

# Simultaneous XMM-Newton Radio Observations of the Mode-switching Radio Pulsar PSR B1822-09

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Netherlands Organisation for Scientific Research (NWO)

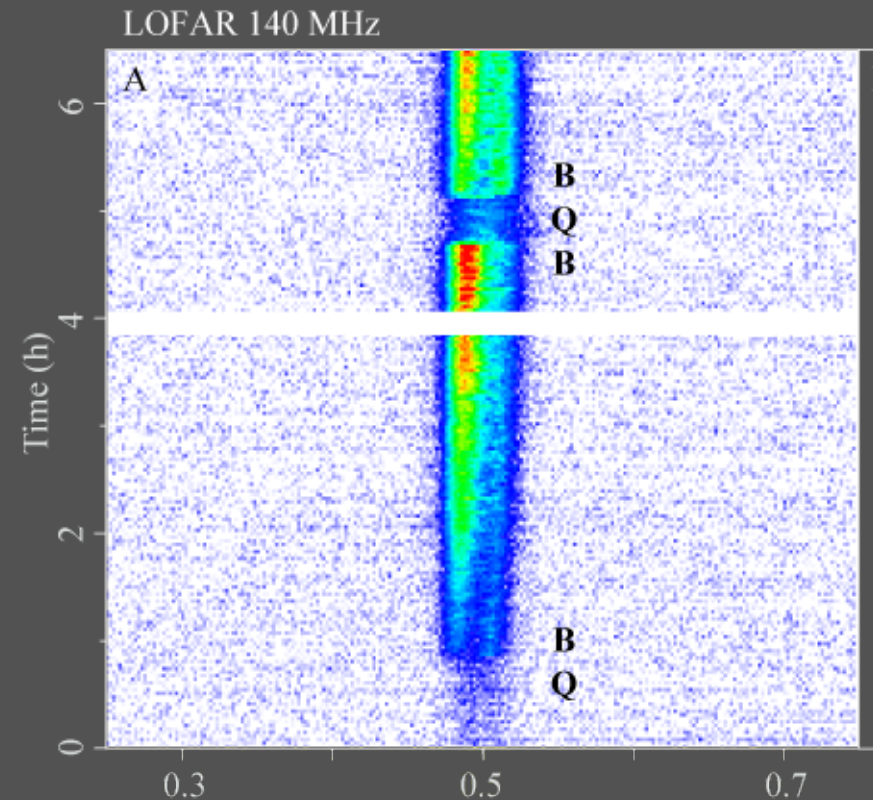
# Outline

- Introduction: XMM-Newton revealed synchronous X-ray and radio-mode changing in pulsar PSR B0943+10.  
(Hermsen et al. 2013, Science 339, 436)
- Simultaneous radio and X-ray observations of PSR B1822-09: results
- Conclusions / Dilemma's

# PSR B0943+10; 6x6 hrs of simultaneous XMM-Newton X-ray and GMRT/LOFAR radio observations

## Characteristics of PSR B0943+10

- $\dot{P} = 1.10 \text{ s}$
- $\dot{P} = 3.5 \times 10^{-15}$
- $\dot{E} = 1.0 \times 10^{32} \text{ ergs s}^{-1}$
- $B_p = 2 \times 10^{12} \text{ G}$
- $T = 5.0 \times 10^6 \text{ yr}$
- **Nearly aligned rotator**
- **mode switching between radio B(right) and Q(quiet) modes**



# XMM-Newton **PN**, Maximum-Likelihood Maps, 0.2-10 keV

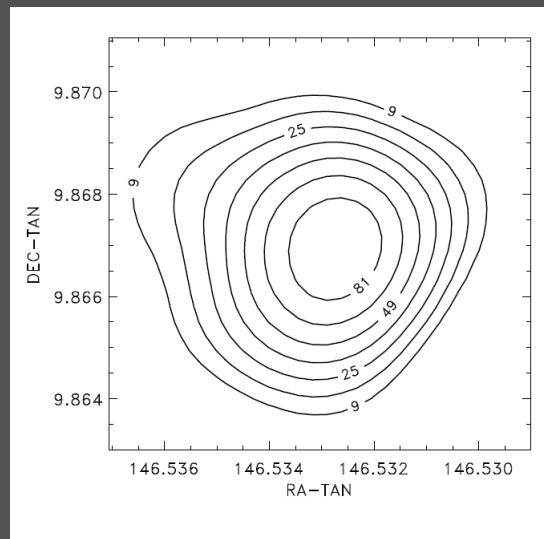
**B-mode windows**

$t_{\text{eff}} 39.7 \text{ ks}$

9.9  $\sigma$  detection

$174 \pm 36 \text{ cnts}$

$(0.44 \pm 0.07)$   
 $10^{-2} \text{ cnts/s}$



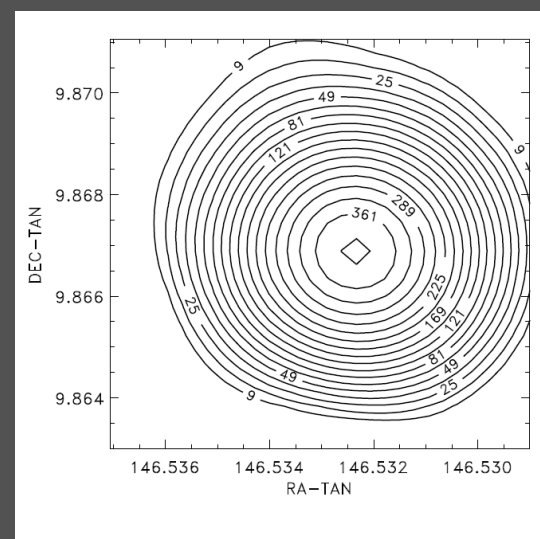
**Q-mode windows**

$t_{\text{eff}} 43.5 \text{ ks}$

20  $\sigma$  detection

$470 \pm 33 \text{ cnts}$

$(1.08 \pm 0.08)$   
 $10^{-2} \text{ cnts/s}$



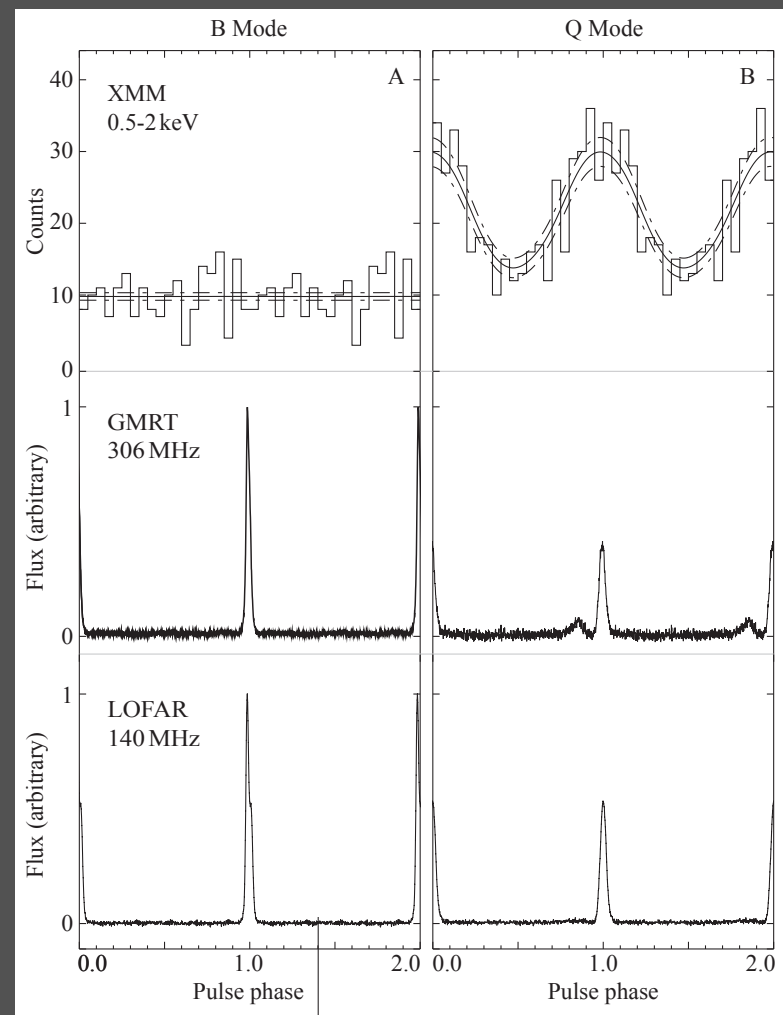
Discovery of correlated Radio – X-ray mode changes  
Anti correlation !

# Discovery of X-ray pulsation only in **radio Q mode intervals** !

XMM-Newton  
EPIC PN + MOS-1  
+ MOS-2

Difference between X-ray emissions in radio B and Q mode is addition of a pulsed X-ray component in Q mode !

