r}_i - \mathbf{r}_p|^3} (\mathbf{r}_p - \mathbf{r}_i) \\ i = 1, \dots, N$$

$$M \frac{d\mathbf{u}_p}{dt} = - \sum_{i=1}^N \mathbf{F}_i$$

schemes

2nd-order leap-frog scheme

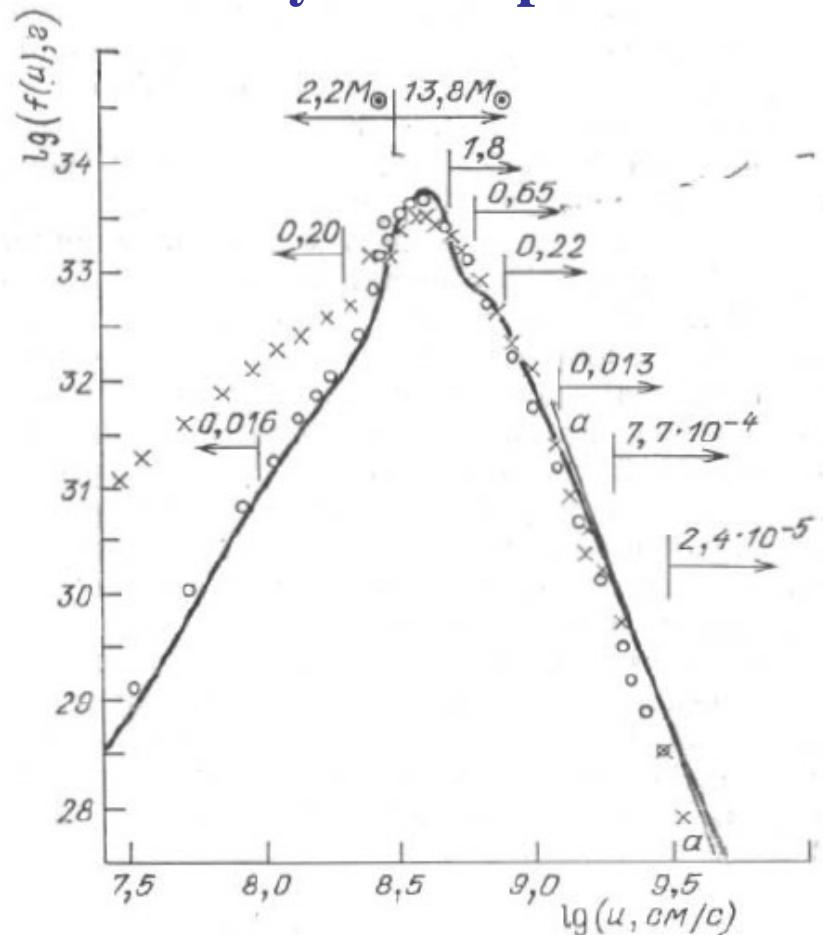
4th-order Runge-Kutta scheme

5th-order Runge-Kutta-England scheme