X-ray pulse aligned with radio main pulse with precursor

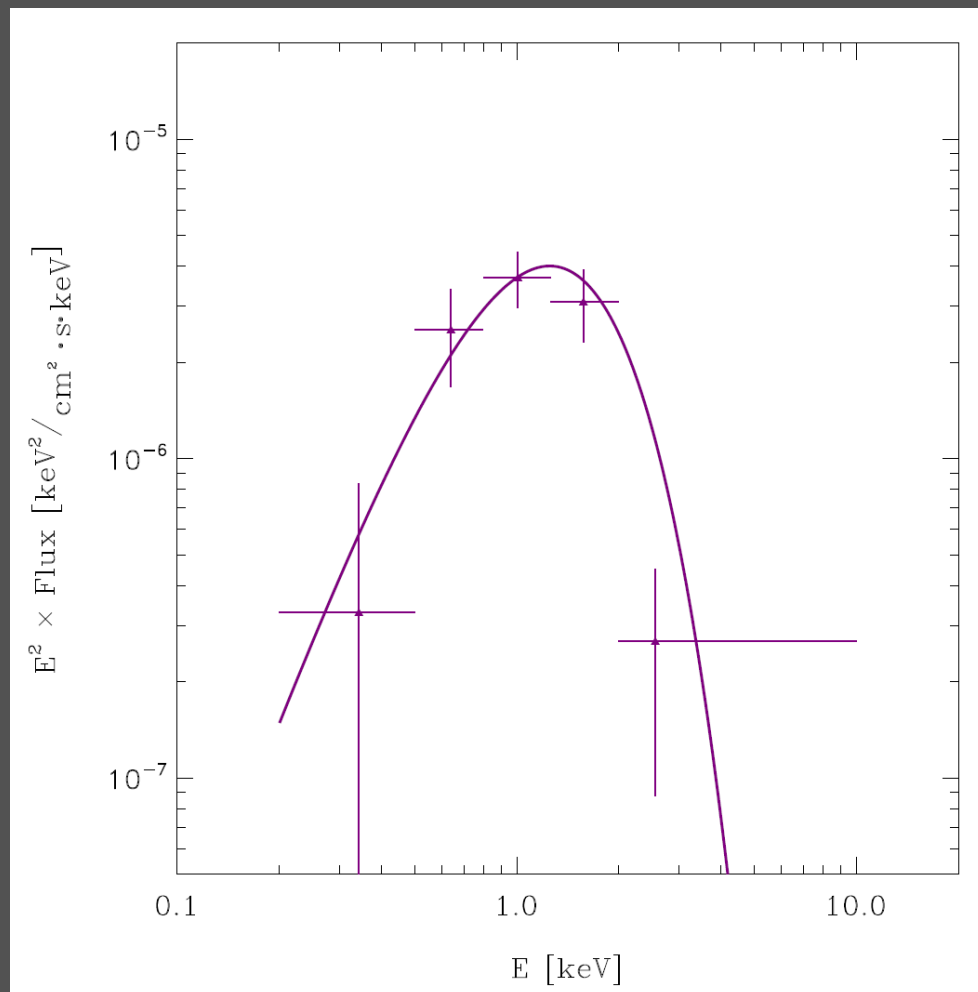


## Pulsed fractions

. Pulsed fractions of PSR B093+10 as a function of energy, defined as the ratio of the flux in the pulse profile over the total flux of the point source measured in the sky maps. Errors are  $1 \sigma$ .

Energy interval keV	Pulsed fraction
0.2 - 0.5	$0.10 \pm 0.15$
0.5 - 0.8	$0.44 \pm 0.16$
0.8 - 1.3	$0.62 \pm 0.14$
1.3 - 2.0	$0.60 \pm 0.17$
2.0 - 10	$0.72 \pm 0.53$

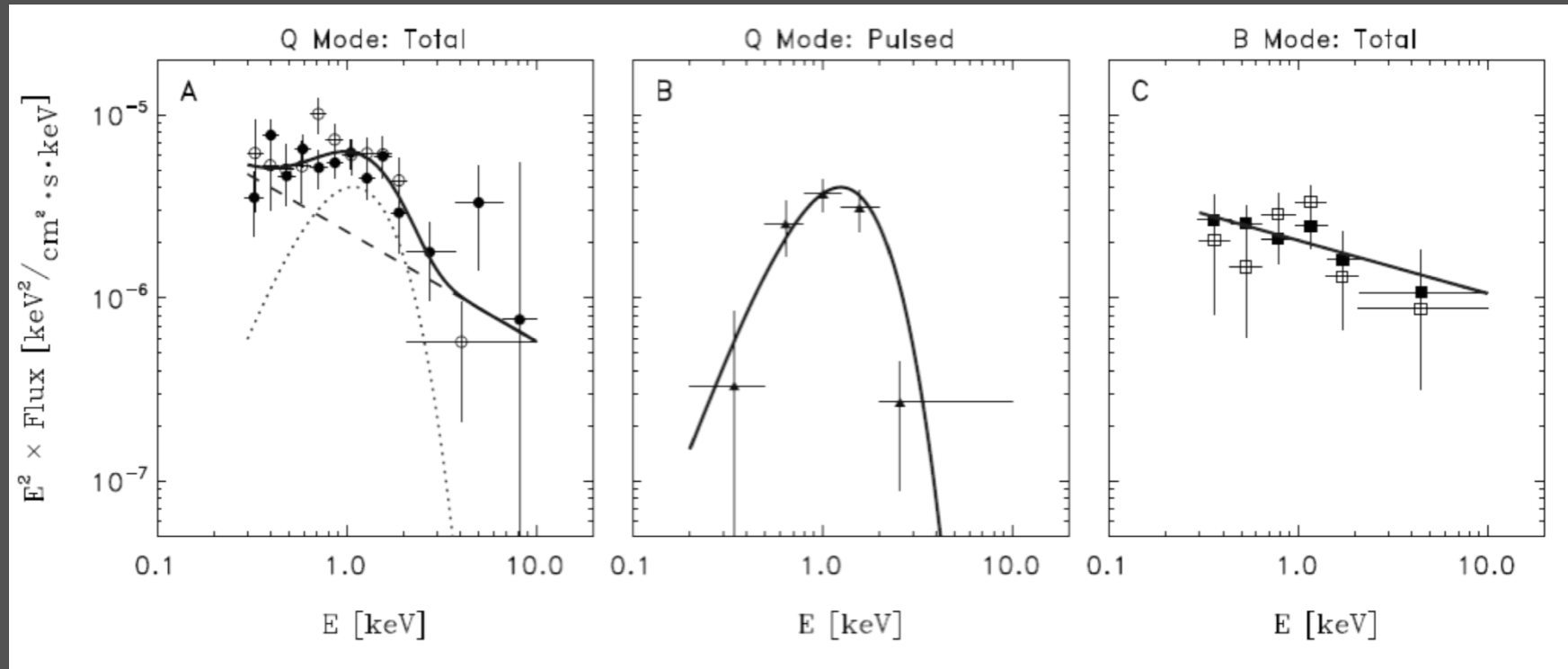
## Pulsed emission X-ray spectrum of PSR B0943+10: radio **Q-mode**



- **Best fit: BB;  $\chi^2/\nu = 1.14/3$ ,  $\sim 78\%$**
- $N_{\text{H}} = 4.3 \times 10^{20} \text{ cm}^{-2}$  (fixed)
- BB:  $kT = 0.319 \pm 0.012 \text{ keV}$
- $F_{\text{BB}} (0.5\text{-}8 \text{ keV}) = (7.8 \pm 1.6) 10^{-15} \text{ erg cm}^{-2} \text{ s}^{-1}$  (unabsorbed)
- $R_{\text{hot spot}} \approx 18 \text{ m}$  ( $d = 630 \text{ pc}$ )