with automatic step selection

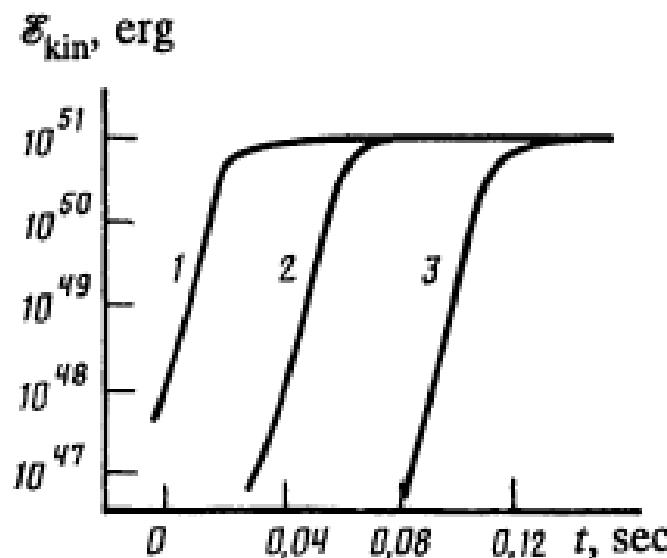


Simulation setup

Velocity radial profile



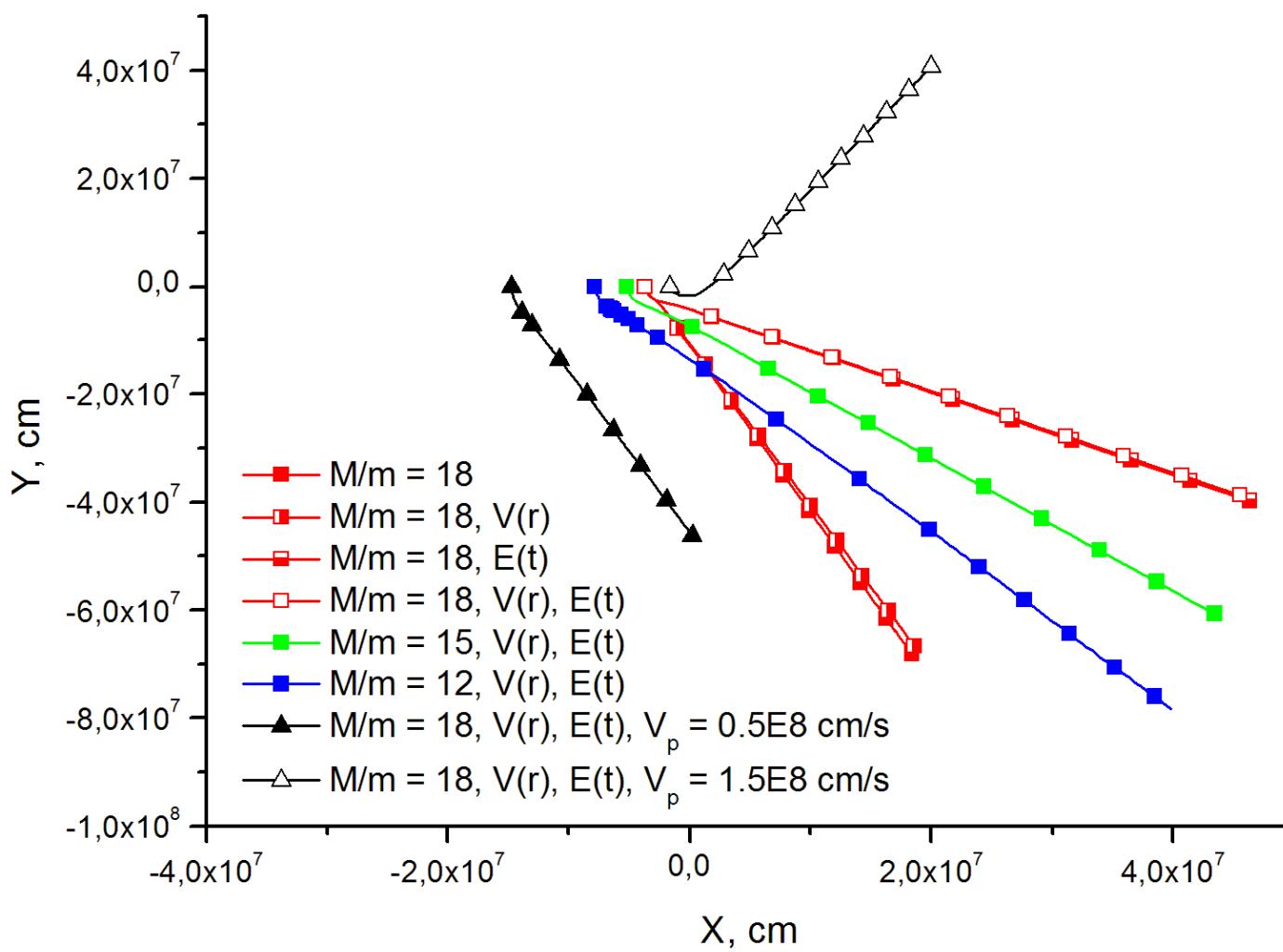
Energy release



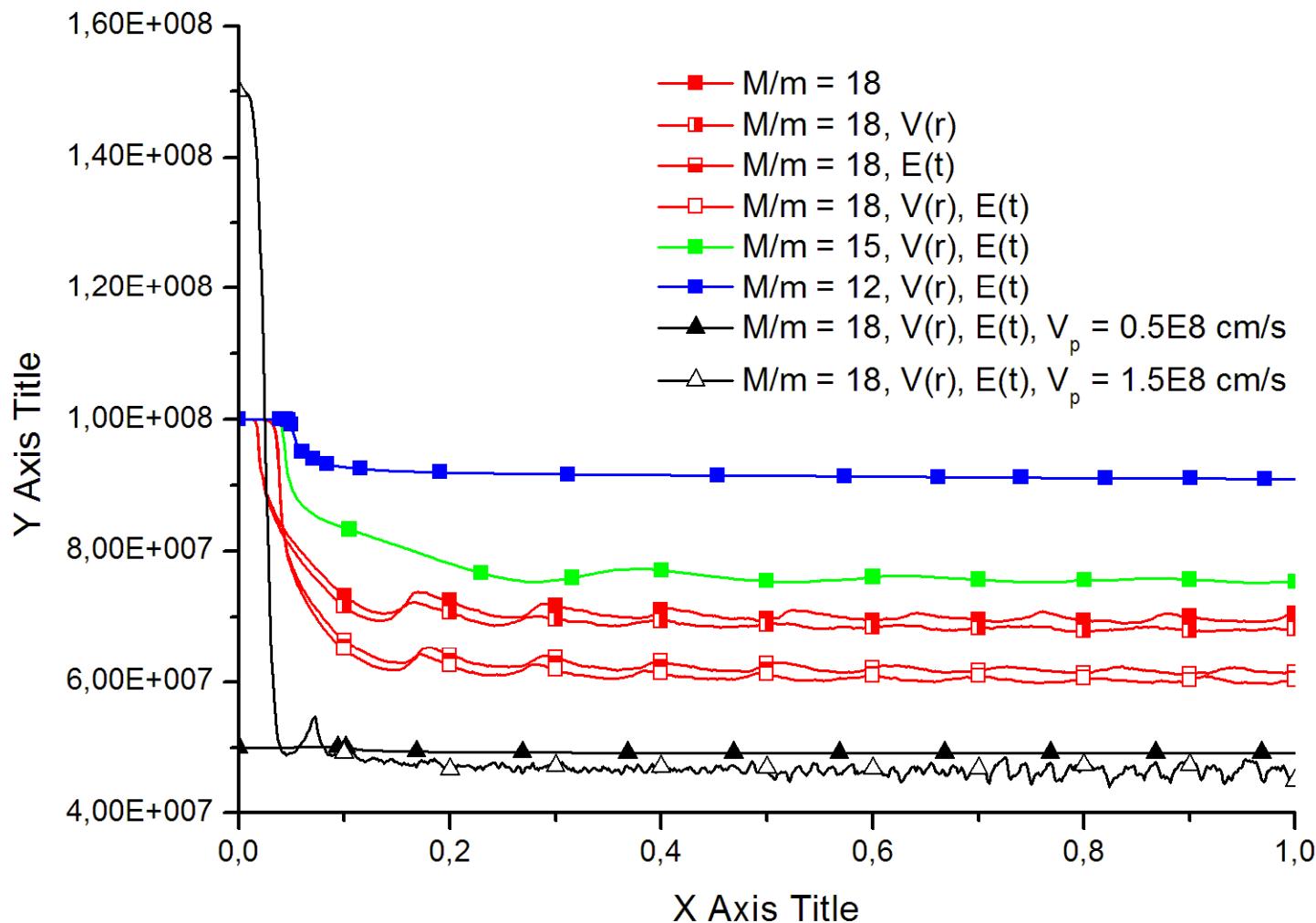
Blinnikov S.I. et al.,
Soviet Astr., 34, 1990

Imshennik, Nadyozhin,
Sov. Sci. Rev. E, 8 , 1989

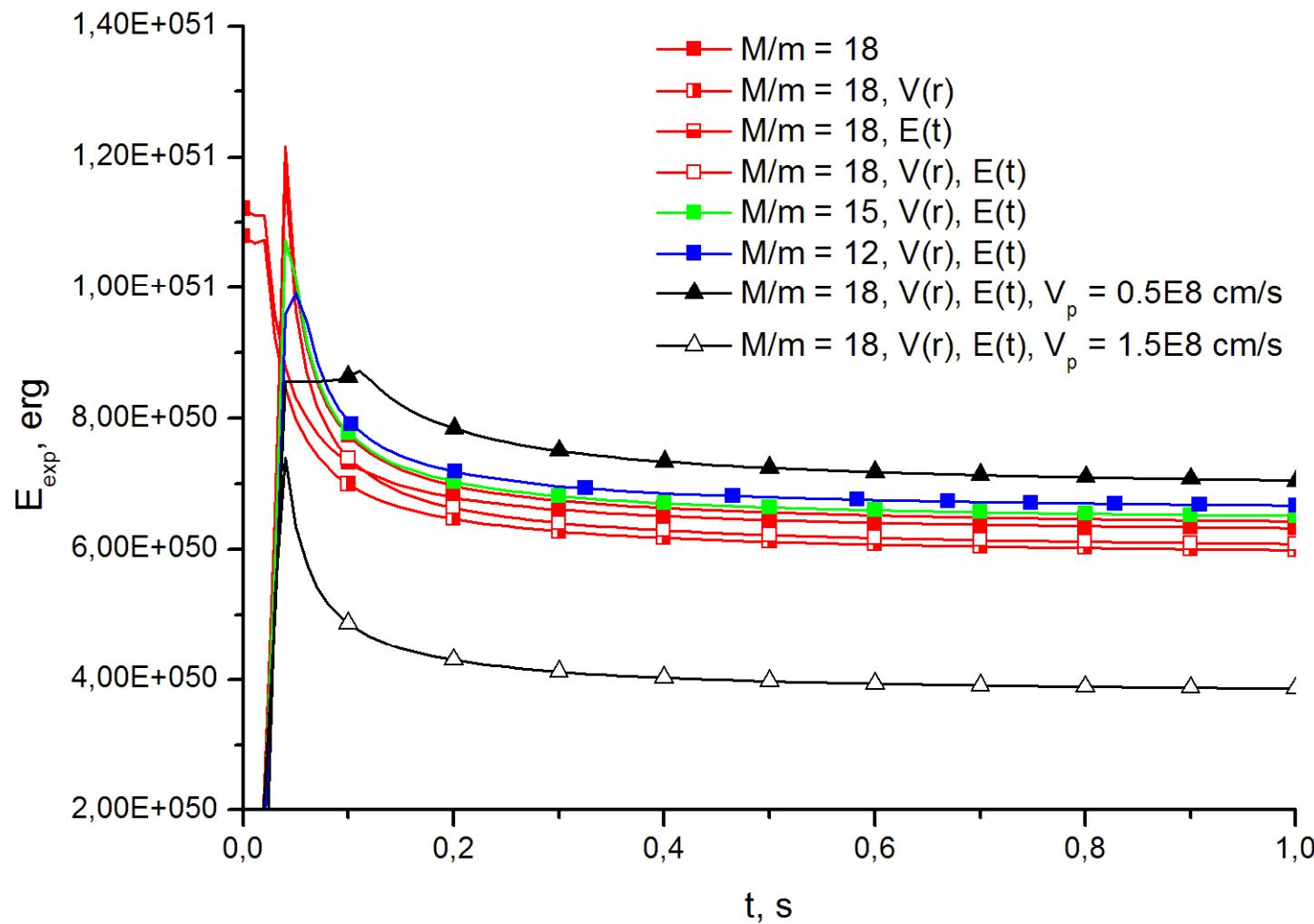
Pulsar track