PL fit;  $\chi^2/\nu = 9.50/3$ ,  $\sim 2.3\%$

# Conclusions on X-ray spectral characteristics



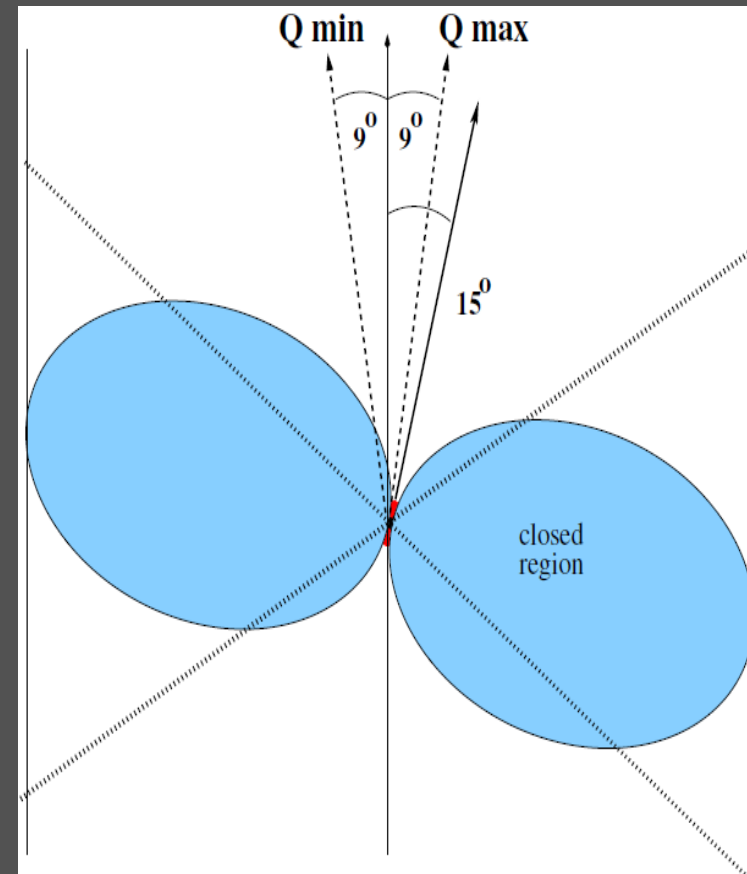
- Radio B mode: unpulsed, non-thermal emission (index and normalization within  $1 \sigma$  the same as for power-law component in Q mode)
- Radio Q mode: same unpulsed non-thermal emission plus pulsed thermal emission



Many unanswered questions, dilemma's:

- The polar cap region is viewed continuously: how to produce a 100%-pulsed thermal component in the Q mode?
- How to switch off a thermal X-ray component in the B mode while the radio emission becomes bright and ordered.

etc., etc, .....



Improved geometrical model

## New X-ray and radio campaign on PSR B1822-09

Characteristics: **PSR B0943+10**

- $P = 1.10 \text{ s}$
- $\dot{P} = 3.5 \times 10^{-15}$
- $\dot{E} = 1.0 \times 10^{32} \text{ erg s}^{-1}$
- $B_p = 2.0 \times 10^{12} \text{ G}$
- $T = 5.0 \times 10^6 \text{ yr}$

**PSR B1822-09**

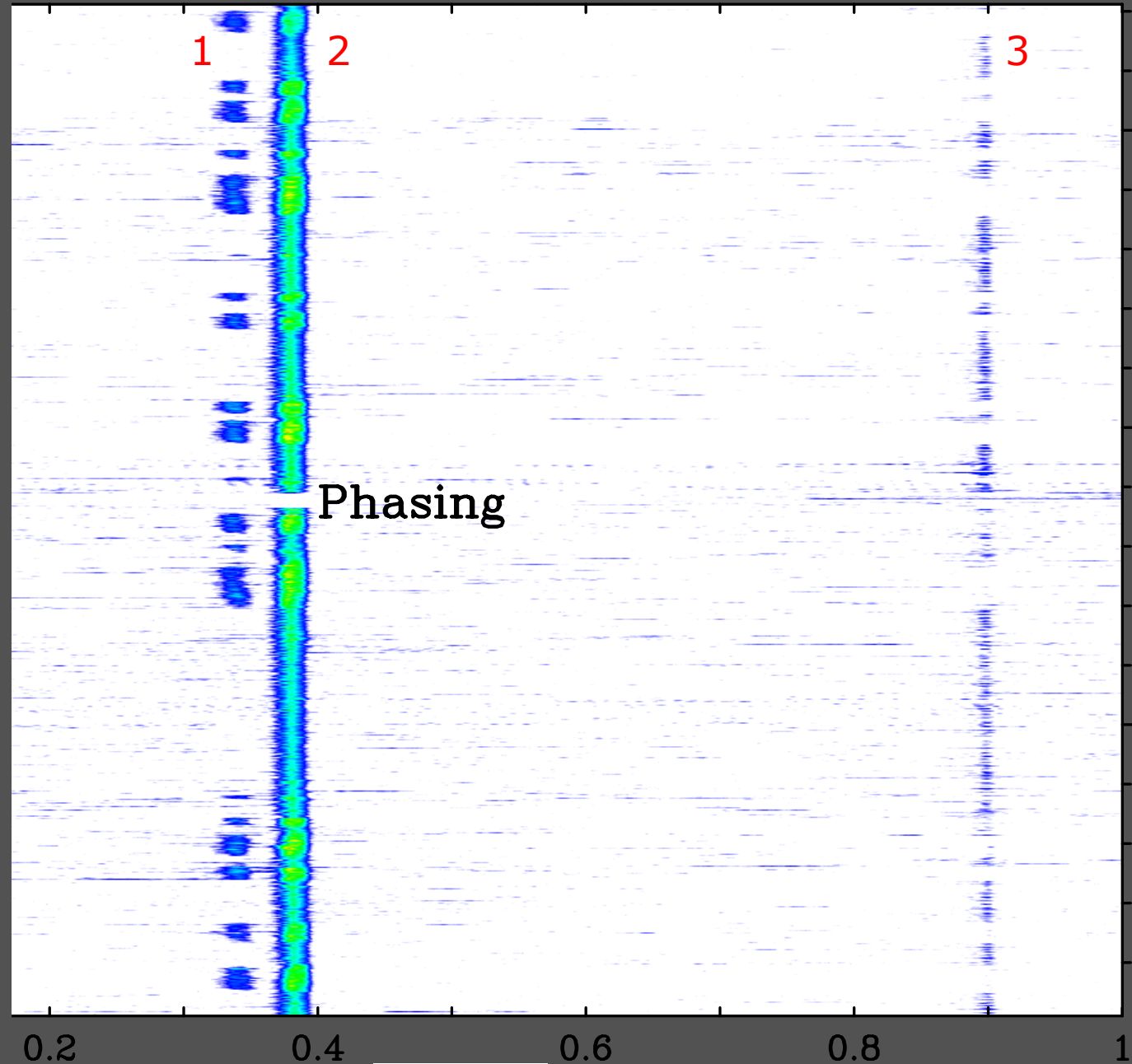
- $0.77 \text{ s}$
- $5.2 \times 10^{-14}$
- $4.5 \times 10^{33} \text{ erg s}^{-1}$
- $6.4 \times 10^{12} \text{ G}$
- $2.3 \times 10^5 \text{ yr}$
- nearly **aligned** rotator
- nearly **orthogonal** rotator  
(but, Malov & Nikitina 2011, 2013: aligned rotator)
- **Both pulsars mode switching between radio B(right) and Q(quiet) mode**

PSR B1822-09  
@ 624 MHz  
(GMRT)

Mode switching

- 1: Precursor
- 2: Main pulse
- 3: Interpulse

Typical mode  
durations less  
than 5 minutes



## XMM-Newton observation times (ks) in September, October 2013, and March 2014

Date /CCDs	10/09 2013	18/09 2013	22/09 2013	28/9 2013	30/09 2013	06/10 2013	10/03 2014	12/03 2014	Mode
PN	23.1	21.1	24.8	21.1	27.9	21.1	21.1	34.1	Large Window
MOS-1	24.8	22.8	26.5	22.8	29.6	22.8	22.8	35.8	Small Window
MOS-2	24.8	22.8	26.5	22.8	29.6	22.8	22.8	35.8	Small Window

Simultaneous radio observations with the **WSRT, Lovell and GMRT**

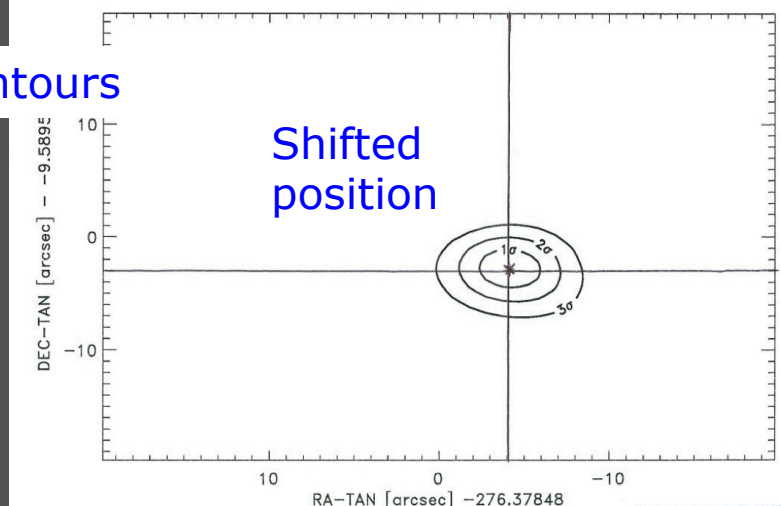
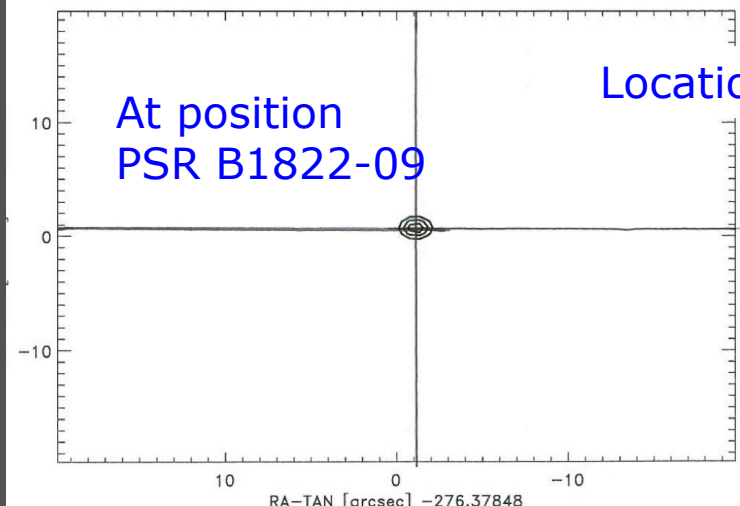
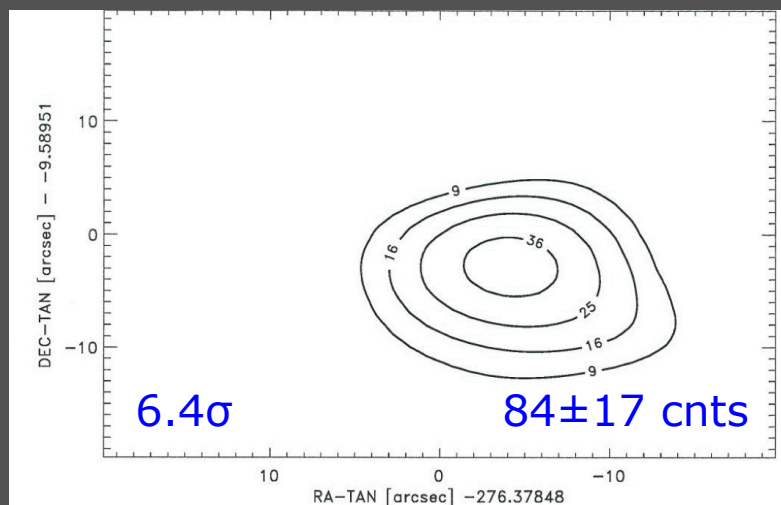
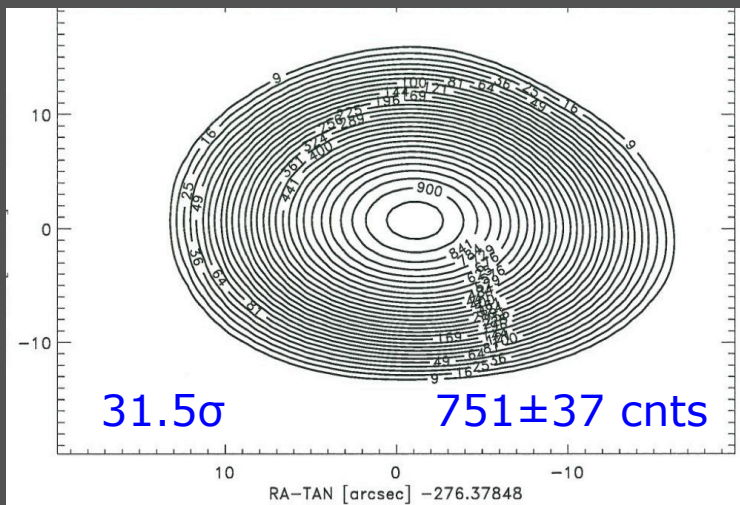
Total XMM-Newton

- PN 194.3 ks
- MOS-1 209.3 ks
- MOS-2 209.3 ks

# X-ray spatial analysis of skymaps

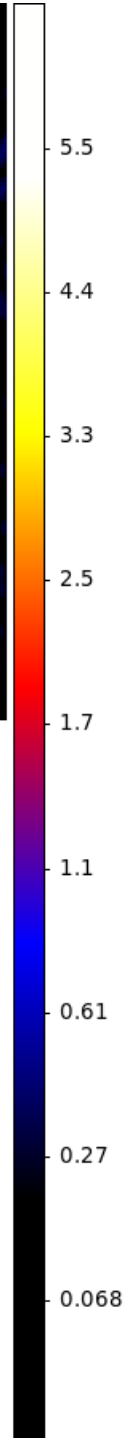
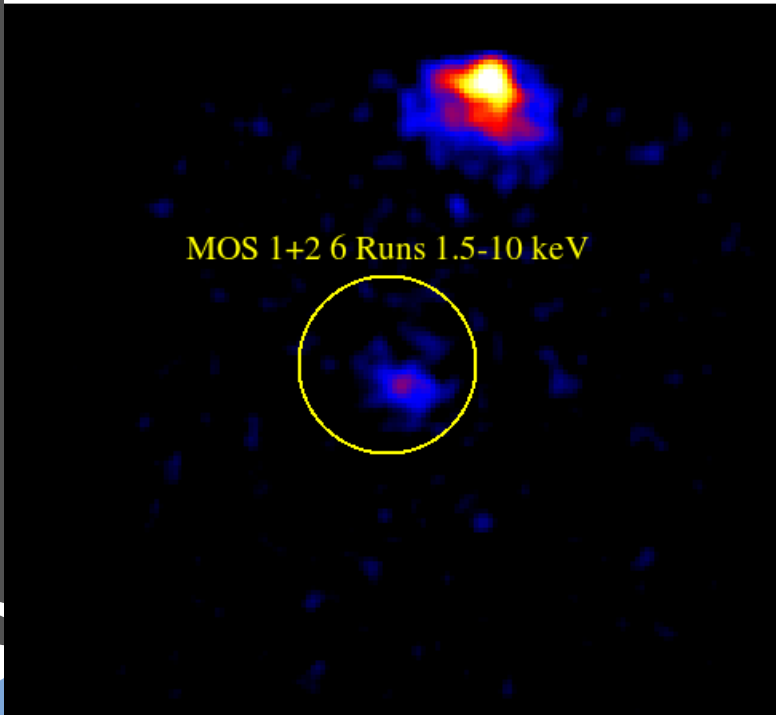
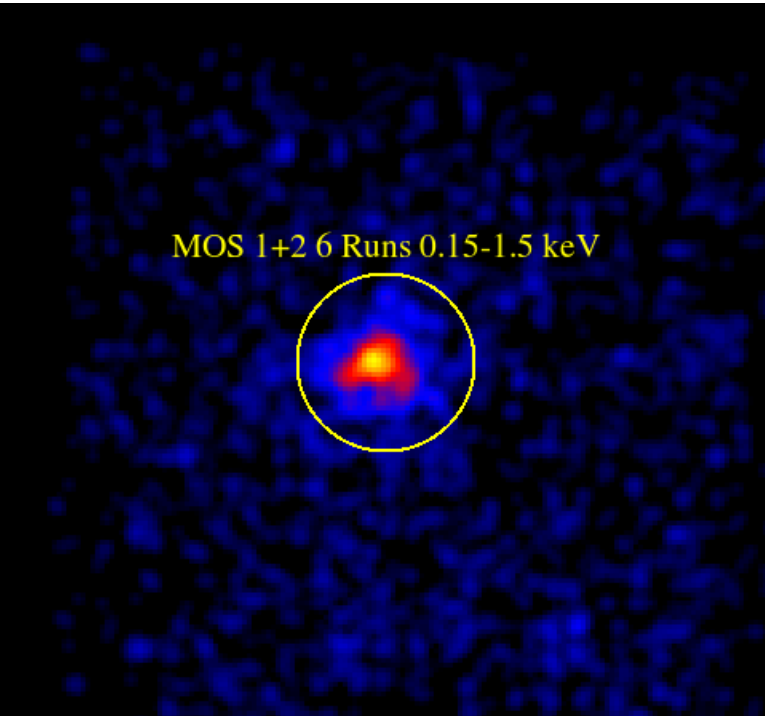
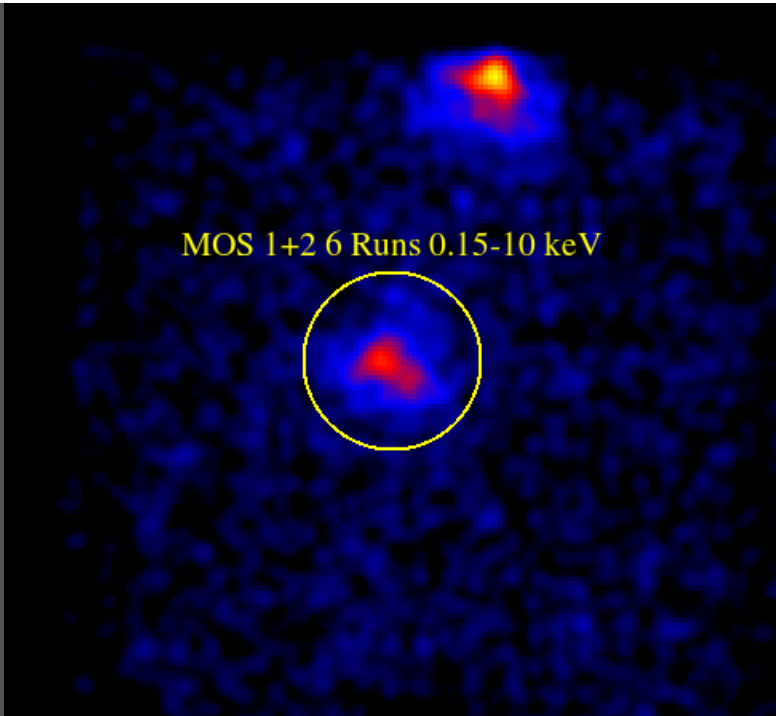
- **Maximum Likelihood Analysis**
- Two sources are detected separated by  $5.3'' \pm 0.5''$
- A soft-spectrum source at the position of PSR J1822-09, dominating below 1.4 keV
- A hard-spectrum source dominating above 1.4 keV