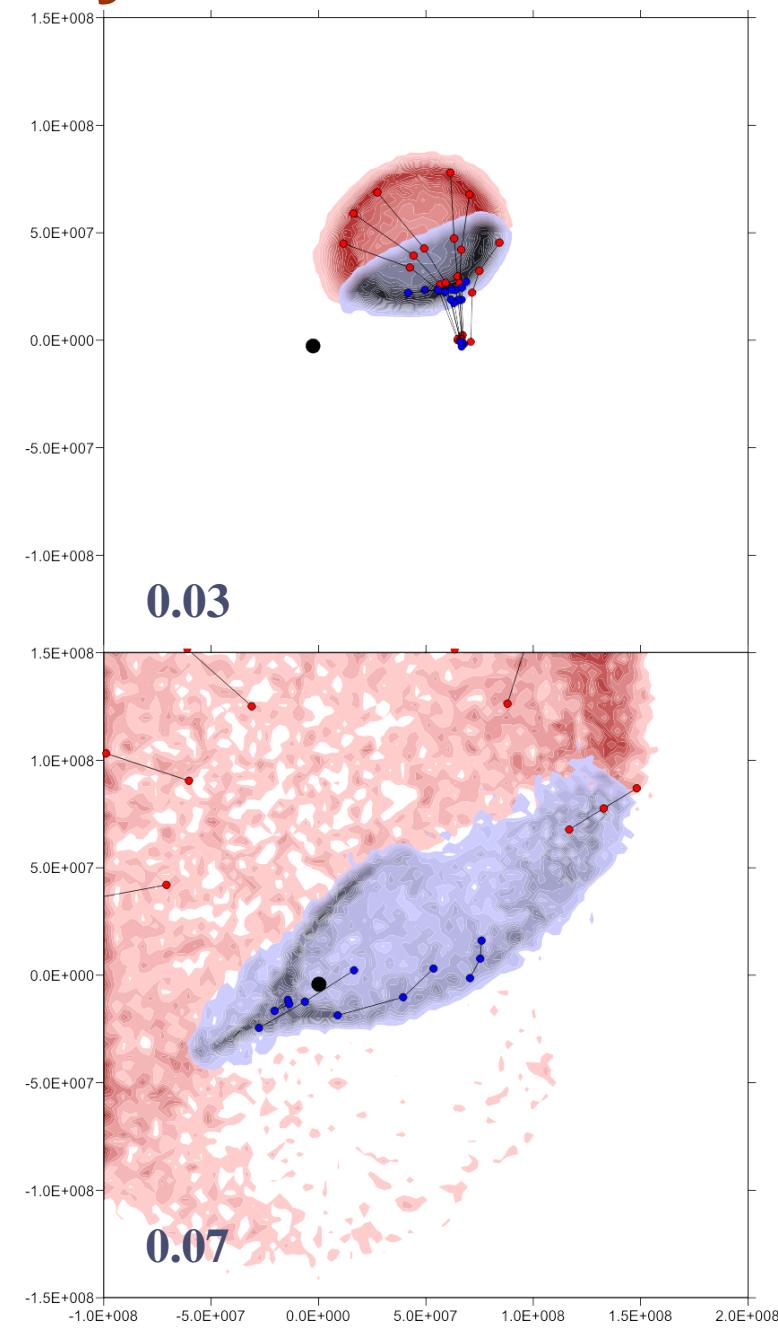
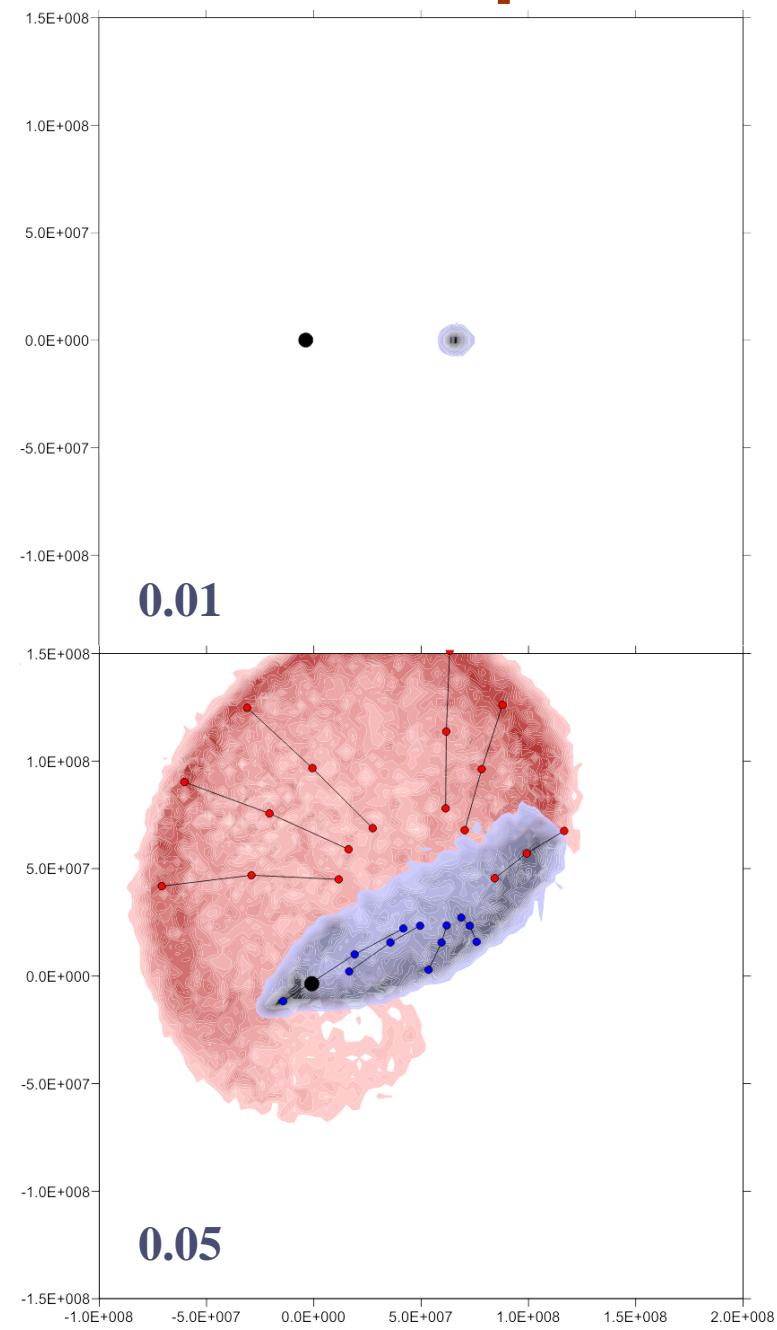
Pulsar velocity