# Example Maximum Likelihood maps **PN detector, Runs 1-3**



**0.3 – 2 keV**

**2-10 keV**



Raw skymaps for MOS1+2,  
Runs 1-6

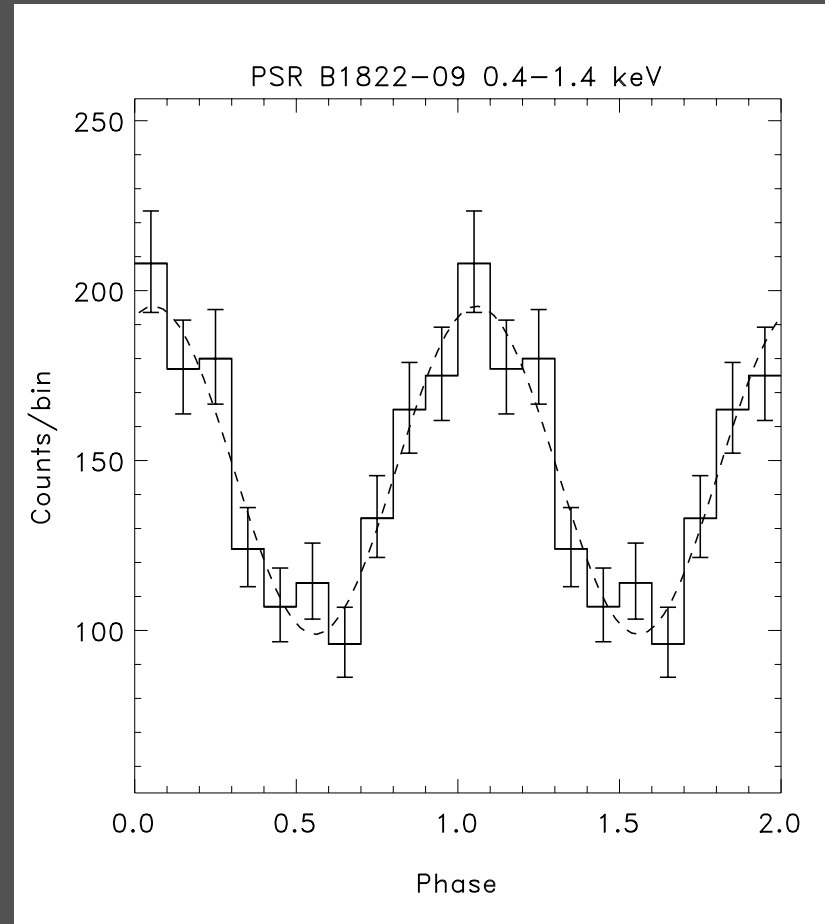
Total number of counts  
PSR1822-09  
Runs 1-8  
PN 2144 cnts  
MOS1+2 810 cnts

# X-ray timing analysis (Runs 1-8, PN+MOS1+2)

Detection of X-ray pulsation in energy band 0.4-1.4 keV

Phase folding with ephemeris from Jodrell Bank

- Broad sinusoidal X-ray pulse shifted by  $\sim 0.07$  in phase with respect to radio main pulse (0.0)
- No indication for X-ray pulse from radio interpulse
- Pulsed fraction  $\sim 35\%$



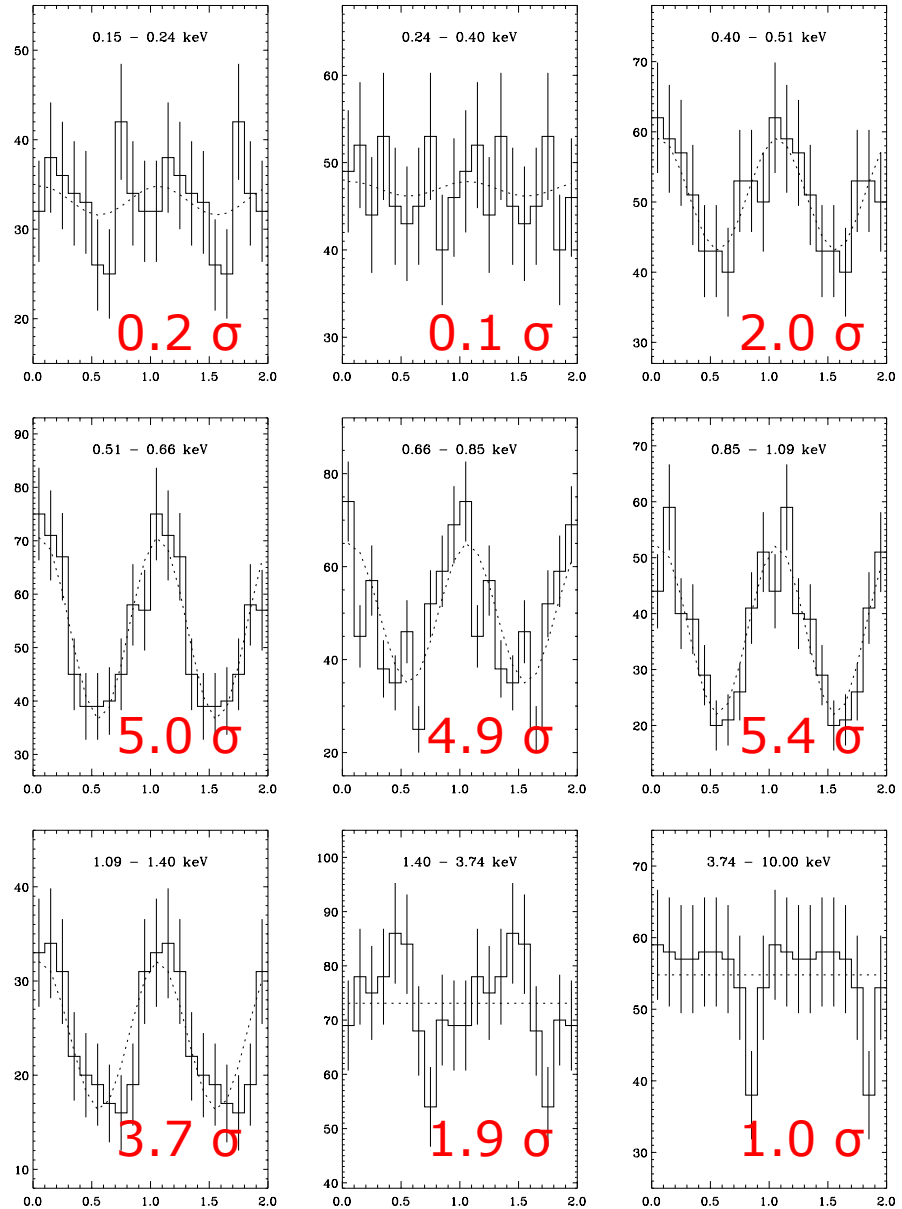
**9.8  $\sigma$**  detection significance



# X-ray timing analysis

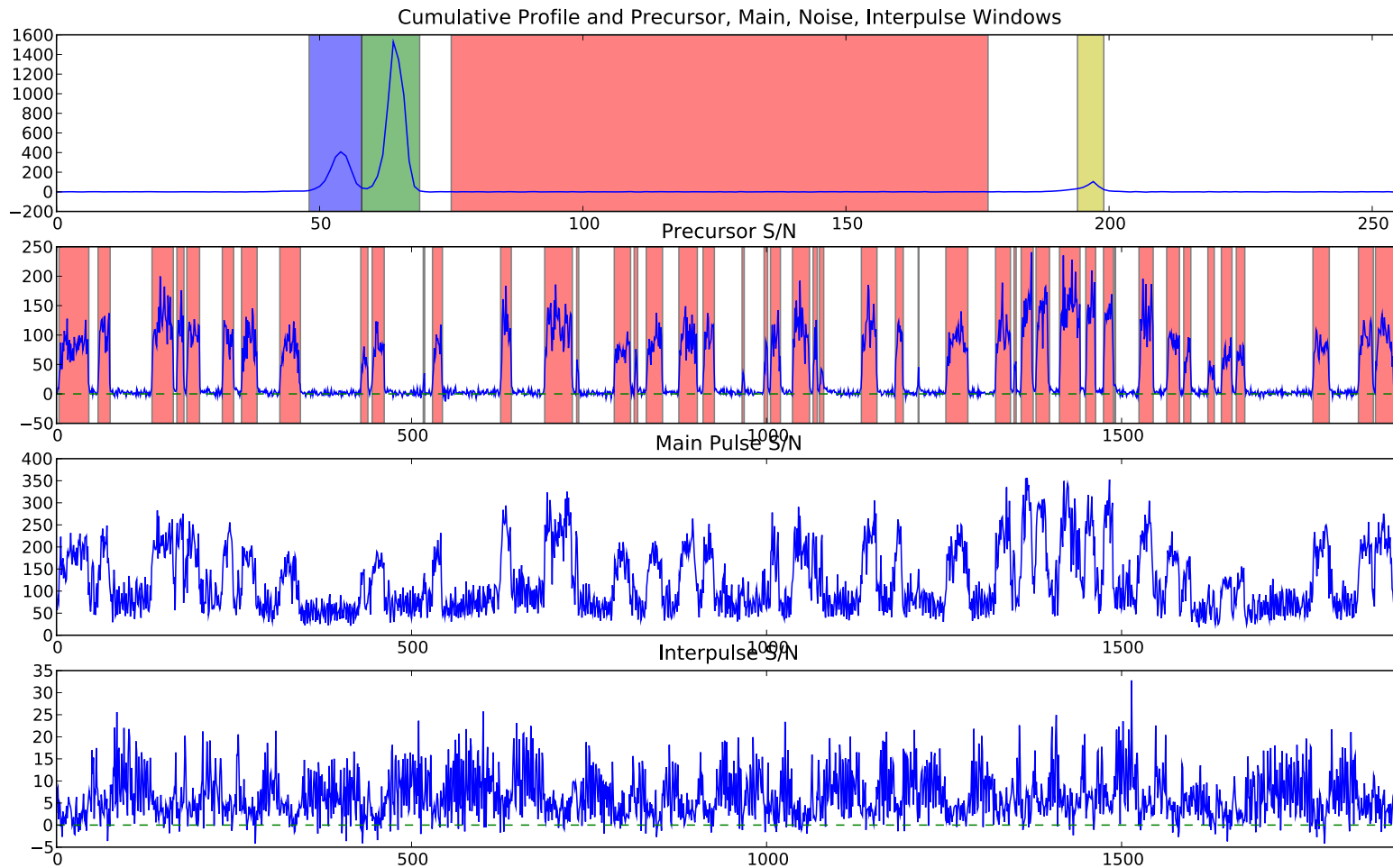
Profiles in differential energy bands: pulse detections only between 0.4 and 1.4 keV ( $z_1^2$  statistic)

Dashed profiles are fits with profile shape for integral energy band 0.4-1.4 keV at fixed phase.



# X-ray mode switching ?

## PSR B1822-09, 5.55 hrs observing with the WSRT



S/N of detection in bins of 10 s

# X-ray mode switching? Preliminary result for runs 1-6

Q-mode 85.5 ks and B-mode 50.4 ks

- Skymap analysis with Maximum Likelihood Method of count maps separately for PN and MOS1+2 detectors and  $\Delta E$ :
  - ◆ 0.15-0.4 keV & 0.4-1.4 keV fitting 1 source (PSR B1822-09)
  - ◆ 1.4-10 keV fitting 2 sources

# X-ray mode switching? Preliminary result for runs 1-6

Q-mode 85.5 ks and B-mode 50.4 ks

Count rates total emission PSR B1822-09

$\Delta E$	0.15 – 0.4 keV	0.4 – 1.4 keV	1.4 – 10 keV
No mode selection	$(2.52 \pm 0.21) \times 10^{-3}$	$(8.97 \pm 0.34) \times 10^{-3}$	$(0.73 \pm 0.27) \times 10^{-3}$
B-mode	$(2.67 \pm 0.35) \times 10^{-3}$	$(7.98 \pm 0.53) \times 10^{-3}$	$(0.96 \pm 0.44) \times 10^{-3}$
Q-mode	$(2.41 \pm 0.27) \times 10^{-3}$	$(9.60 \pm 0.44) \times 10^{-3}$	$(0.57 \pm 0.34) \times 10^{-3}$
$\Delta$ B-Q modes	0.6 $\sigma$	-2.4 $\sigma$	0.7 $\sigma$

→ only 2.4  $\sigma$  hint for X-ray mode switching in 0.4-1.4 keV band, which contains pulsed signal

# X-ray mode switching? Preliminary result for runs 1-6

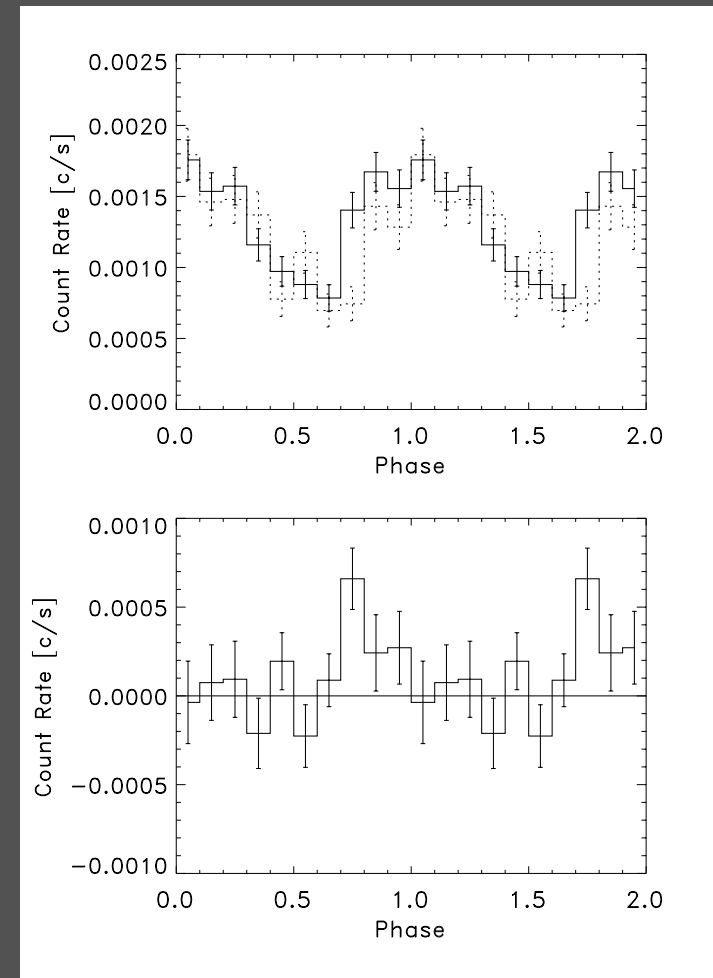
Q-mode 85.5 ks and B-mode 50.4 ks

- Timing/phase resolved imaging for 10 phase bins

Upper figure shows source count rates: broken lines B mode, solid line Q mode

Lower figure: count rate difference (Q-B) as a function of phase

KS-test: 30% chance probability that both distributions are drawn from same parent distribution



# X-ray mode switching? Preliminary result for runs 1-6

Q-mode 85.5 ks and B-mode 50.4 ks

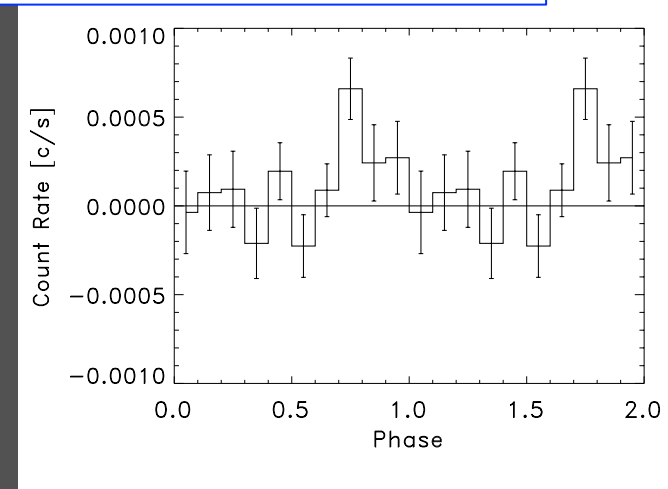
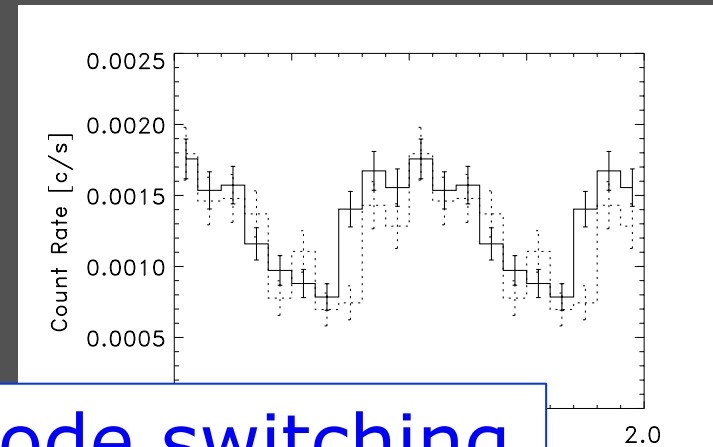
- Timing/phase resolved imaging for 10 phase bins