Explosion energy



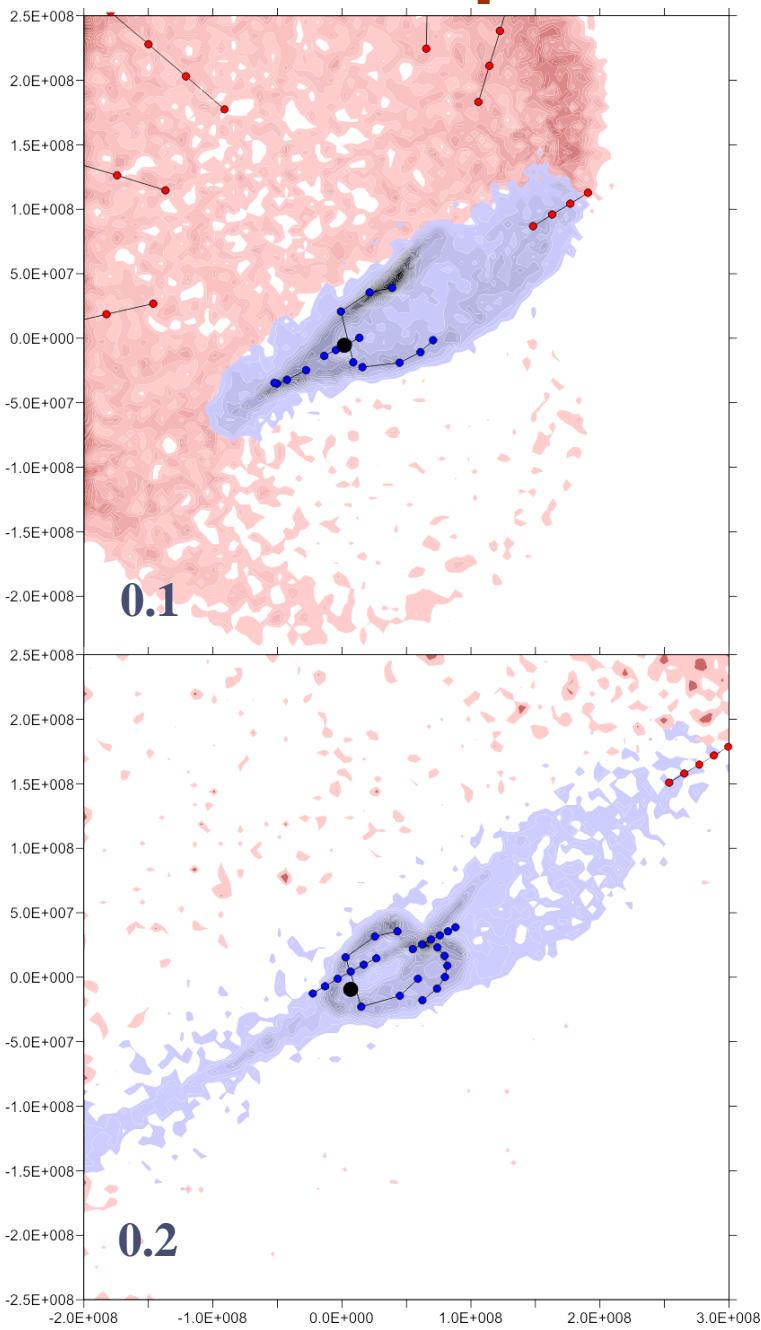
Explosion dynamics



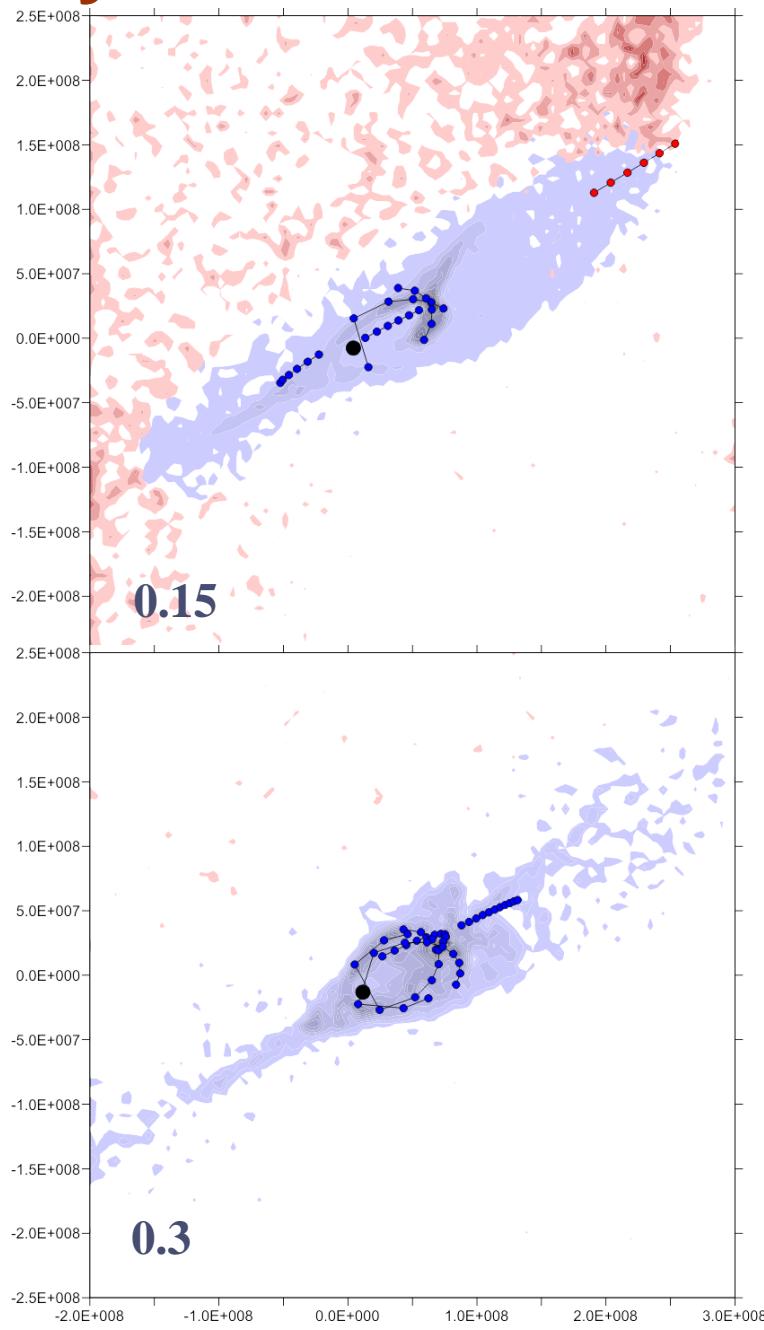
0.05

0.07

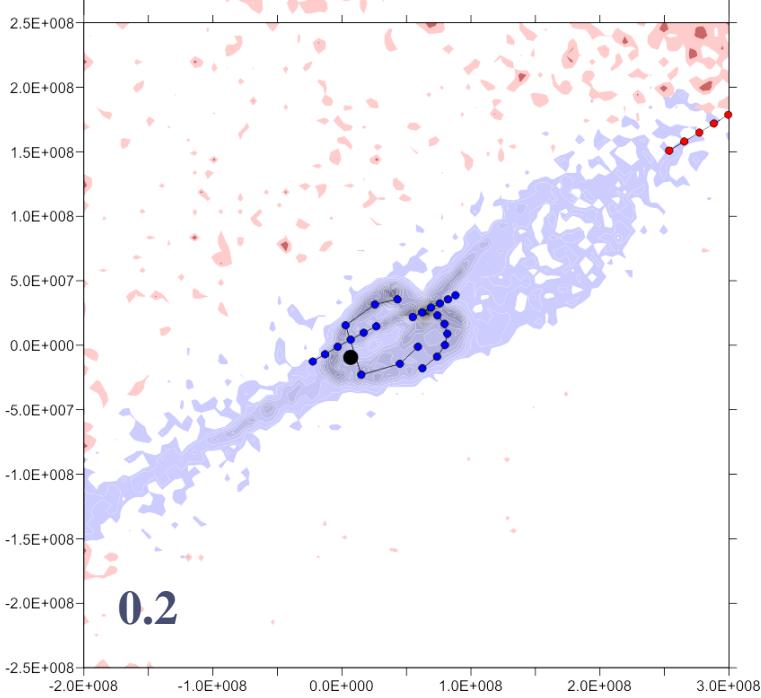
Explosion dynamics



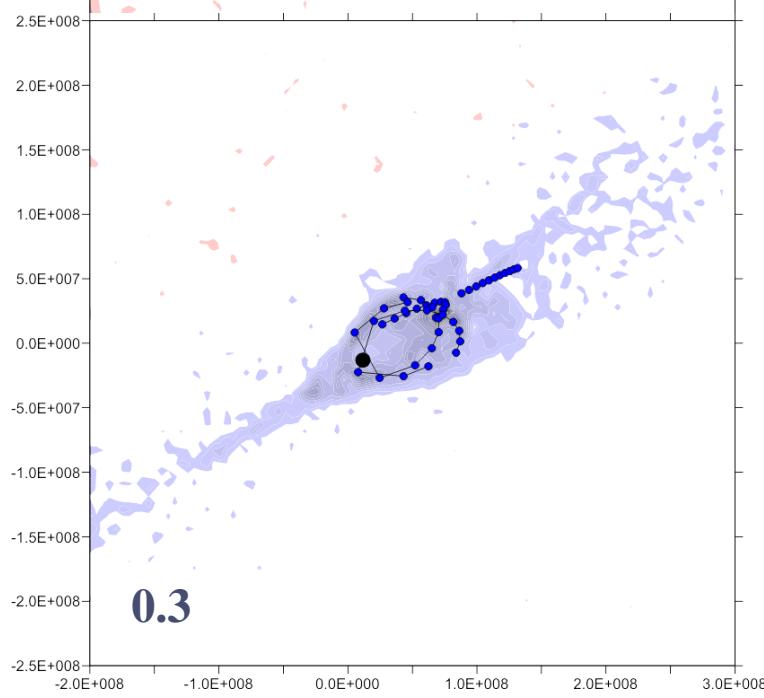
0.1



0.15



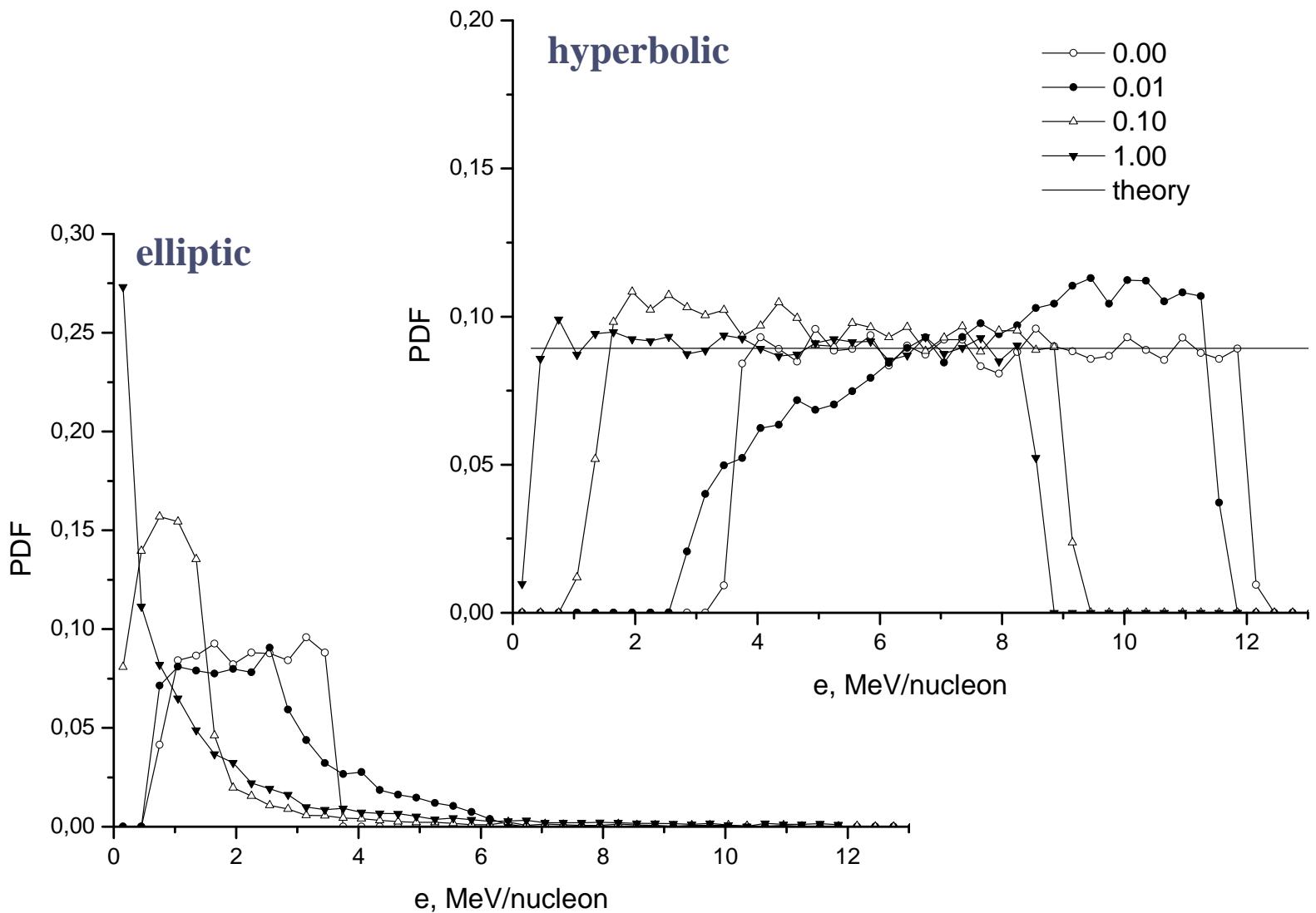
0.2