Upper figure shows source count rates: broken lines B mode, solid line Q mode

**No significant evidence for mode switching**

Lower figure: count rate difference (Q - B) as a function of phase

**KS-test:** 30% chance probability that both distributions are drawn from same parent distribution

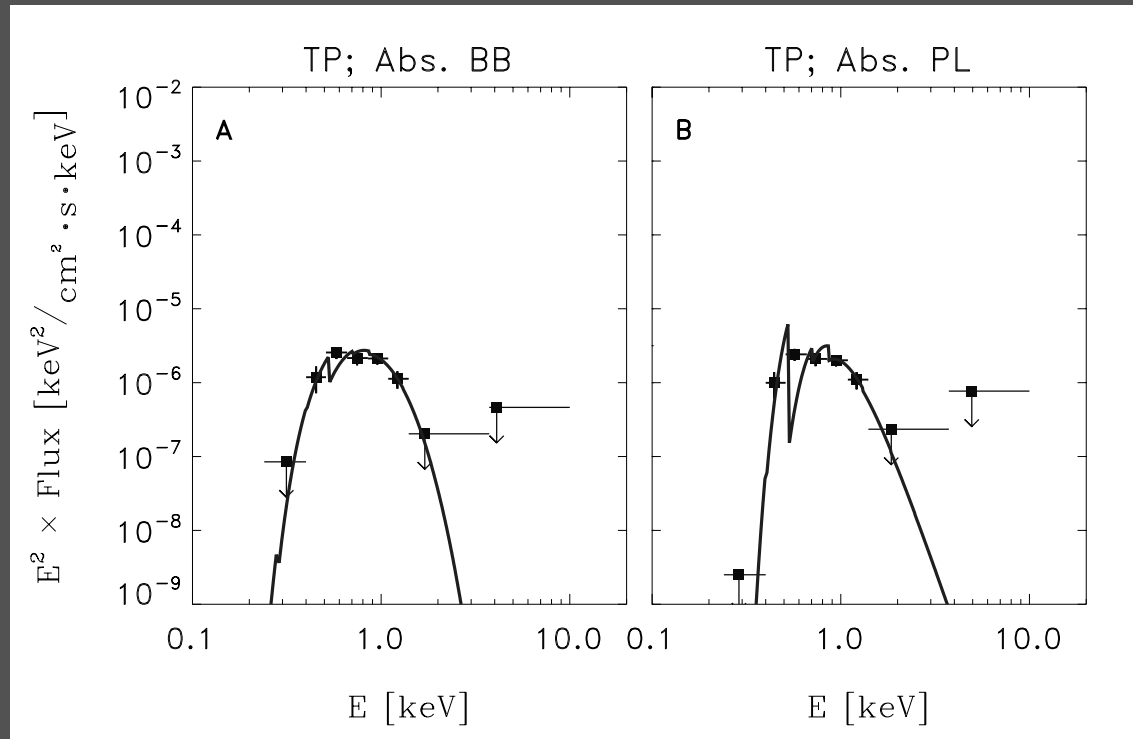


# Spectral analysis

- Distance PSR B1822-09:  
Upper limit 1.9 kpc (Johnston et al. 2001),  
often quoted  $d \sim 1$  kpc (e.g. Zhou et al. 2005).
- $N_{\text{H}}$  at  $\sim 1.9$  kpc is  $\sim 3 \cdot 10^{21} \text{ cm}^{-2}$
- $N_{\text{H}}$  is in initial analysis treated as free parameter



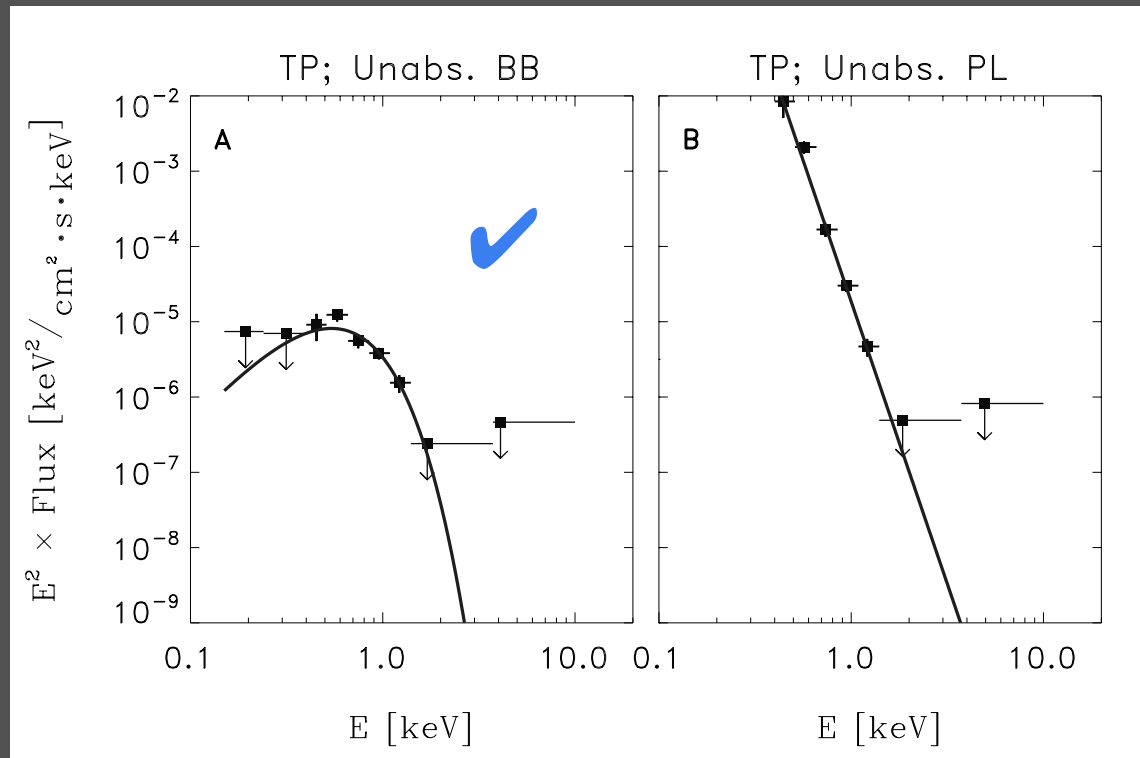
# Pulsed emission X-ray spectrum of PSR B1822-09: 'absorbed'



- **BB fit:  $\chi^2_{\nu} = 1.14$  (d.o.f. 6)**
- $N_H = (2.03^{+0.35}_{-0.30}) 10^{21} \text{ cm}^{-2}$
- $BB_{\text{norm}} = 0.0047 \pm 0.0007$
- $kT = 0.138^{+0.003}_{-0.004} \text{ keV}$

- **PL fit:  $\chi^2_{\nu} = 1.11$  (d.o.f. 6)**
- $N_H = (9.28^{+0.36}_{-0.30}) 10^{21} \text{ cm}^{-2}$
- $\alpha = (1.85 \pm 0.28) 10^{-5}$   
ph  $\text{cm}^{-2} \text{ keV}^{-1}$  at 1 keV
- $\Gamma = -9.5^{+0.5}_{-0.4}$

# Pulsed emission X-ray spectrum of PSR B1822-09 ('unabsorbed')



- BB fit:  $\chi^2_{\nu} = 1.14$  (d.o.f. 6)

- $N_{\text{H}} = (2.03^{+0.35}_{-0.30}) 10^{21} \text{ cm}^{-2}$

- BB:  $kT = 0.138^{+0.003}_{-0.004} \text{ keV}$

- $R_{\text{hot spot}} \approx 213 \text{ m}$  ( $d = 1 \text{ kpc}$ )

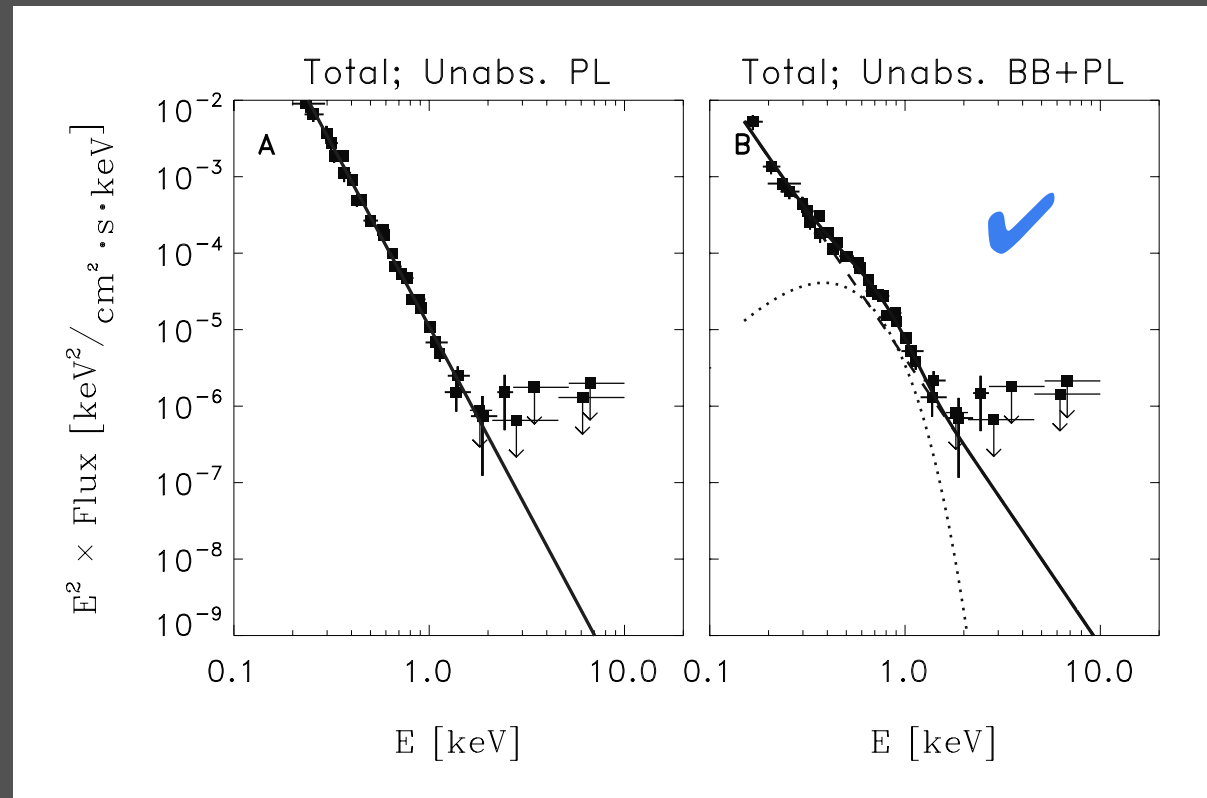
- PL fit:  $\chi^2_{\nu} = 1.11$  (d.o.f. 6)

- $N_{\text{H}} = (9.28^{+0.36}_{-0.30}) 10^{21} \text{ cm}^{-2}$

- $\Gamma = -9.5^{+0.5}_{-0.4}$

→ Values for  $N_{\text{H}}$  and  $\Gamma$  in PL fit are not acceptable/realistic

# Total emission X-ray spectrum of PSR B1822-09 ('unabsorbed')



- PL fit:  $\chi^2_{\nu} = 1.65$  (d.o.f. 32)  
 **$\sim 1\%$  prob.**

- $N_{\text{H}} = (4.15 \pm 0.05) 10^{21} \text{ cm}^{-2}$

- $\Gamma = -6.77 \pm 0.07$

- BB fit :  $\chi^2_{\nu} = 1.60$  (d.o.f. 32)  
 **$\sim 1\text{-}2\%$  prob.**

- BB+PL fit:  $\chi^2_{\nu} = 1.43$  (d.o.f. 30)  
 **$\sim 7.5\%$  prob.**

- $N_{\text{H}} = (2.85 \pm 0.08) 10^{21} \text{ cm}^{-2}$

- $\text{BB}_{\text{norm}} = 0.093 \pm 0.010$

- $kT = 0.098 \pm 0.002 \text{ keV}$

- $R_{\text{hot spot}} \approx 948 \text{ m}$  ( $d = 1 \text{ kpc}$ )

- $a = (1.85 \pm 0.28) 10^{-5} \text{ ph cm}^{-2} \text{ keV}^{-1}$  at 1 keV

- $\Gamma = -5.76 \pm 0.15$

# JOINT FIT TO **TOTAL AND PULSED** SPECTRA OF PSR B1822-09

Assuming that:

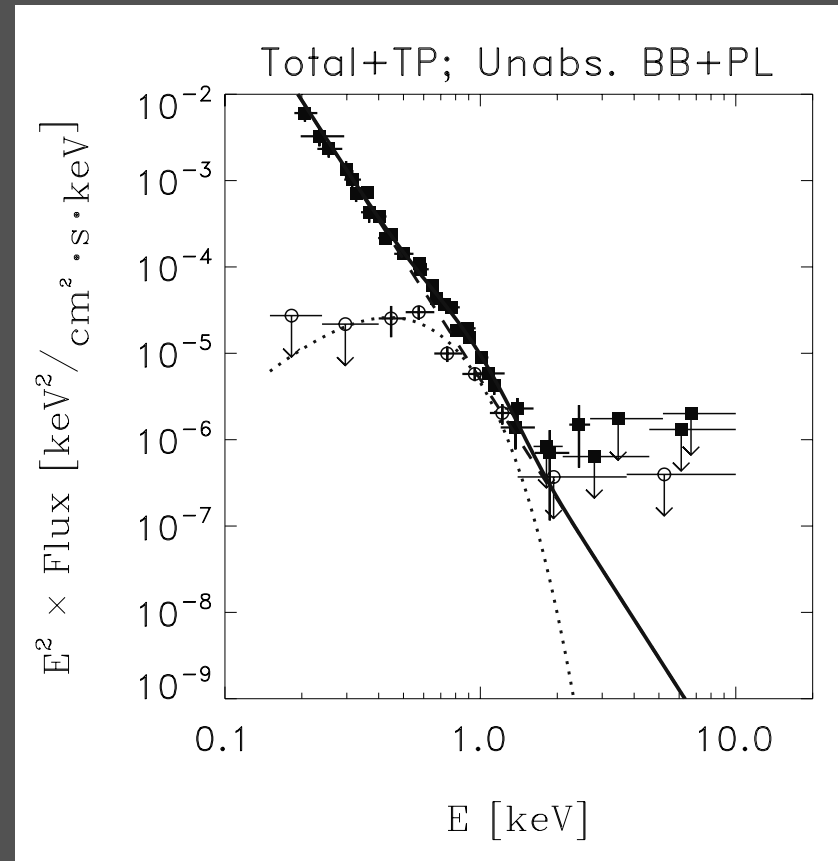
- pulsed emission is a BB and equals BB component in total spectrum.

- Unpulsed emission has a PL shape

- **BB+PL fit:  $\chi^2_{\nu} = 1.40$  (d.o.f. 39)  
~5% prob.**

- $N_{\text{H}} = (3.39 \pm 0.08) 10^{21} \text{ cm}^{-2}$
- $\text{BB}_{\text{norm}} = 0.0345 \pm 0.0046$
- $kT = 0.112 \pm 0.002 \text{ keV}$
- $R_{\text{hot spot}} \approx 578 \text{ m}$  ( $d = 1 \text{ kpc}$ )

- $a = (5.0 \pm 0.4) 10^{-6} \text{ ph cm}^{-2} \text{ keV}^{-1}$  at 1 keV
- $\Gamma = -6.62 \pm 0.13$



# Conclusions / Dilemma's

- **PSR B1822-09** has been detected with XMM-Newton with a pulsed fraction of  $\sim 35\%$  in the energy band 0.4-1.4 keV
- The pulse profile is sinusoidal; maximum within 0.1 phase from the peak of the radio main pulse; no pulse detected at the phase of the weak radio inter pulse.
- There is no significant evidence for simultaneous X-ray-radio mode-switching by PSR B1822-09. What causes this difference with PSR B0943+10?
- The spectrum of the total emission is best fitted with a thermal BB-component plus a non-thermal PL component.
- The spectrum of the pulsed component is best fitted with a thermal component, consistent with a heated polar cap passing through our line-of-sight for an orthogonal rotator. But, what about a non-thermal unpulsed component?
- The X-ray-radio characteristics of PSR B1822-09 do not shed more light on the interpretation of the X-ray/radio characteristics of PSR B0943+10



Thank you for listening!