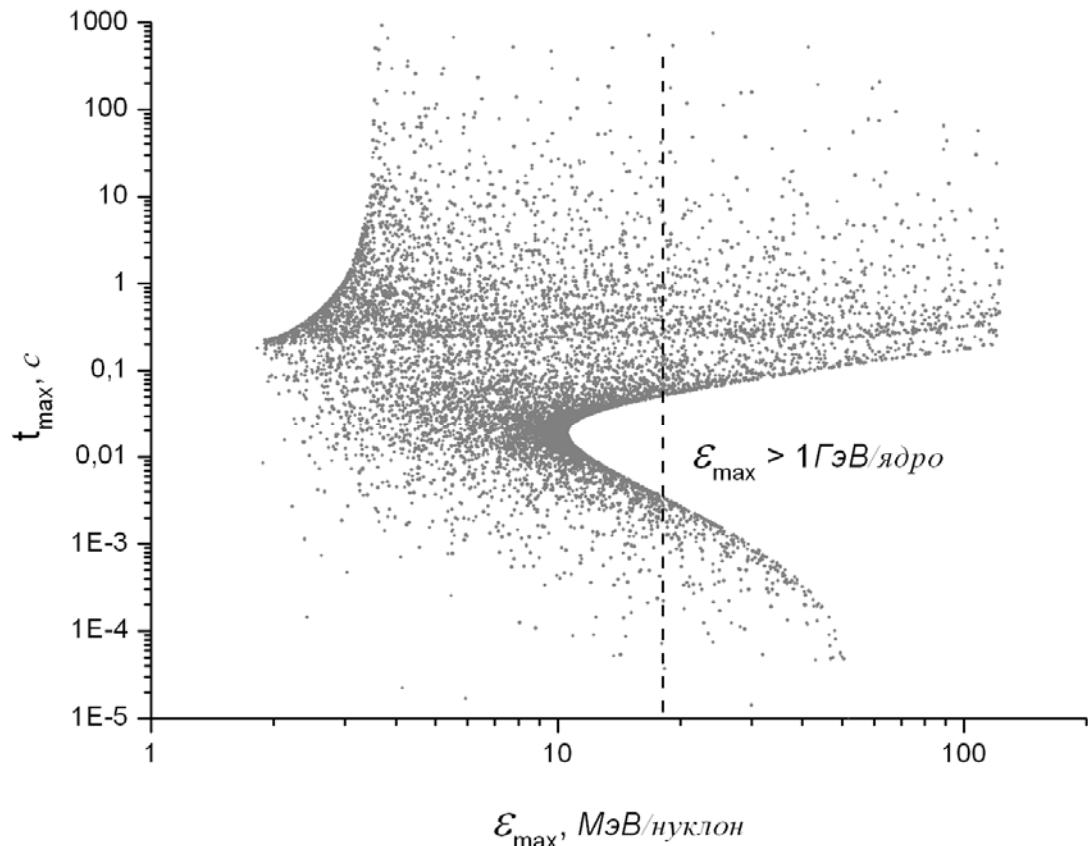
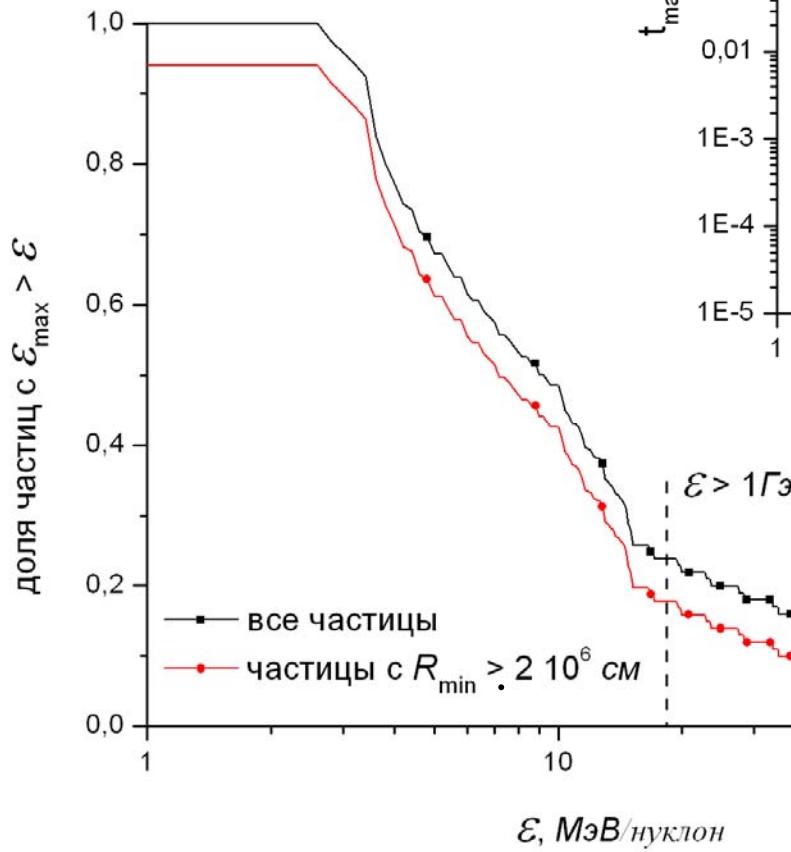
0.3



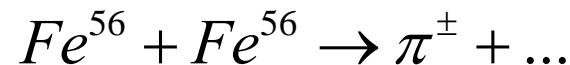
Energy spectrum



Captured matter



Muon neutrinos



$$\sigma(E_{Fe}) - ?? \quad \left(\text{for } E_{Fe} \geq E_{Fe\,th} \right)$$

Ryazhskaya O.G., UFN, 2007

main reaction

